

# Heavy Neutral Lepton Searches at ICARUS-NuMI

JOSU HERNANDEZ-GARCIA

BASED ON ARXiv: 2408.03383. IN COLLABORATION WITH A. DE ROECK AND A. CHATTERJEE

LIGHT DARK WORLD 2025 | SEPT 2025

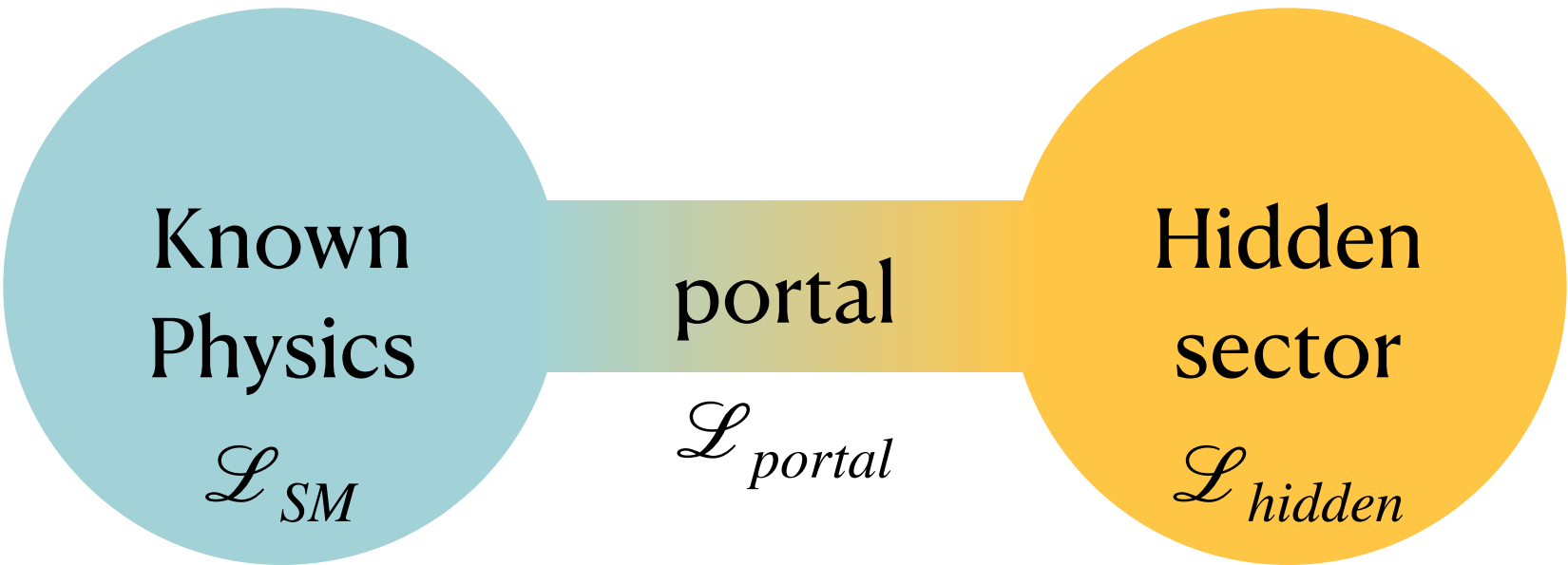
# LLPs: Portals to BSM Physics

Long-Lived Particles (LLPs) could be the portal to hidden sectors addressing different Standard Model open problems



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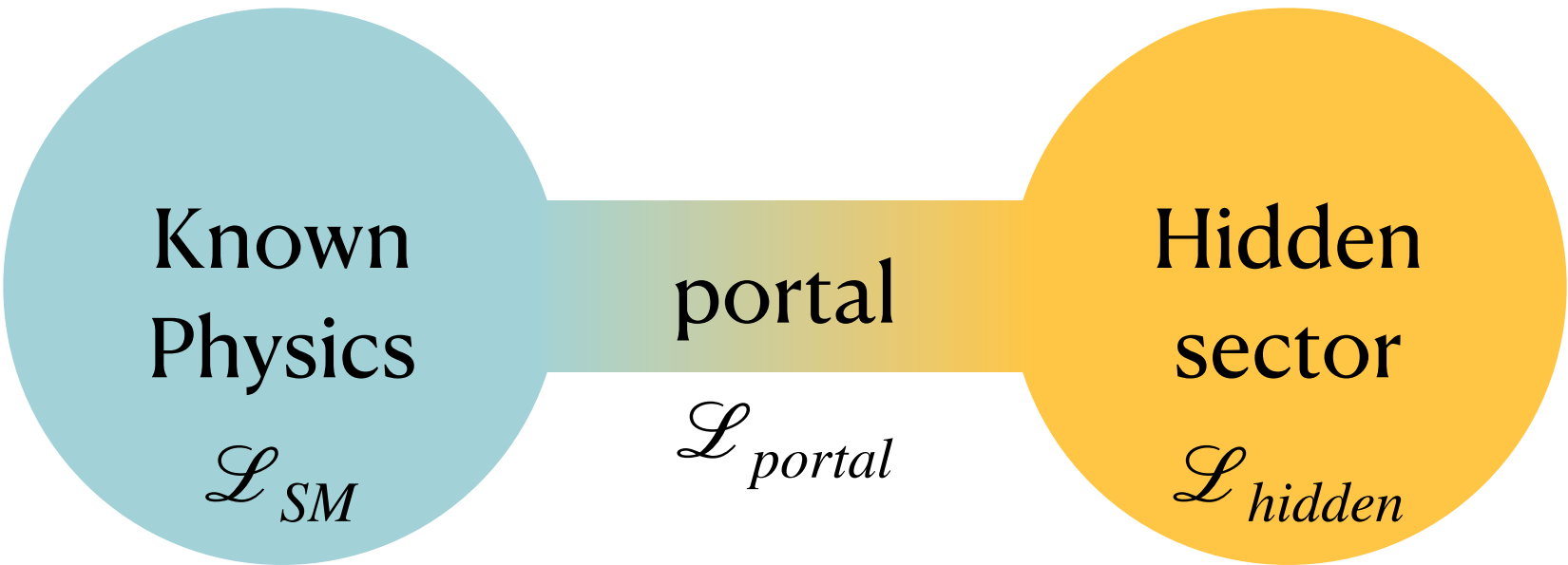
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BSM portal	New particle	New parameters	SM open problem
Fermion portal	HNLs – $N$	$M_N, U_{\alpha N}$	Neutrino masses
Axion portal	ALPs – $a$	$m_a, f_a$	Strong CP
Scalar portal	Dark scalar – $S$	$m_S, \theta$	Vacuum stability
Vector portal	Dark photon – $Z'$	$m_{Z'}, \epsilon$	DM candidate

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# LLPs signals at neutrino experiments

Neutrinos are feebly-interacting particles

- intense sources needed
- large detectors needed

Interactions  $\rightarrow$  scale with detector **mass**

LLPs decays  $\rightarrow$  scale with detector **volume**

Generally lower backgrounds if

- Detector has low density/mass
- decays can be fully reconstructed  $\Rightarrow$   
instrumented detector (ECAL, magnetic field, ...)

# Phenomenology of Heavy Neutral Leptons

Origin of neutrino **masses** and **mixings**

HNL **mixes** with active neutrinos

$$\nu_\alpha = \sum_i U_{\alpha i} \nu_i + U_{\alpha N} N$$

HNL **CC** and **NC** interactions:

$$\mathcal{L}_{N,int} \supset -\frac{m_W}{v} \bar{N} U_{\alpha N}^* \gamma^\mu \ell_{L\alpha} W_\mu^+ - \frac{m_Z}{\sqrt{2}v} \bar{N} U_{\alpha N}^* \gamma^\mu \nu_{L\alpha} Z_\mu$$

Minimal model: 2 HNLs needed.  
Phenomenology given by

**4 free parameters:**

- HNL mass,  $M_N$
- mixing with three  $\nu_\alpha$ ,  $U_{\alpha N}$

Heavy Neutral Leptons (HNLs)

$U_{\alpha N} \ll 1 \Rightarrow$  **weaker** than weak interaction

Experimental results shown **single flavor dominance**  $\Rightarrow$  HNL coupled to **one flavor** at a time:

$$|U_{eN}|^2 : |U_{\mu N}|^2 : |U_{\tau N}|^2 = 1 : 0 : 0$$

**e dominance**

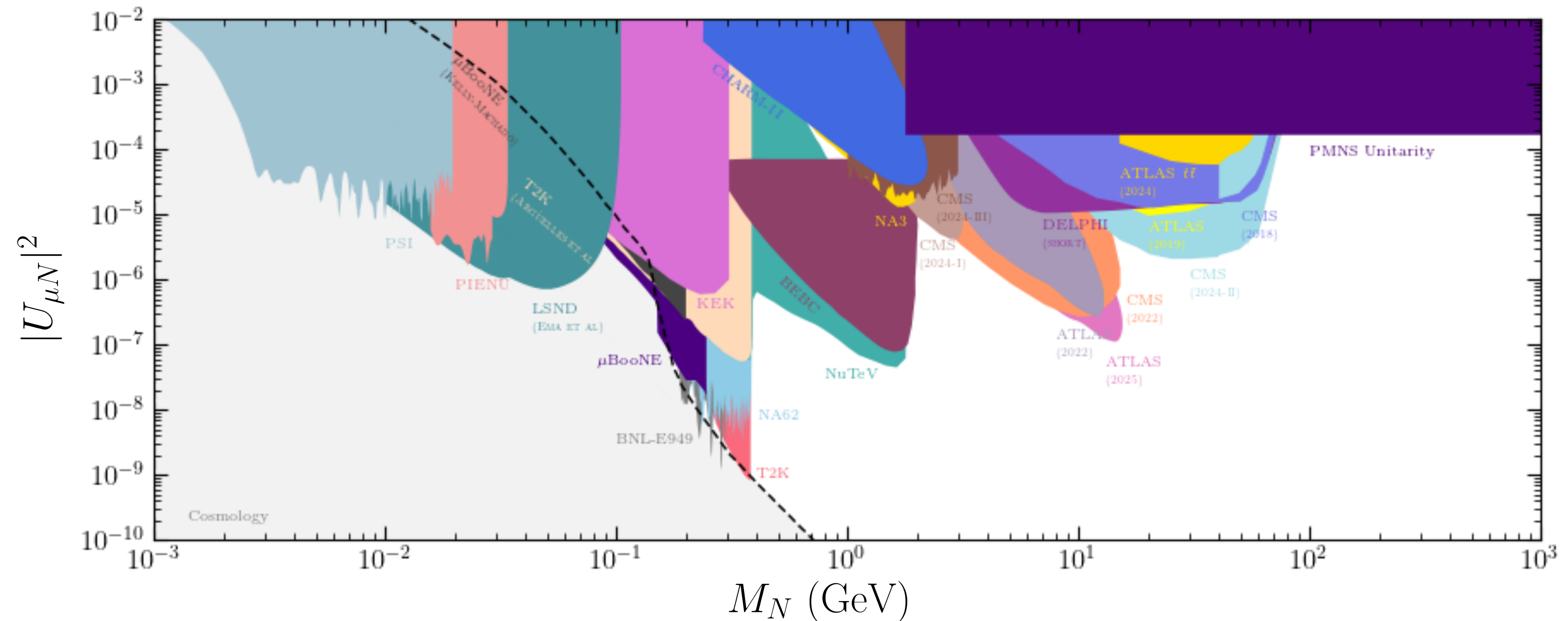
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**$\mu$  dominance**

$$|U_{eN}|^2 : |U_{\mu N}|^2 : |U_{\tau N}|^2 = 0 : 0 : 1$$

**$\tau$  dominance**

# Heavy Neutral Leptons: present bounds

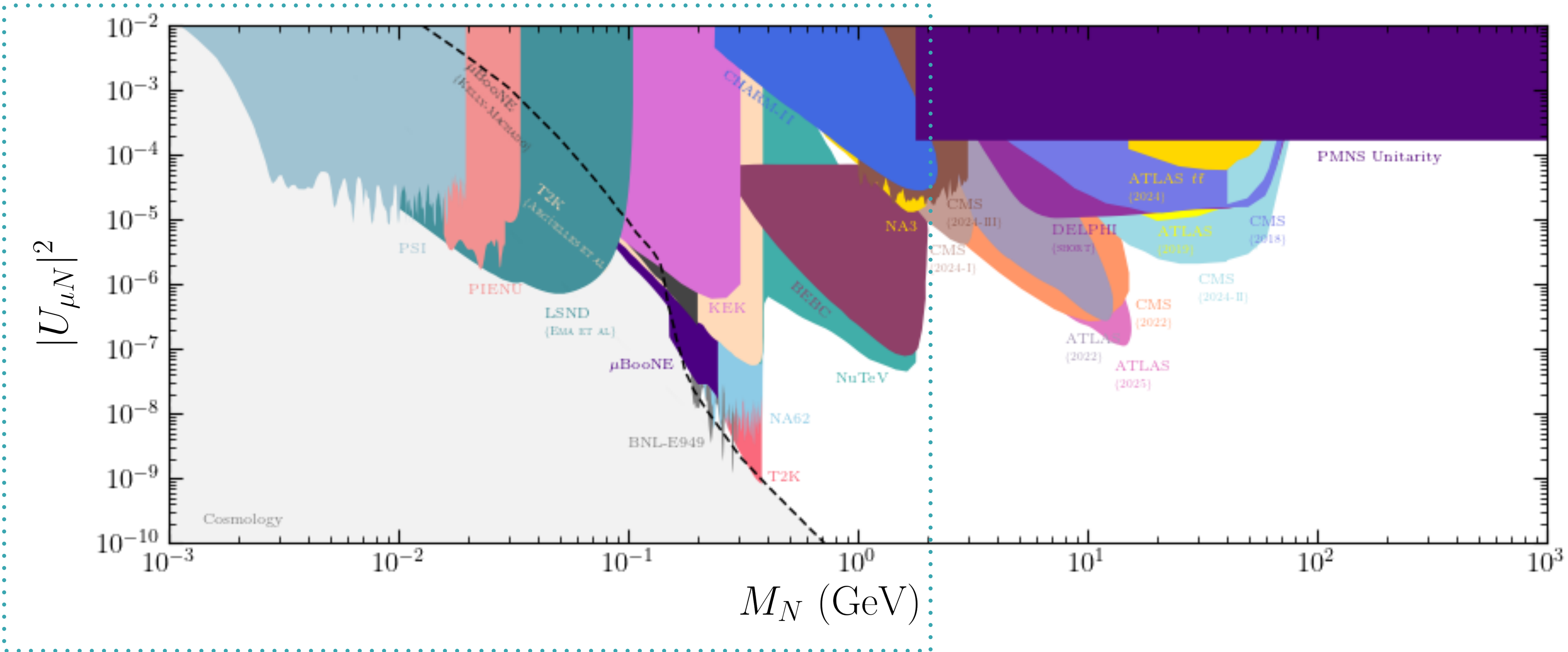


Present bounds on the HNL mixing from  $\nu$  oscillations, kinks in beta-decay, peaks in meson decays, beam-dumps, colliders, and non-unitarity of the PMNS.

<https://github.com/mhostert/Heavy-Neutrino-Limits>  
 E. Fernandez-Martinez, M. Gonzalez-Lopez, JHG, M. Hostert, J. Lopez-Pavon. JHEP 09 2023 001



# Heavy Neutral Leptons: present bounds



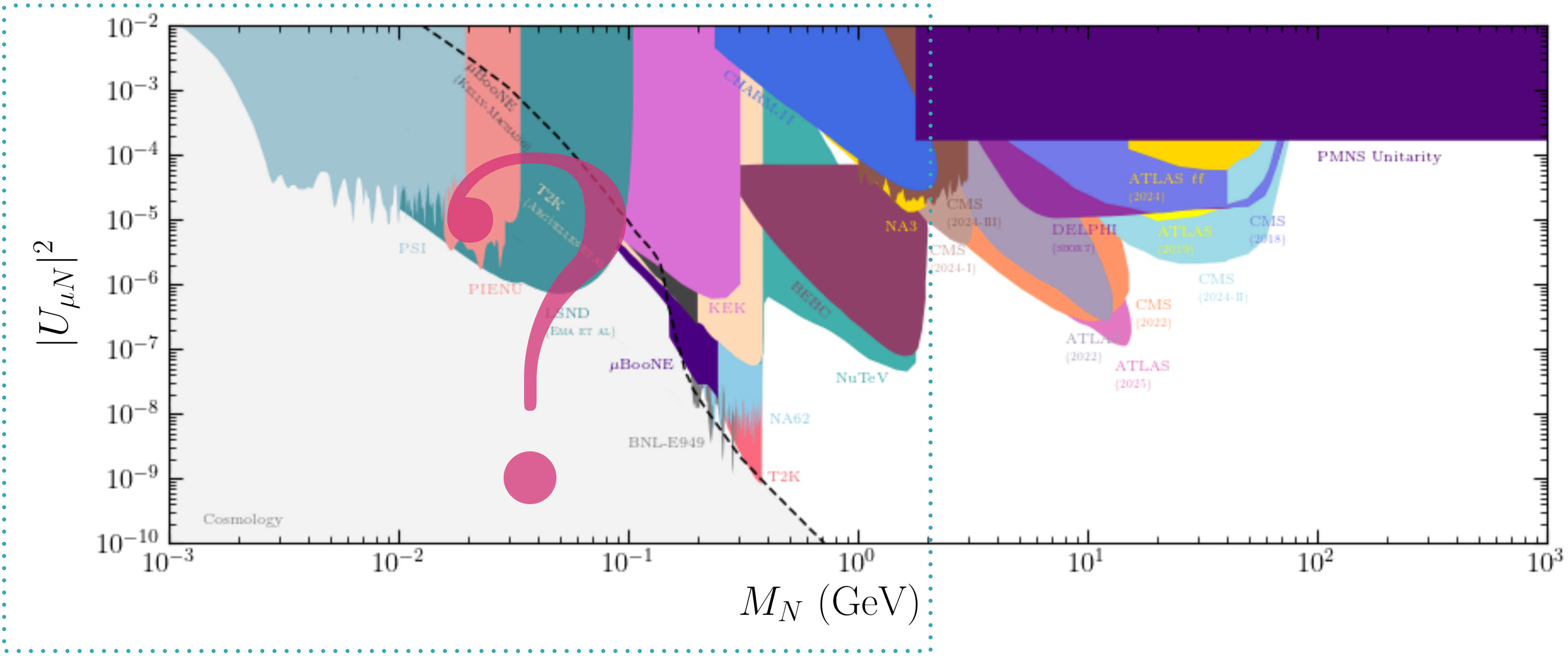
Present bounds in the HNL mass window up to  $\sim 2$  GeV led by peak searches and beam-dump experiments

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# Heavy Neutral Leptons: present bounds



Bounds from on-going and near future experiments.

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# The ICARUS experiment at SBN

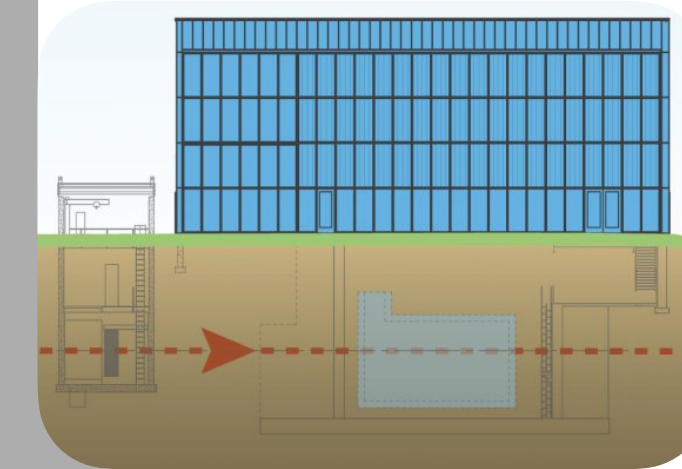
Proposed to:

- Clarify LSND and MiniBooNE anomalies
- Explore sterile neutrino in the eV mass scale

Three liquid Ar TPCs

Taking data since 2015

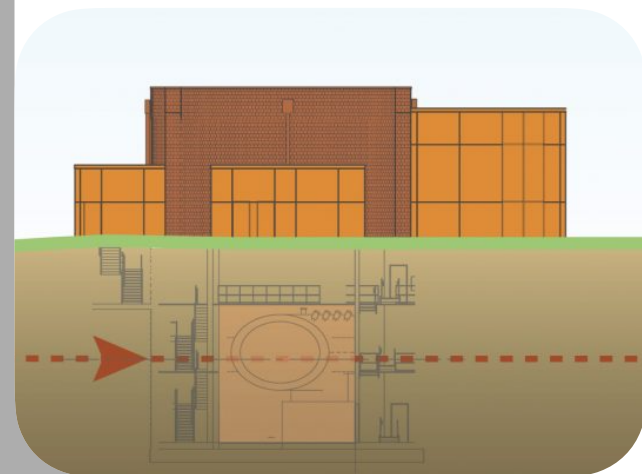
SBND



270 tons  
Baseline: 110 m

2024

$\mu$ BooNE



170 tons  
Baseline: 470 m

2015 – 2021

Short-Baseline Neutrino (SBN)

ICARUS

760 tons  
Baseline: 600 m

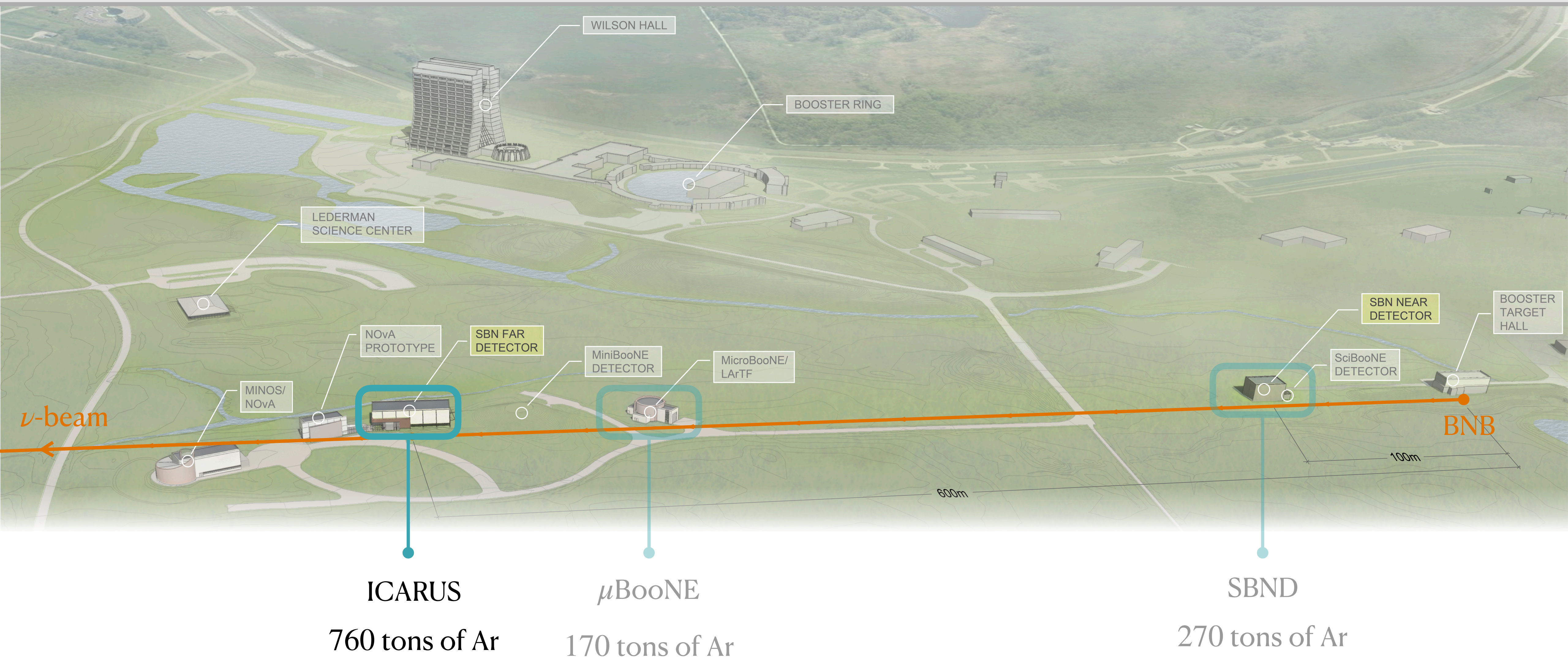


2021

Exposed to the Booster Neutrino Beam (BNB) produced from 8 GeV protons from the booster

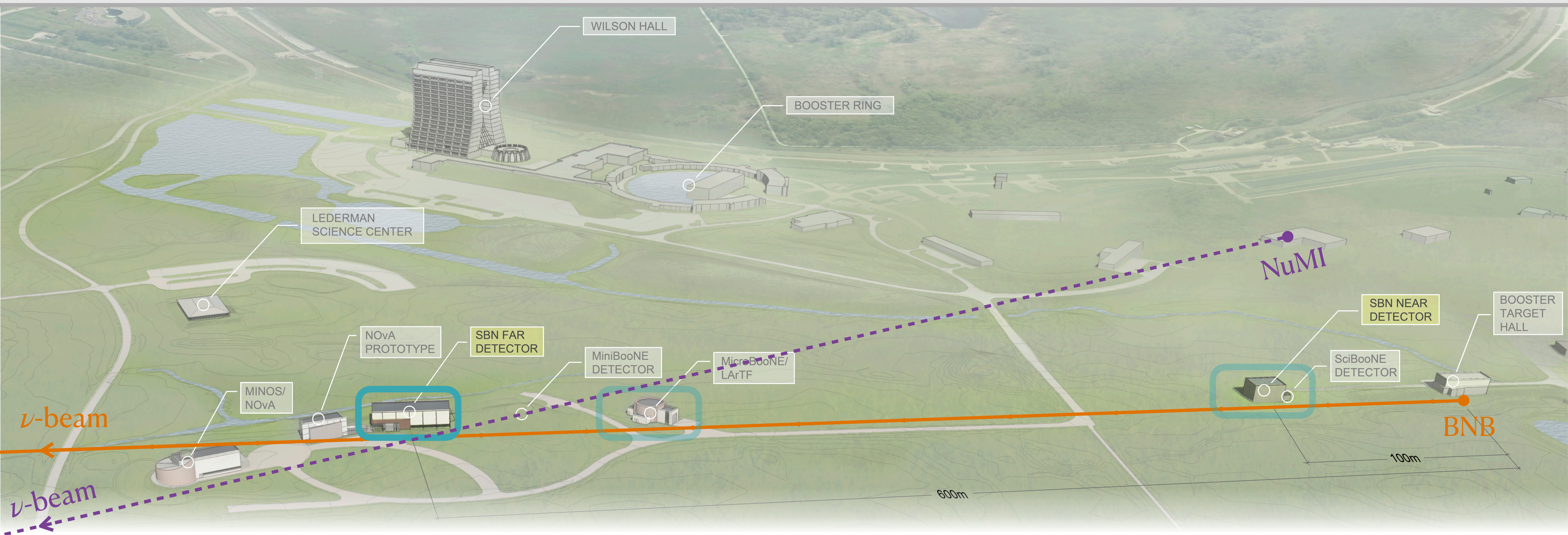


# Experimental setup: SBN





# Experimental setup: BNB and NuMI



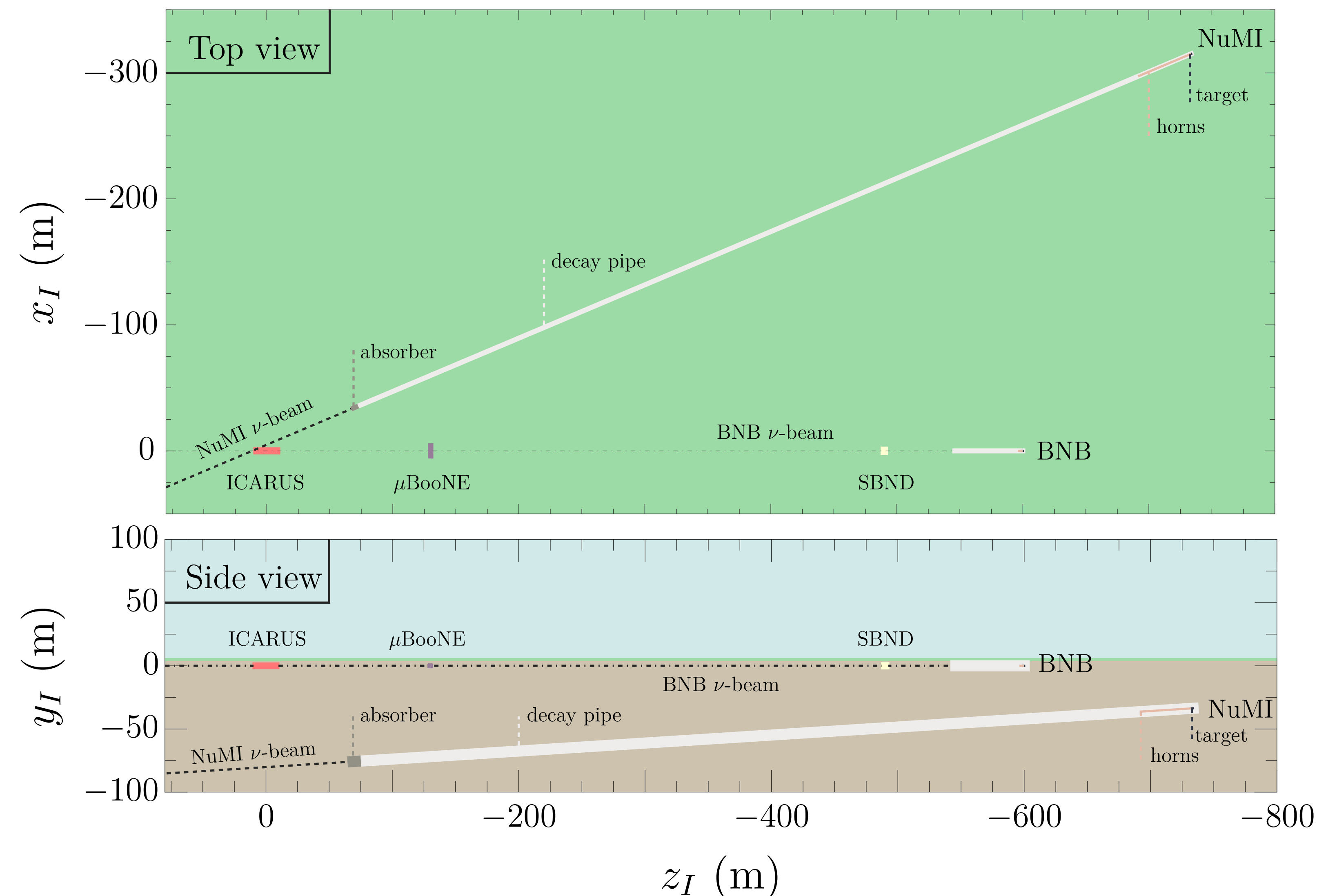
ICARUS is also exposed to the NuMI  $\nu$ -beam.

NuMI: 120 GeV protons from main injector against graphite target, accumulating  $6.6 \cdot 10^{20}$  PoT/year.

C. Rubbia (ICARUS-WA104, LAr1-ND, MicroBooNE), (2015)



# Experimental setup: BNB and NuMI

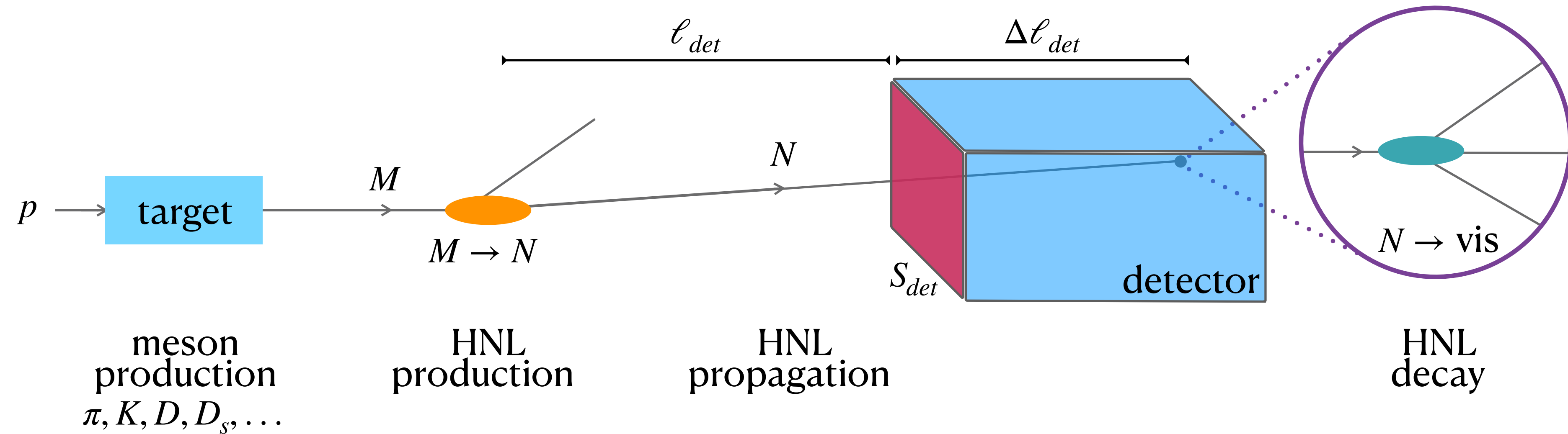


Off-axis NuMI flux  
crossing ICARUS  
400 mrad

NuMI downward  
inclination  
57 mrad

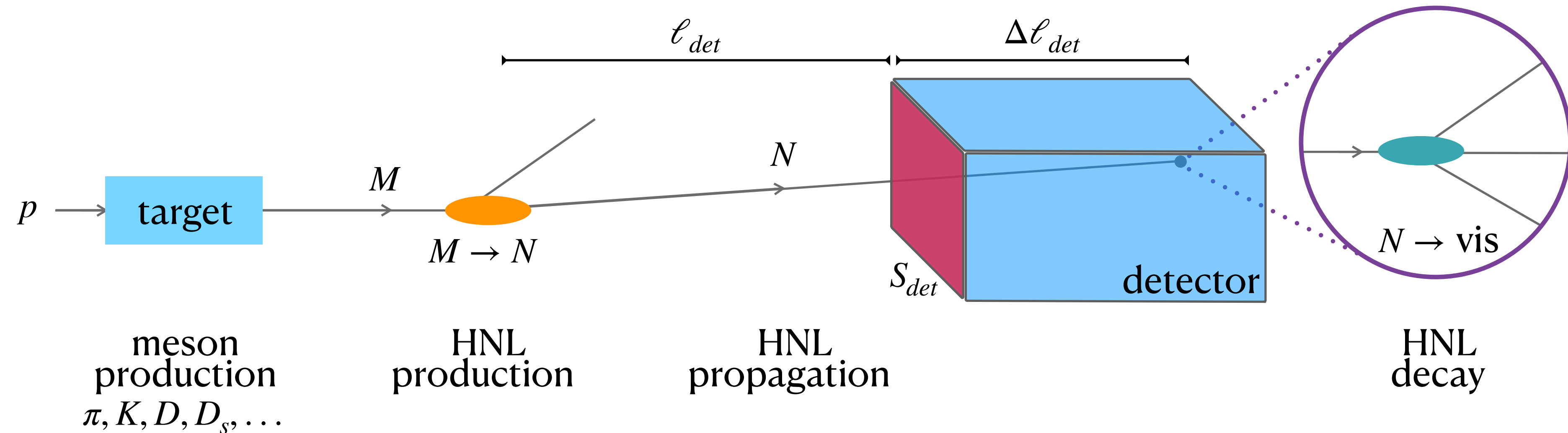
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HNLs could be produced via meson decay at the beam-dump experiment beam



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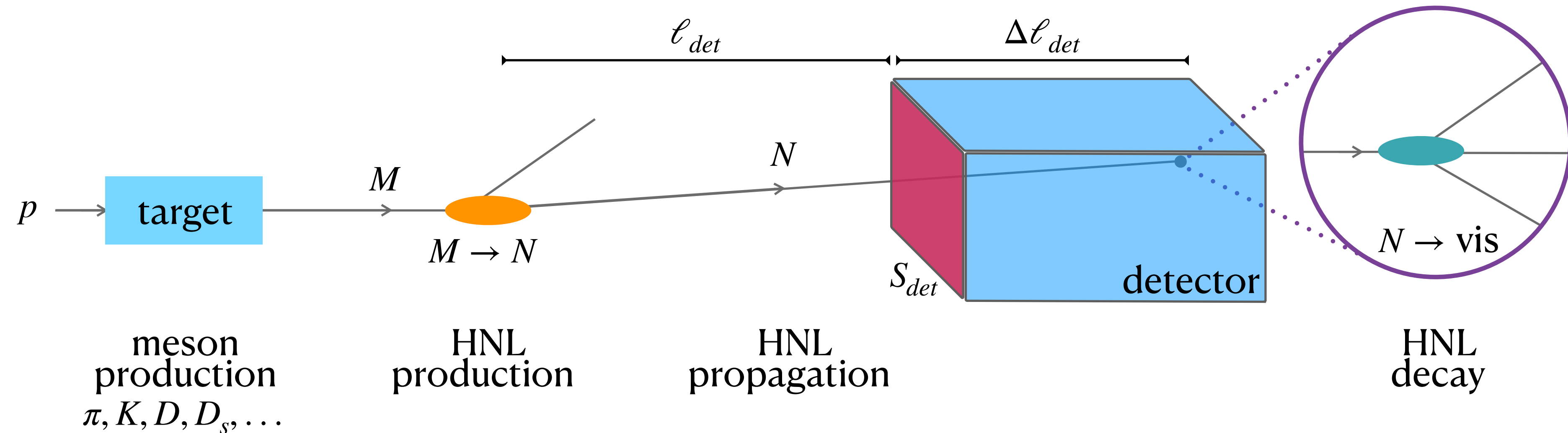


The number of decay events

$$N_{ev} = N_M \text{BR}(M \rightarrow N) \text{BR}(N \rightarrow \text{vis}) \epsilon_{det} \int dS \int dE_\Psi P(c\tau_N/M_N, E_N, \Omega_N) \frac{dn^{M \rightarrow N}}{dS dE_N}$$

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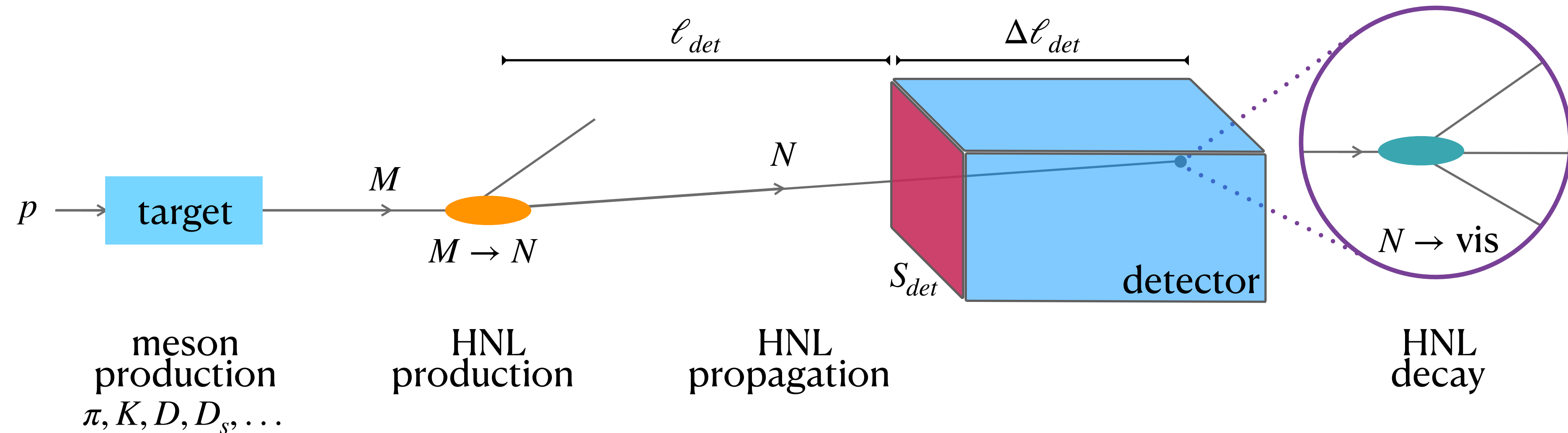
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experiment-dependent  
quantities

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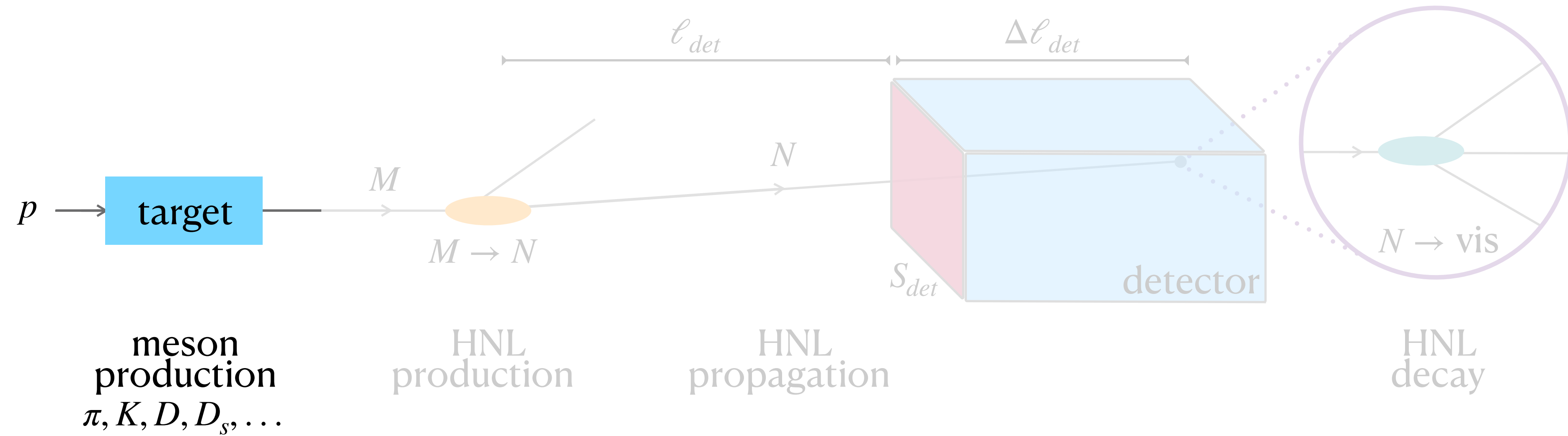


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$$N_{ev} = N_M \underbrace{\text{BR}(M \rightarrow N)}_{\text{HNL production}} \underbrace{\text{BR}(N \rightarrow \text{vis})}_{\text{HNL decay}} \epsilon_{det} \int dS \int dE_N P(c\tau_N/M_N, E_N, \Omega_N) \underbrace{\frac{dn^{M \rightarrow N}}{dS dE_N}}_{\text{model-dependent quantities}}$$

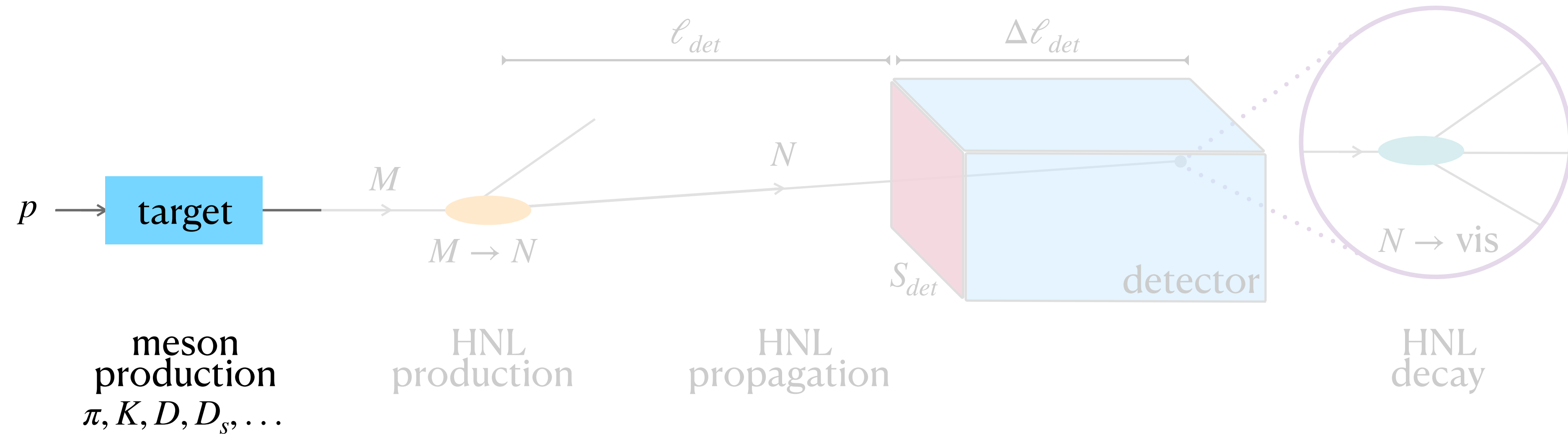
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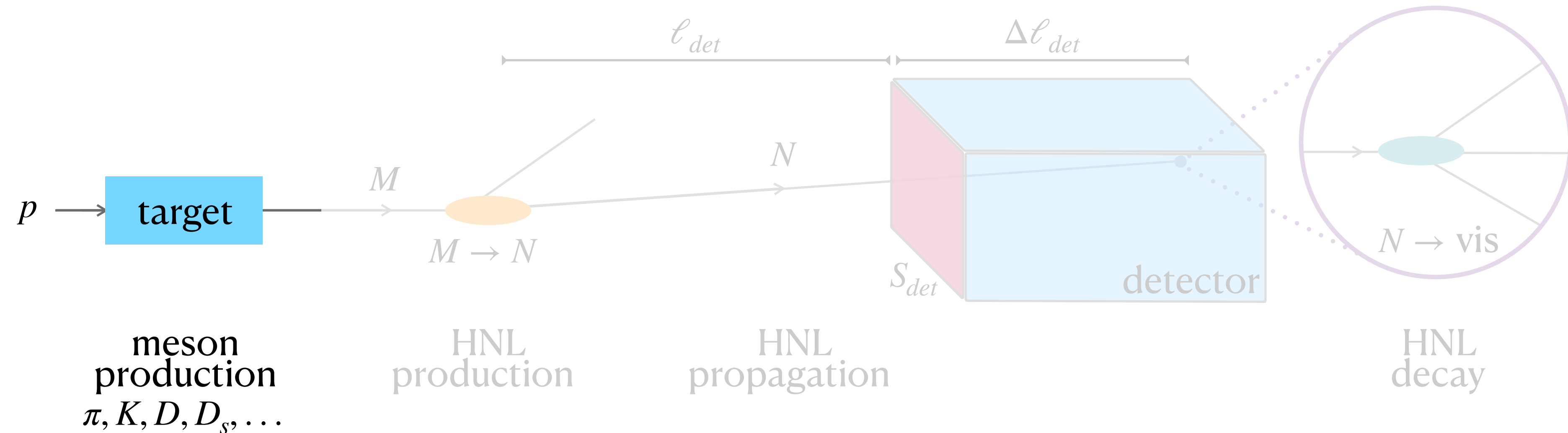
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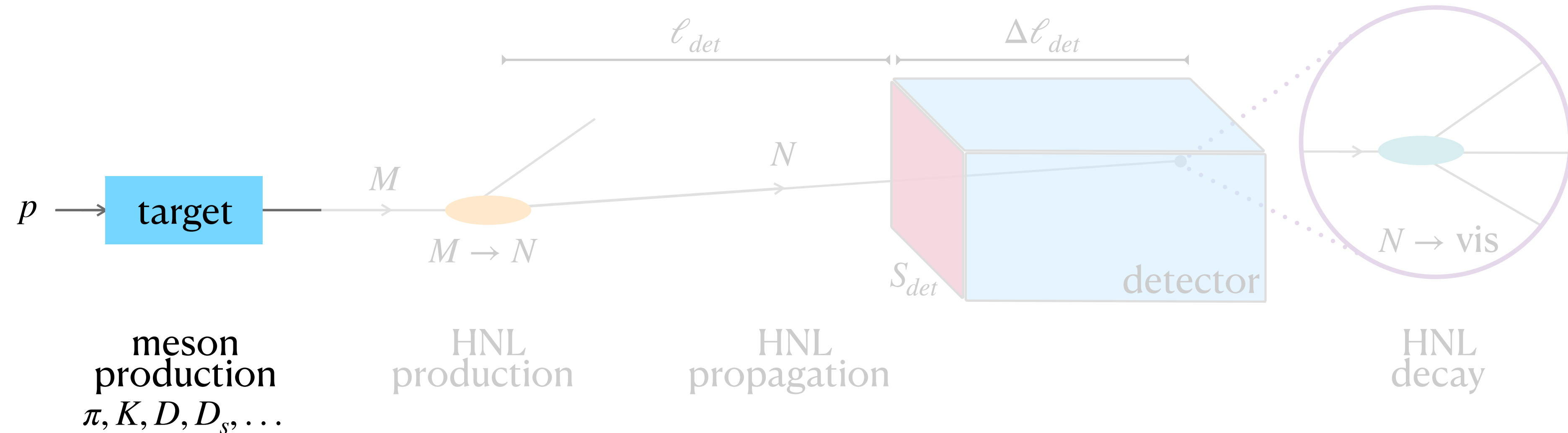
- Meson production at the target station: we consider the mesons with relevant branching fraction to neutrinos
  - $\pi^\pm$  &  $K^\pm$ : GEANT4 simulations including the interaction of 120 GeV protons on the graphite target, the production of mesons, and the re-interaction and focusing in the magnetic horns (G4NuMI)

L. Aliaga Soplin, Neutrino Flux Prediction for the NuMI Beamline, Ph.D. thesis, William-Mary Coll. (2016)



# The simulations

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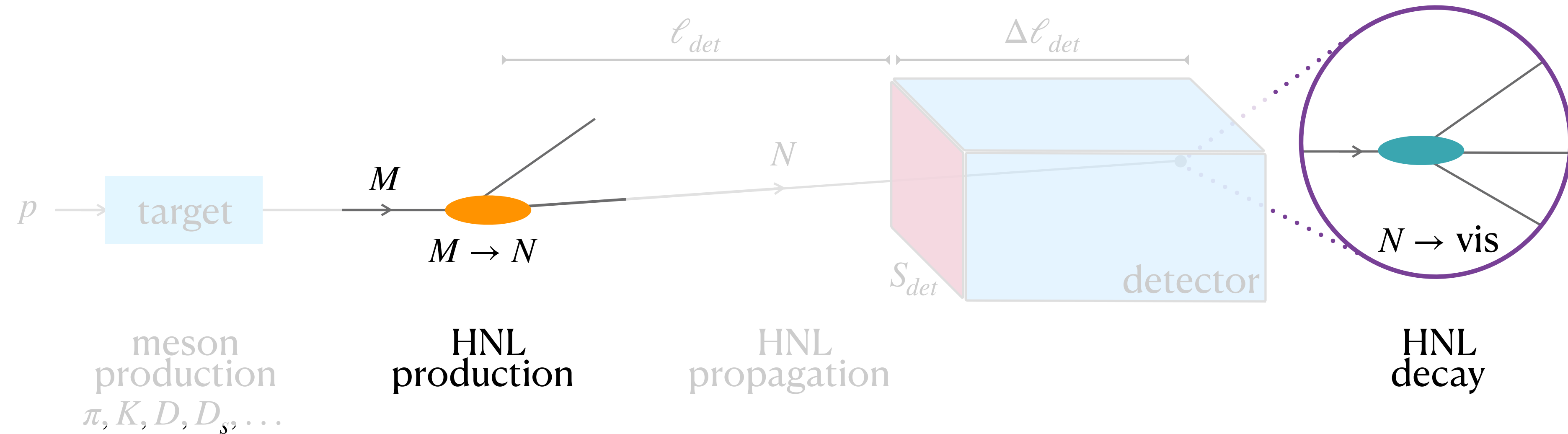
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  - $D^\pm, D_s^\pm$  &  $\tau^\pm$ : simulated with GEANT4 + Pythia8

P. Coloma, E. Fernandez-Martinez, M. Gonzalez-Lopez, J. Hernandez-Garcia, Z. Pavlovic, Eur.Phys.J.C **81** (2021) 1, 78

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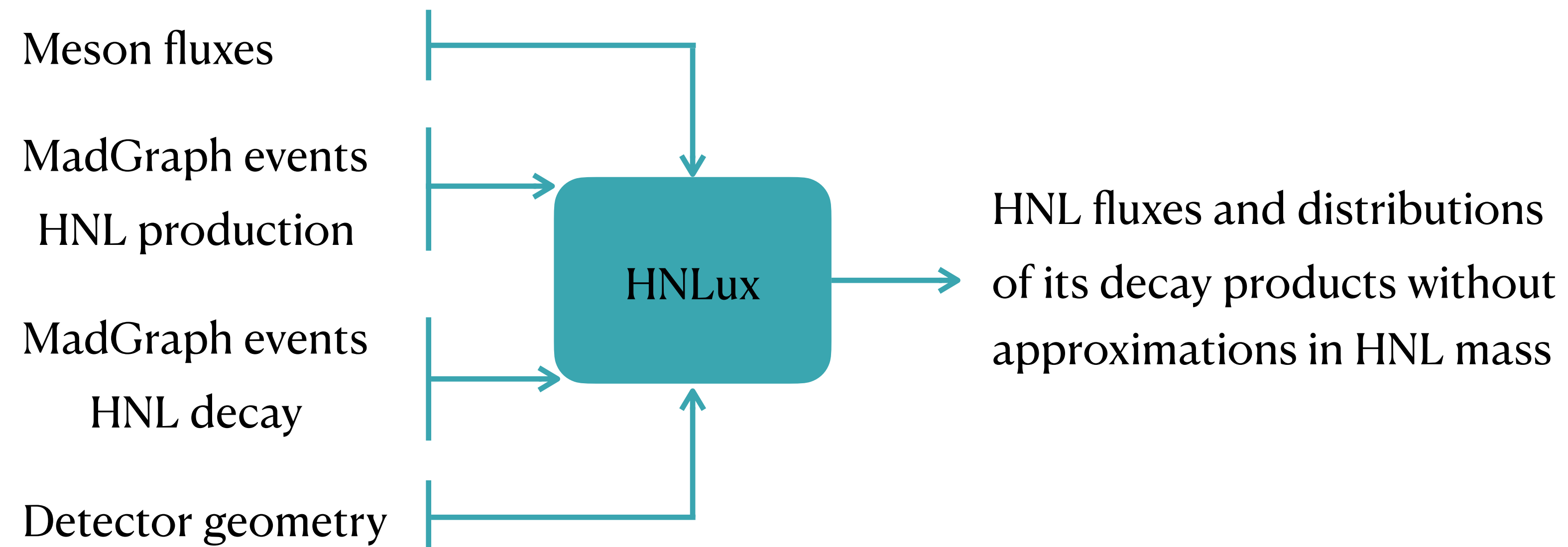
- Meson production at the target station: we consider the mesons with relevant branching fraction to neutrinos
- MonteCarlo simulations of HNL production and HNL decay:

We use a FeynRules model file that contains effective op. describing interactions between light mesons and HNL. Interfacing the model file with an event generator allows for the simulation of fully differential event distributions.

<https://feynrules.irmp.ucl.ac.be/wiki/HNLs>

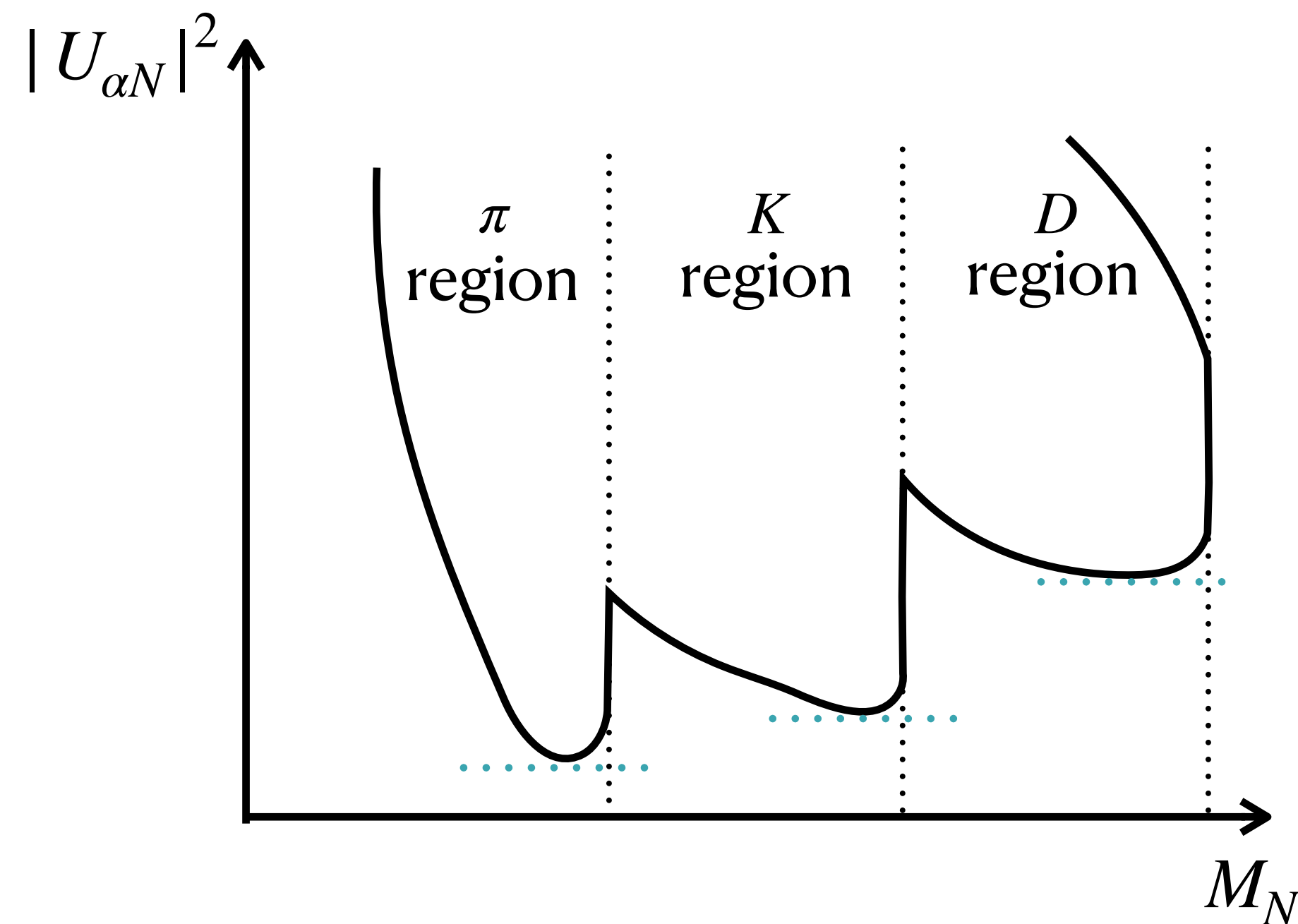
# The simulations

The simulations of HNL fluxes were performed with HNLux:



# The shape of the sensitivity line

The HNL sensitivity line shows a characteristic shape coming from:

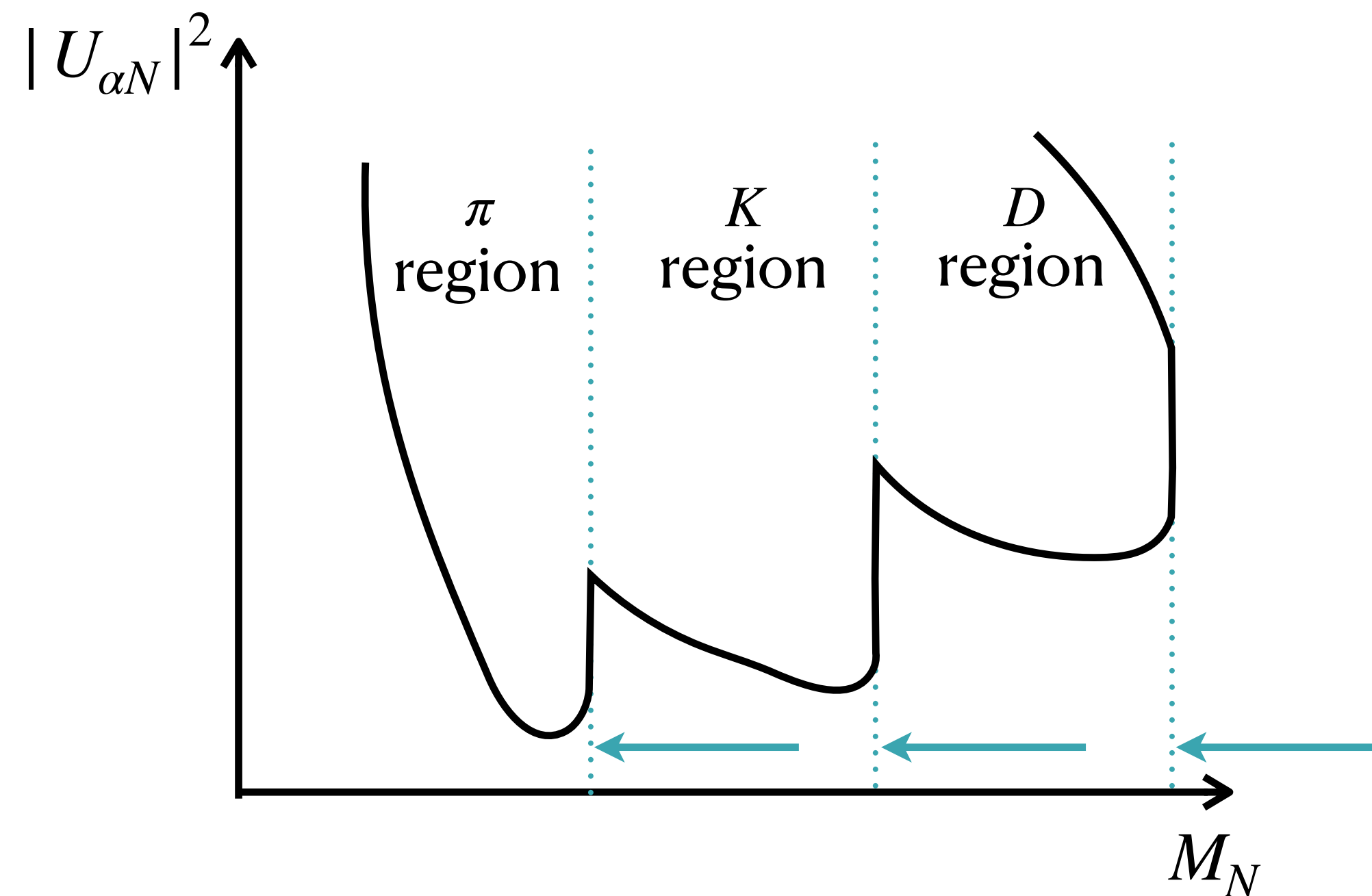


- Meson production yield  $N_M$ 
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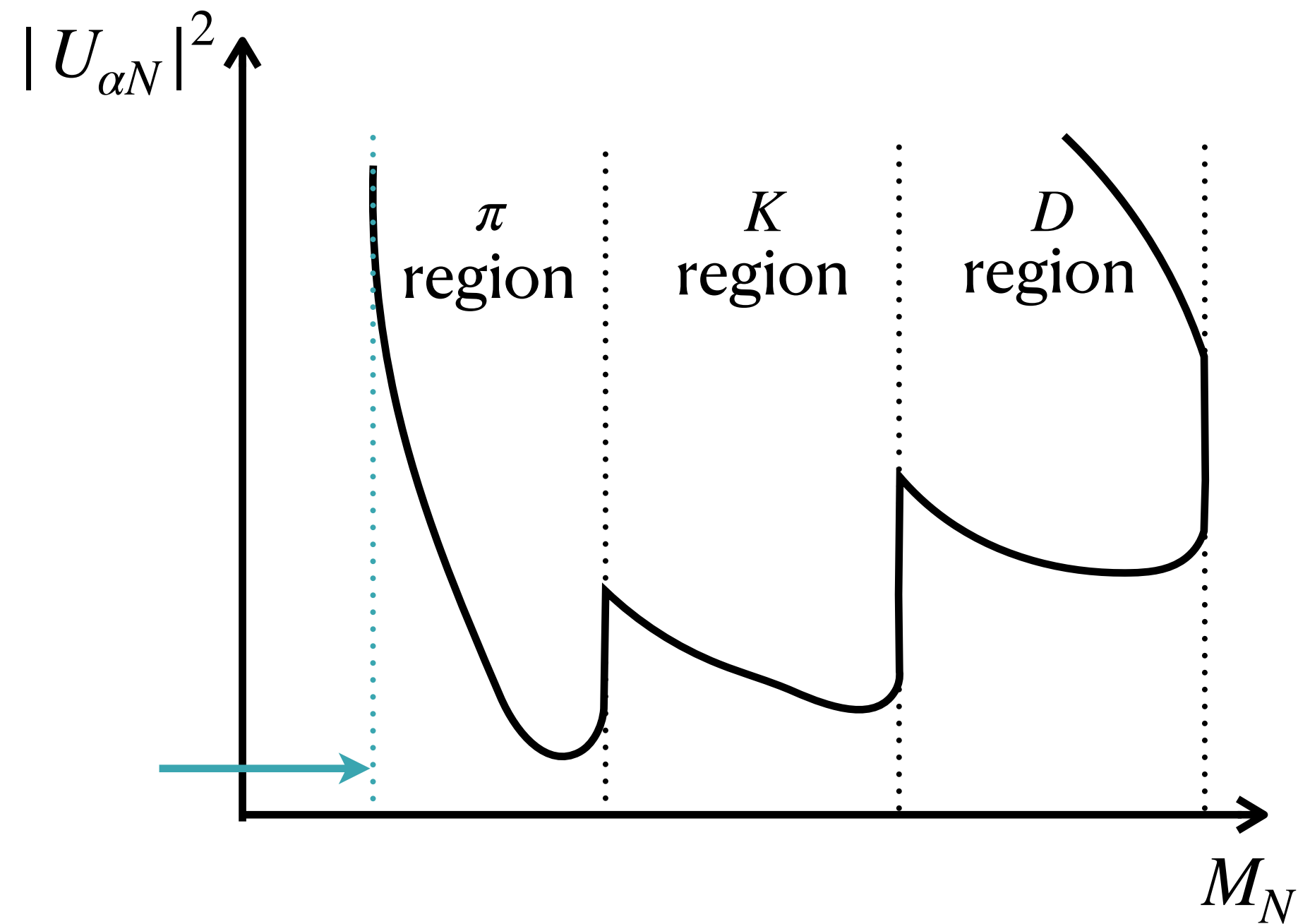
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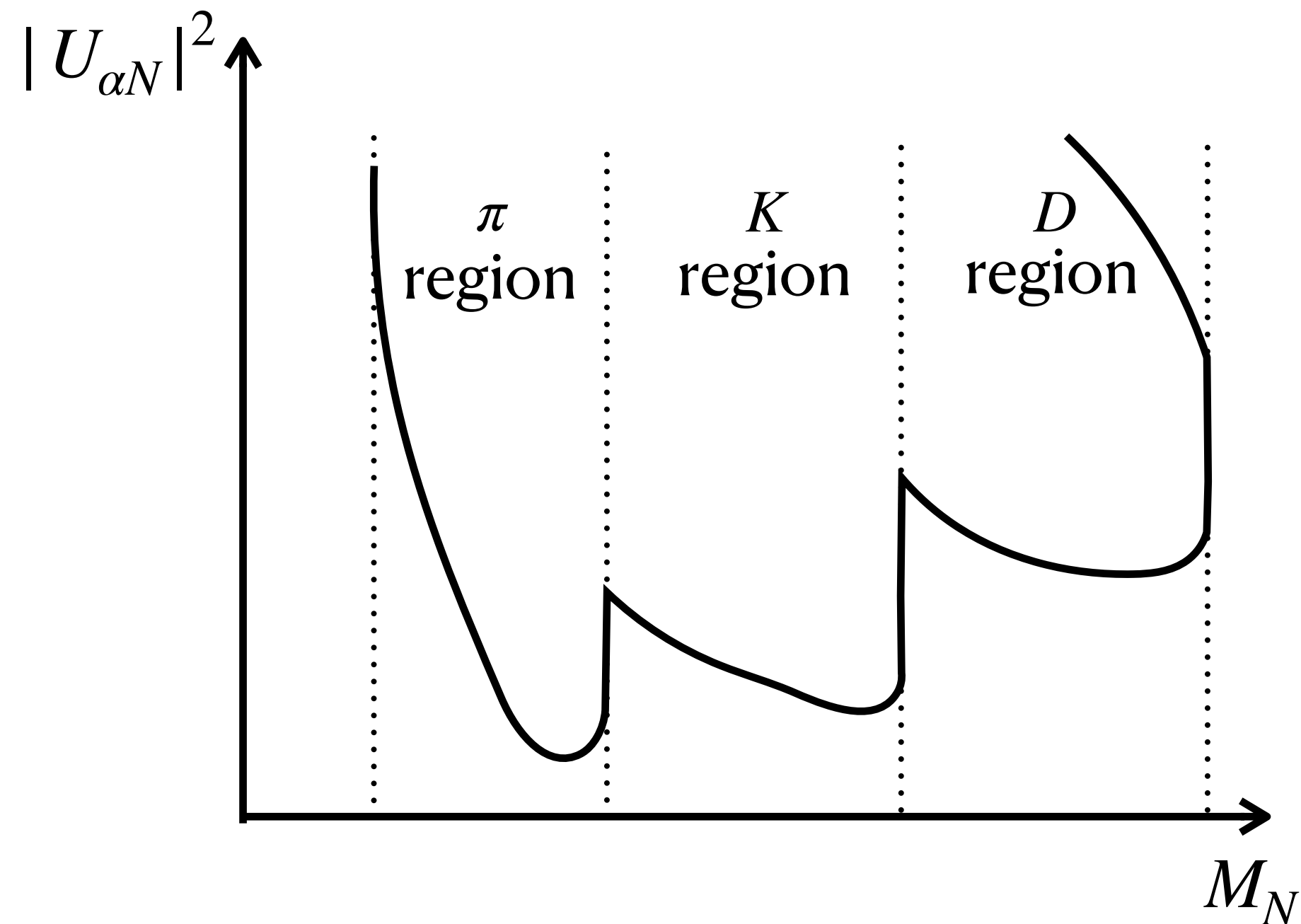
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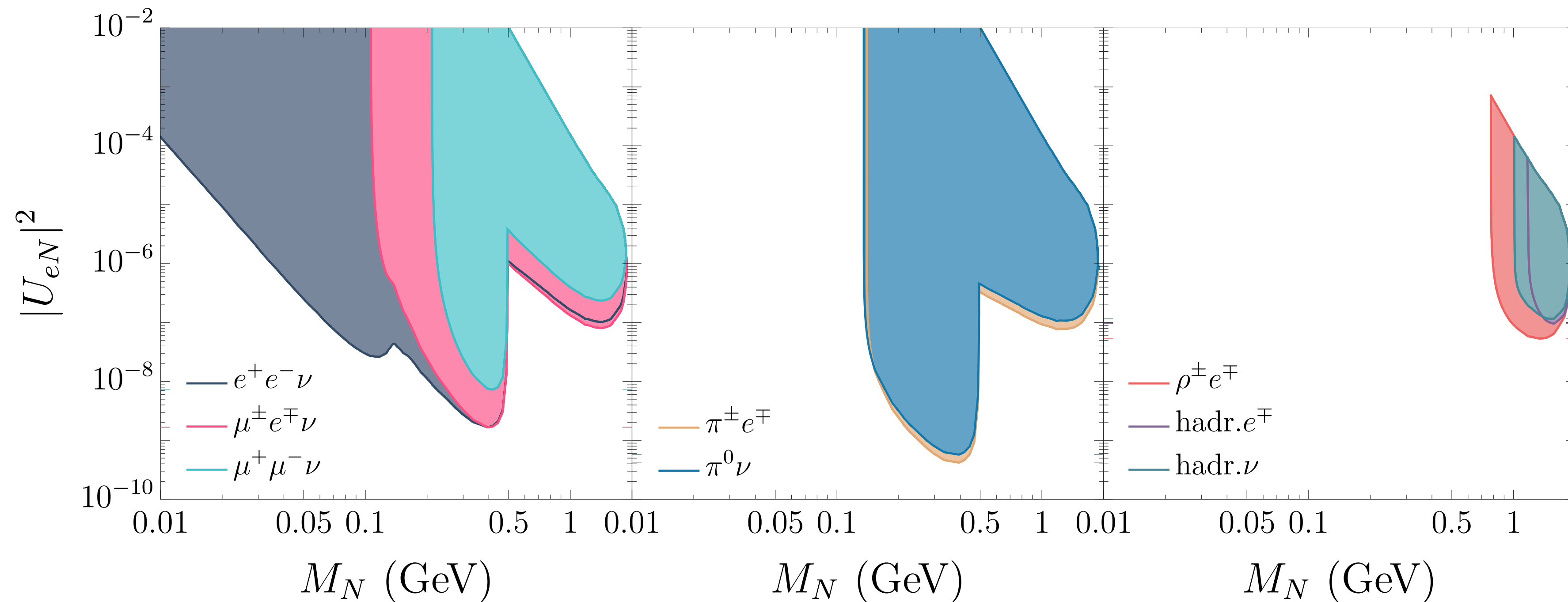
$|U_{\tau N}|^2$  only via  $D/D_s \rightarrow \tau$ , and  $\tau$  decays  $\Rightarrow$  less sensitivity due to suppressed production yield



# HNL sensitivity: signal only

Following the Feldman and Cousins prescription for a Poisson distribution with no background:

$$N_{ev}^{ex} < 2.3 \Rightarrow 90\% \text{ CL}$$



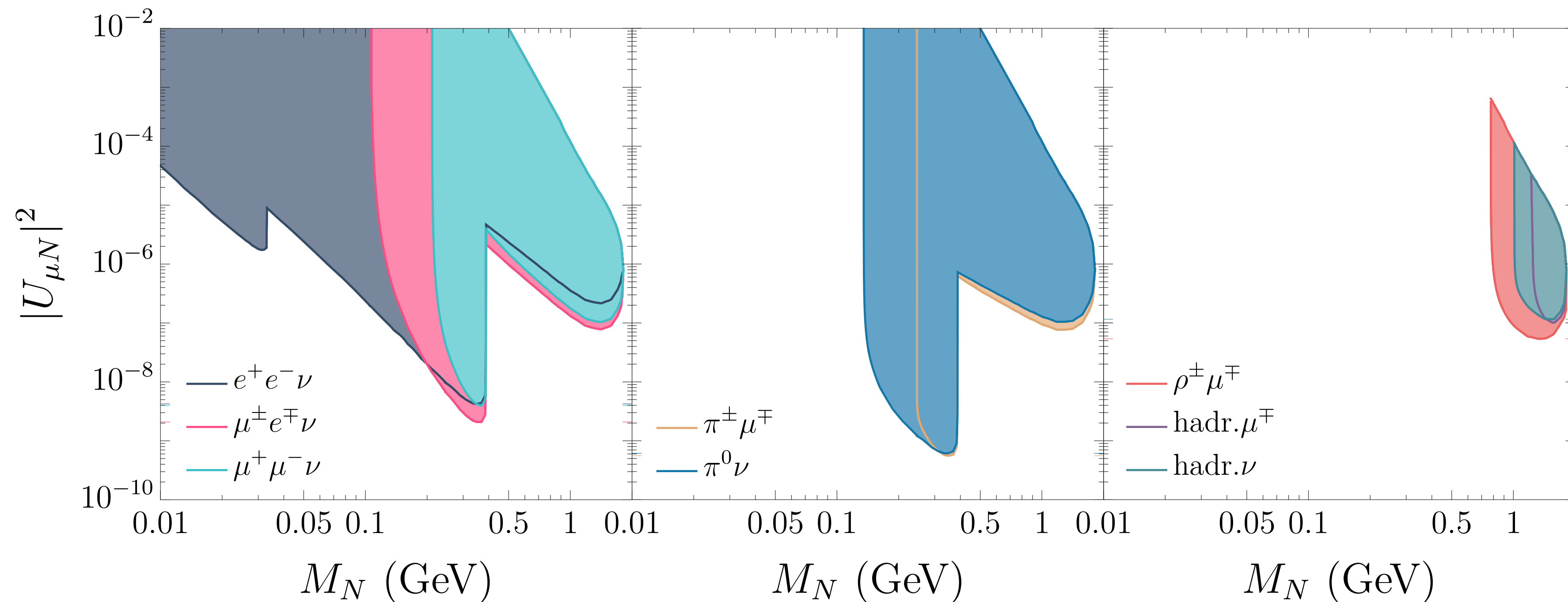
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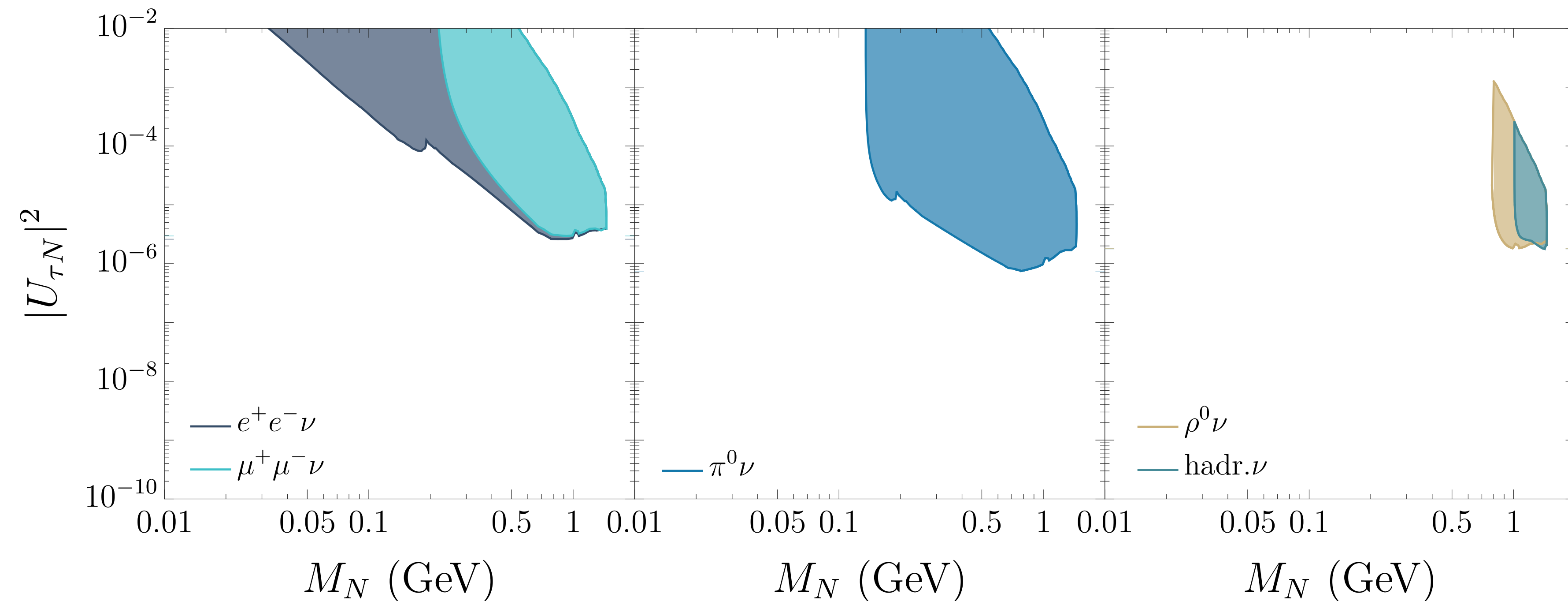


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# The impact of the background

The dominant background comes from the CC and NC SM neutrino interactions with the  $^{40}\text{Ar}$ .

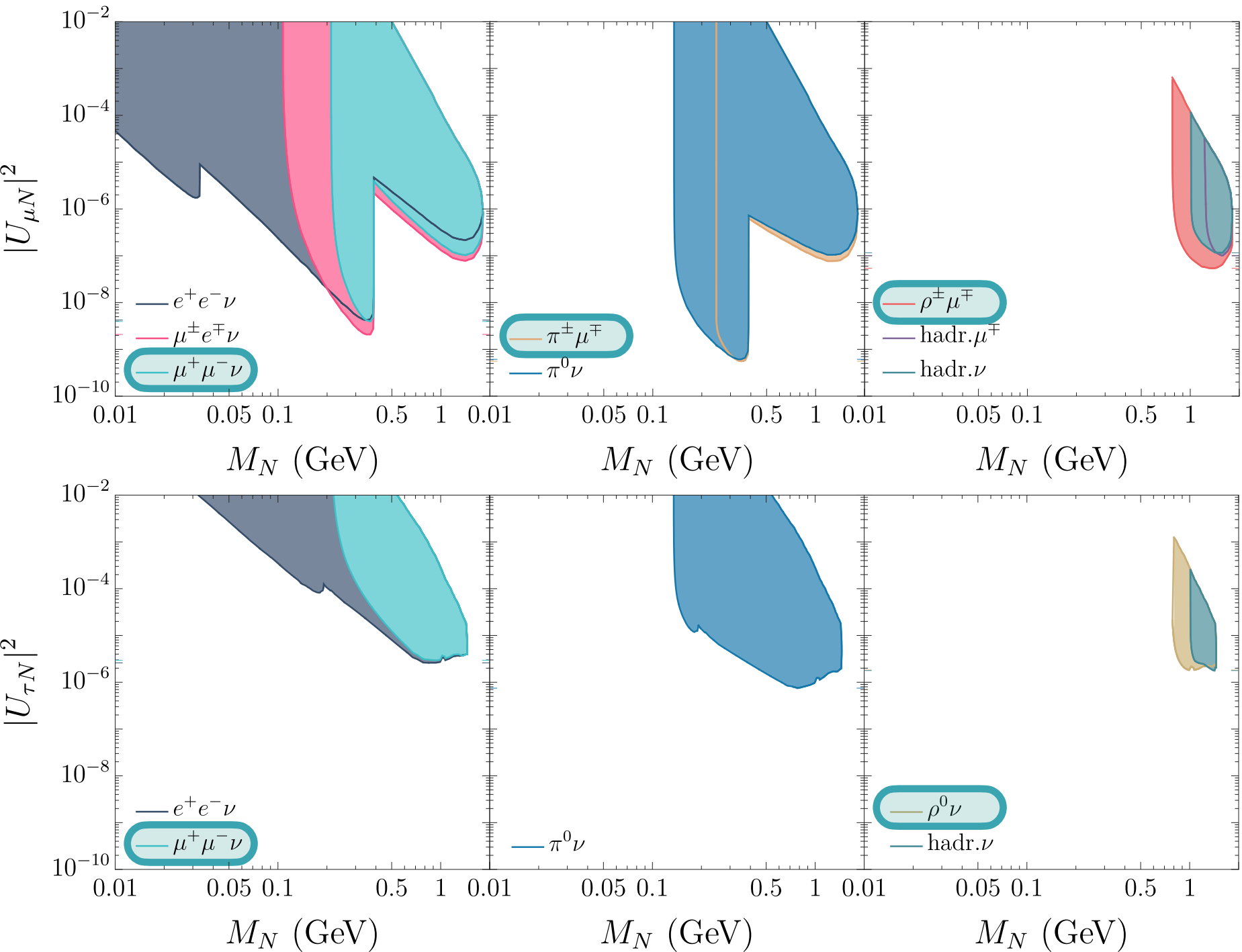
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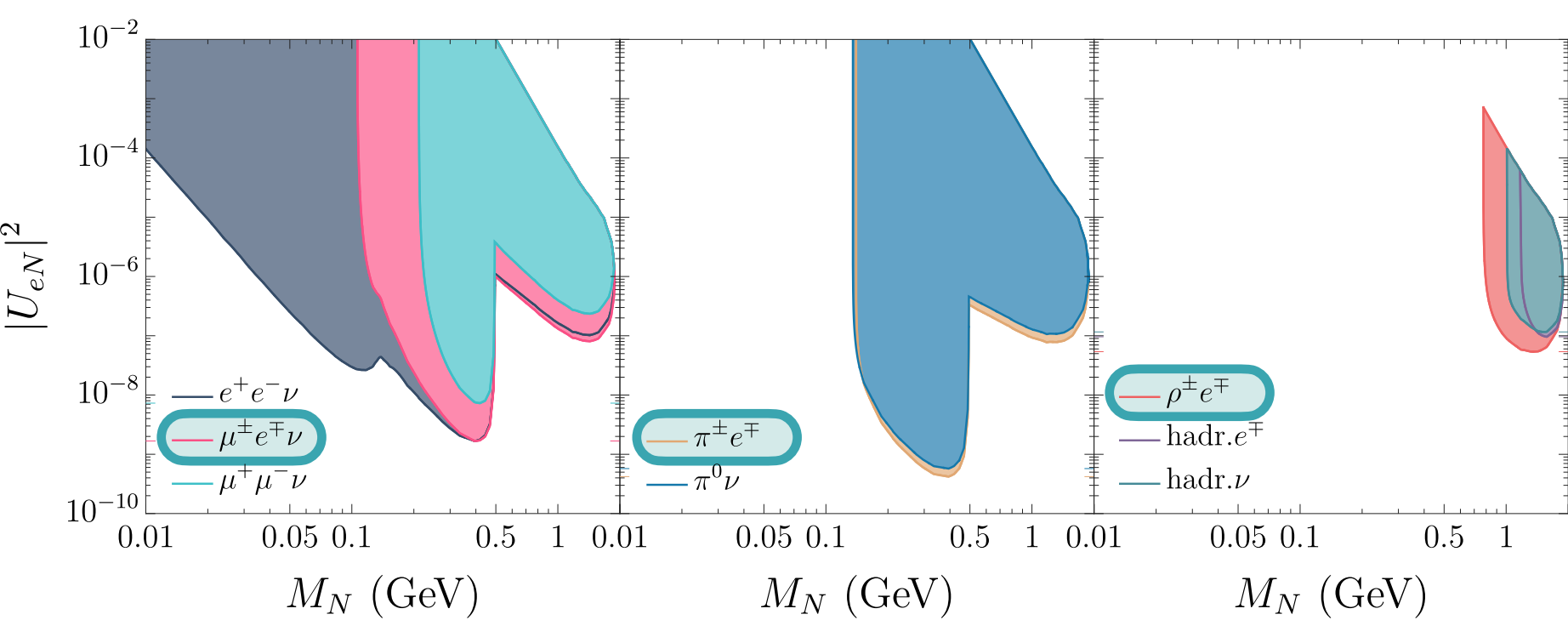
Decay channel	Main source of background
$N \rightarrow \pi^+\mu^-$	Single pion production from $\nu_\mu$ resonant or coherent nucleus scattering
$N \rightarrow \rho^+\mu^-, \pi^+\pi^0\mu^-$	
$N \rightarrow \rho^0\nu, \pi^+\pi^-\nu$	
$N \rightarrow \mu^+\mu^-\nu$	Charm production in CC deep inelastic $\nu_\mu$ nucleus scattering ( <b>suppressed</b> )

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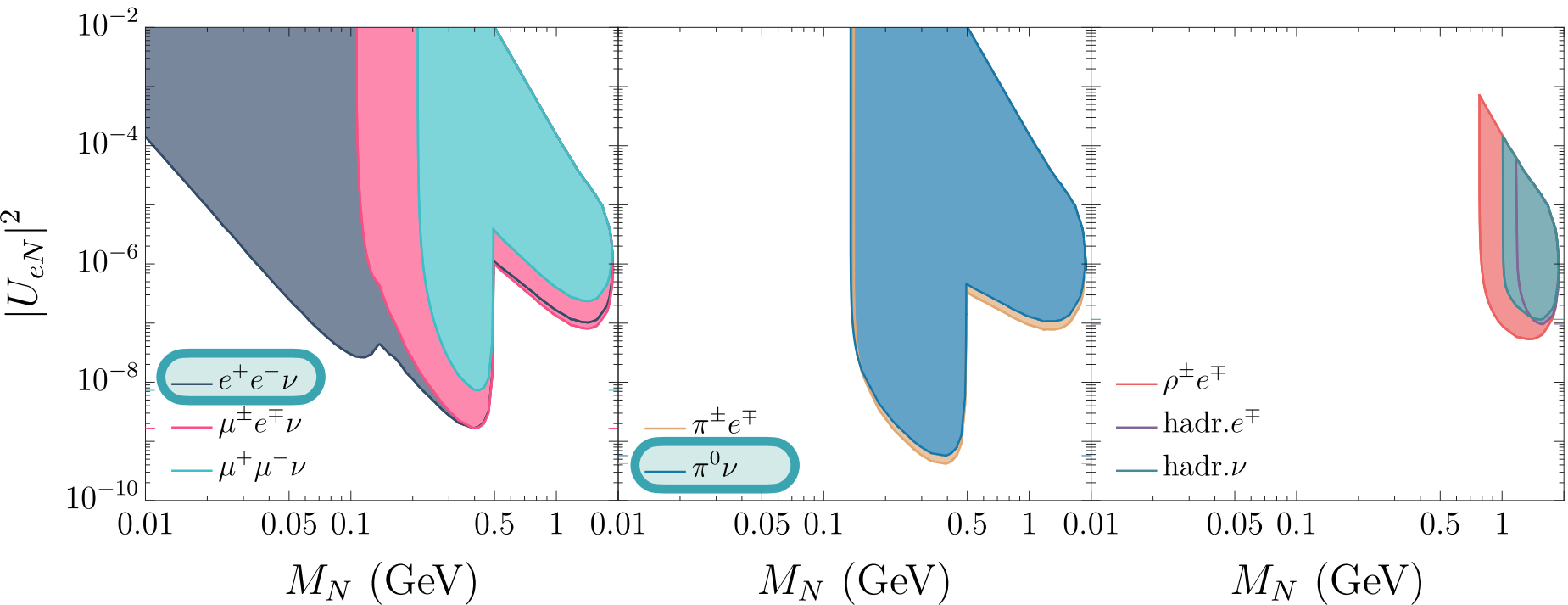
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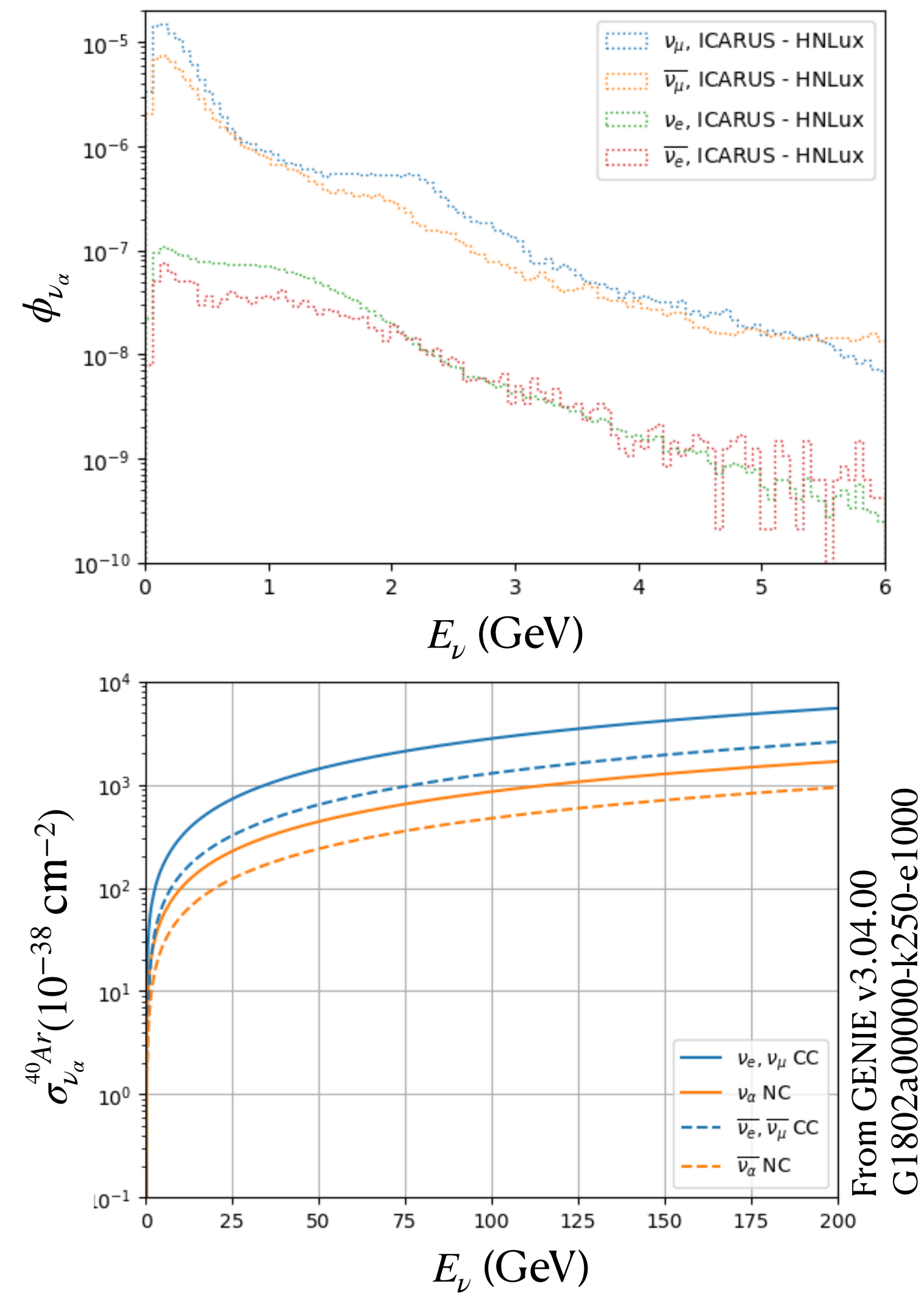
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# $\nu_\alpha$ - $^{40}\text{Ar}$ number of interactions



Number of  $\nu_\alpha$ - $^{40}\text{Ar}$  interactions in the detector given by

$$n_{ev} = n_t \int \phi_{\nu_\alpha} \sigma_{\nu_\alpha}^{40\text{Ar}}$$

number of targets ICARUS

In one year:

	CC $\nu_\alpha$ - $^{40}\text{Ar}$	NC $\nu_\alpha$ - $^{40}\text{Ar}$
$\alpha = e$	19142	2109
$\alpha = \mu$	341257	42610

# Kinematical cuts

Momentum distributions of signal and background very different.

HNLs come from meson decays produced at the target  $\Rightarrow$  HNLs highly boosted:

- Low reconstructed transverse momentum  $p_T$
- Small angle  $\theta$  between  $\mu$ -like tracks

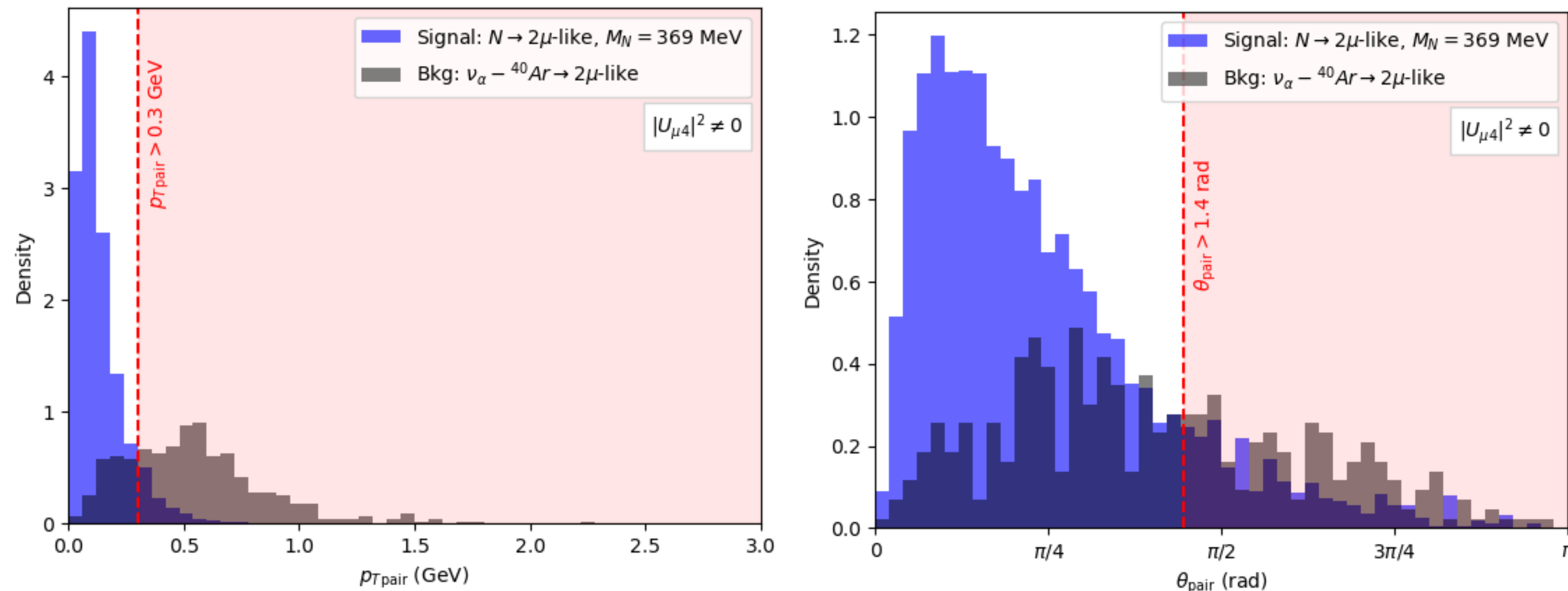


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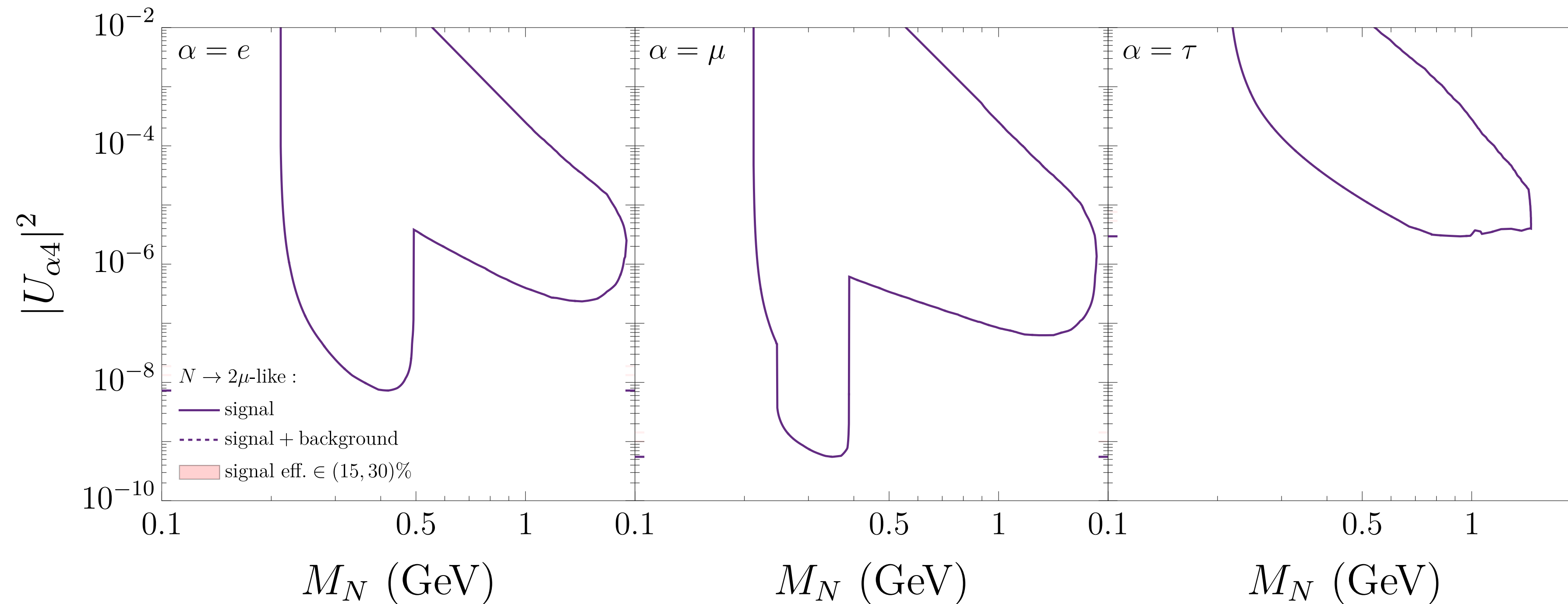


By applying kinematical cuts:

- Bkg. rej. eff. up to  $\simeq 90\%$
- Signal eff. up to  $\simeq 92\%$

# HNL sensitivity: signal + background

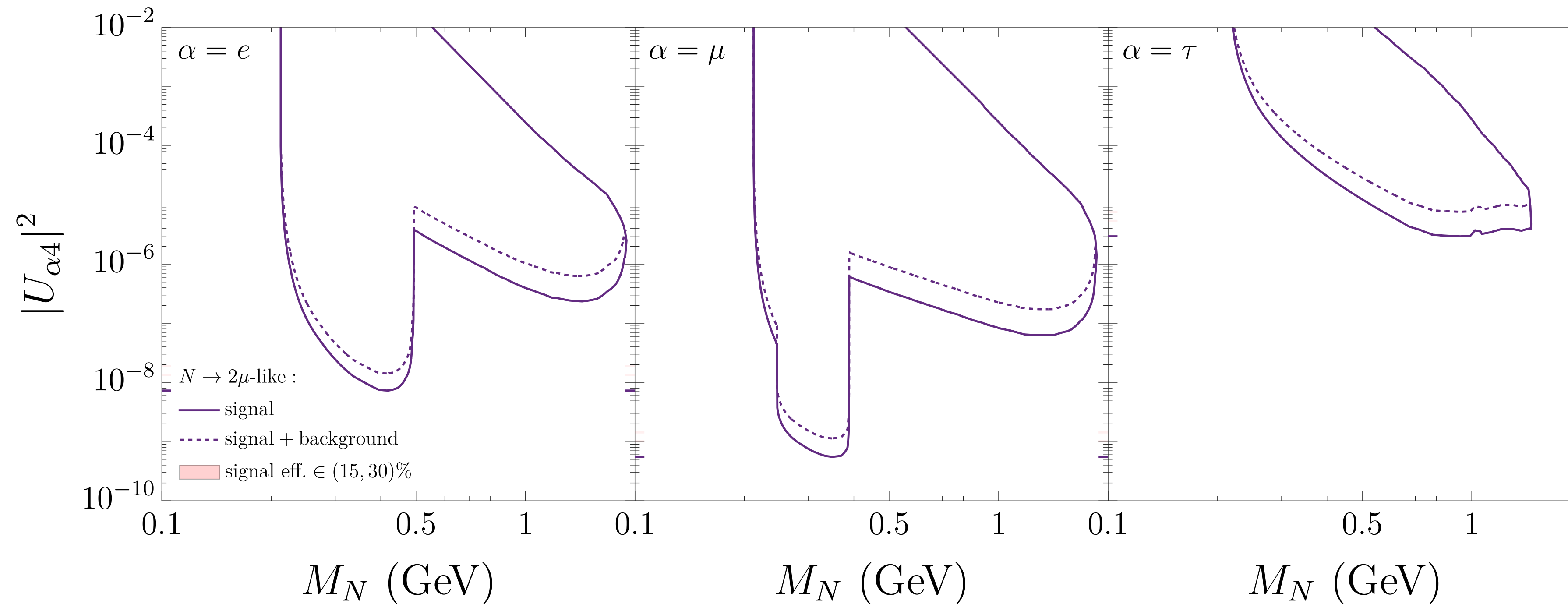
Combining the channels with **two**  $\mu$ -like tracks



- Solid purple line: F&C analysis assuming no background

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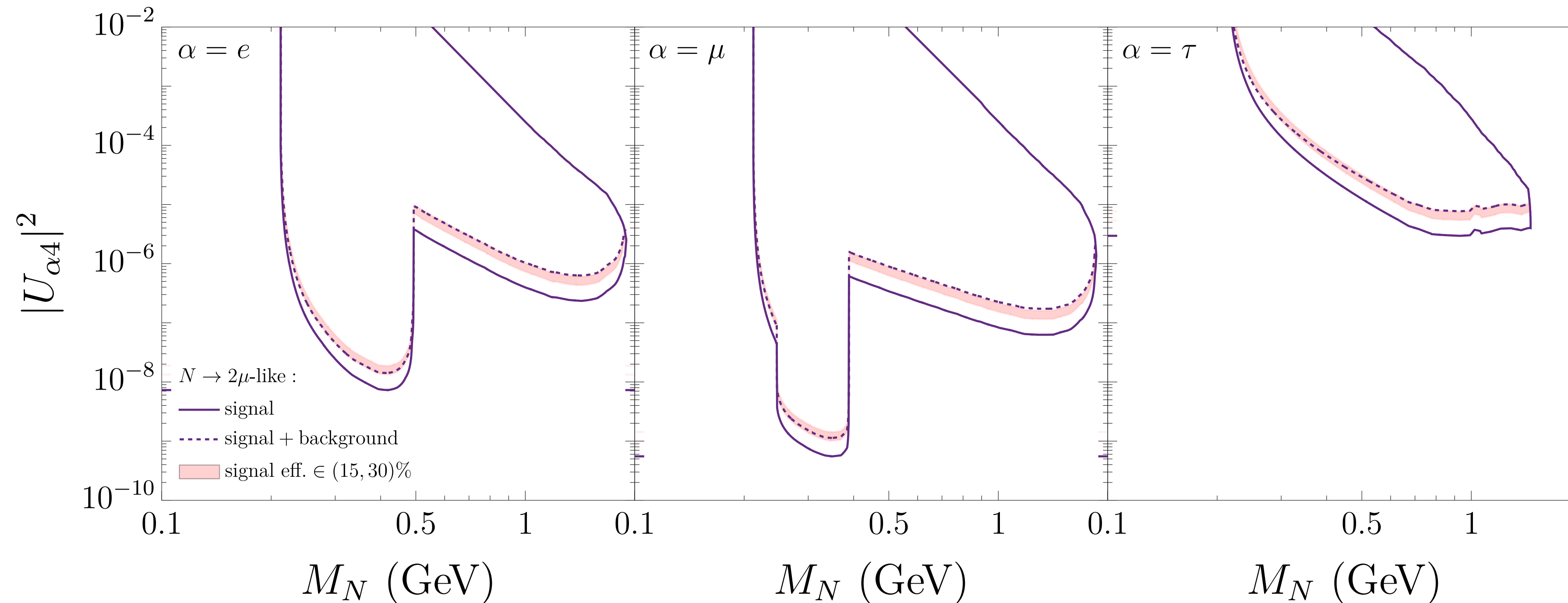
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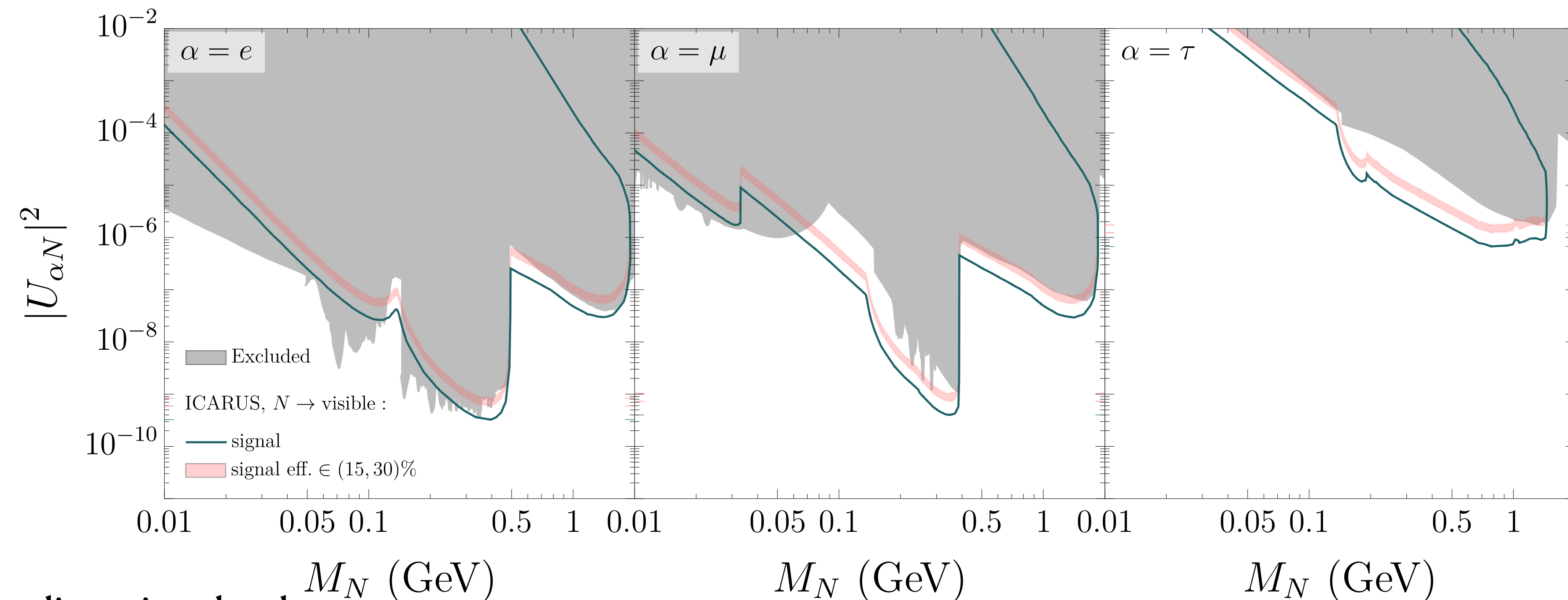
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- Pink band: (15, 30) % signal efficiency cut to the negligible background case that covers the signal + bkg result

# HNL sensitivity: ICARUS over present constraints

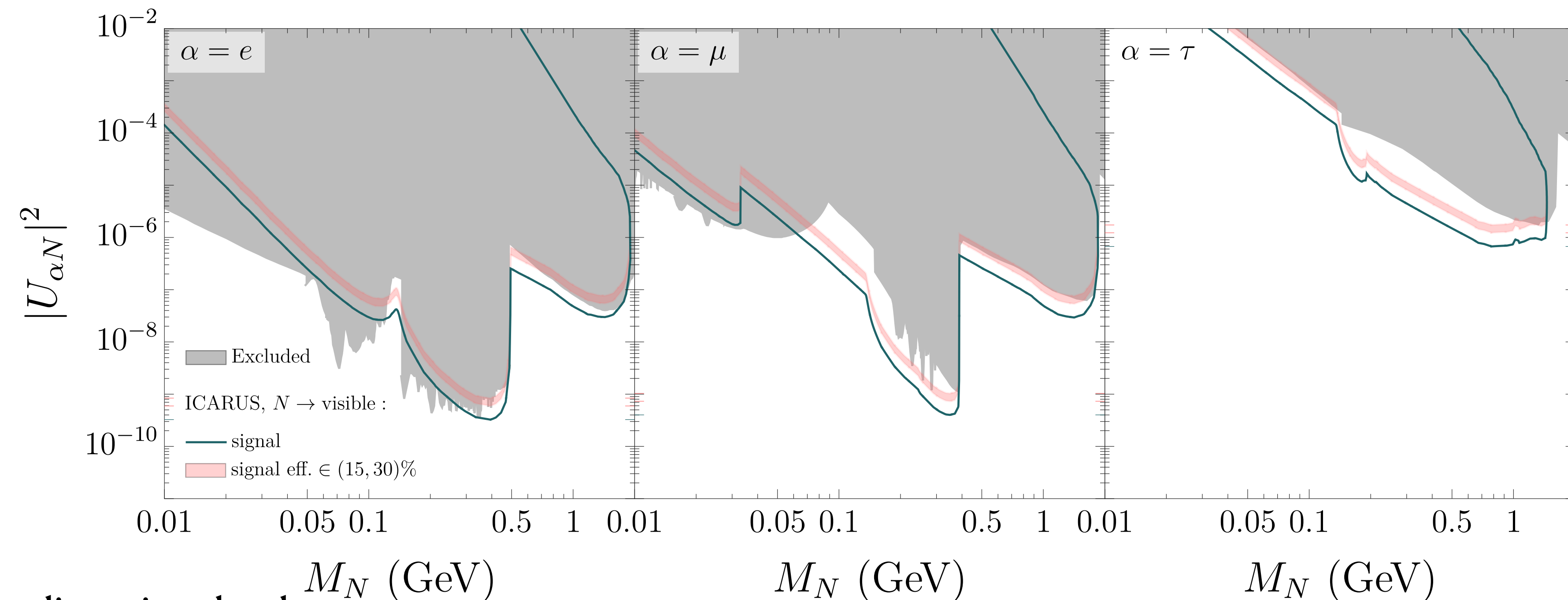
Combining leading HNL decay channels and following the Feldman and Cousins prescription  $\Rightarrow$   
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Sensitivity results compatible with arXiv: 2409.04394



# Summary

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**Thanks!**



This research is supported by the EU H2020 research and innovation programme under the MSC grant agreement No 860881-11 HIDDDeN, as well as by the Spanish Ministerio de Ciencia e Innovación project PID2020-113644GB-I00, the Spanish Research Agency (Agencia Estatal de Investigación) through the project CNS2022-136013, and Severo Ochoa Excellence Program CEX2023-001292-S.

**Back-up**

# Status of NuMI after 2024 power failure

A **major** power failure of a transformer on the NuMI beam line at the end of August 2024 stopped its activity.

As a result, NuMI is not running, and its continuation until the end of 2026 is **not** yet clear.

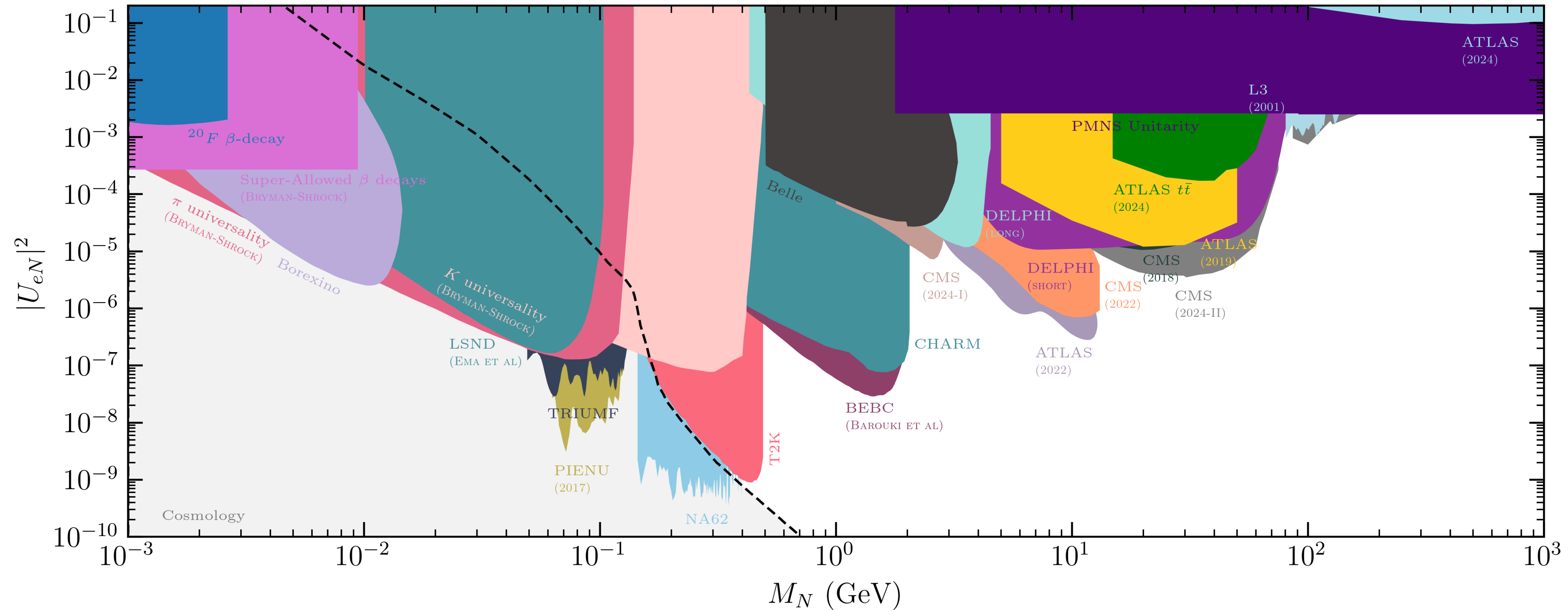
From 2027 on, Fermilab will **focus** on the LBNF beam line for DUNE.

ICARUS has **so far** collected  $\sim 6 \cdot 10^{20}$  PoT  $\Rightarrow$  roughly **half** of what was assumed for our results ( $1.32 \cdot 10^{21}$  PoT).

If NuMI **resumes** its activity during 2026, ICARUS could still collect  $\sim 1.32 \cdot 10^{21}$  PoT.

Otherwise, our sensitivity results should be **corrected** by a factor of  $\sim \sqrt{2}$ .

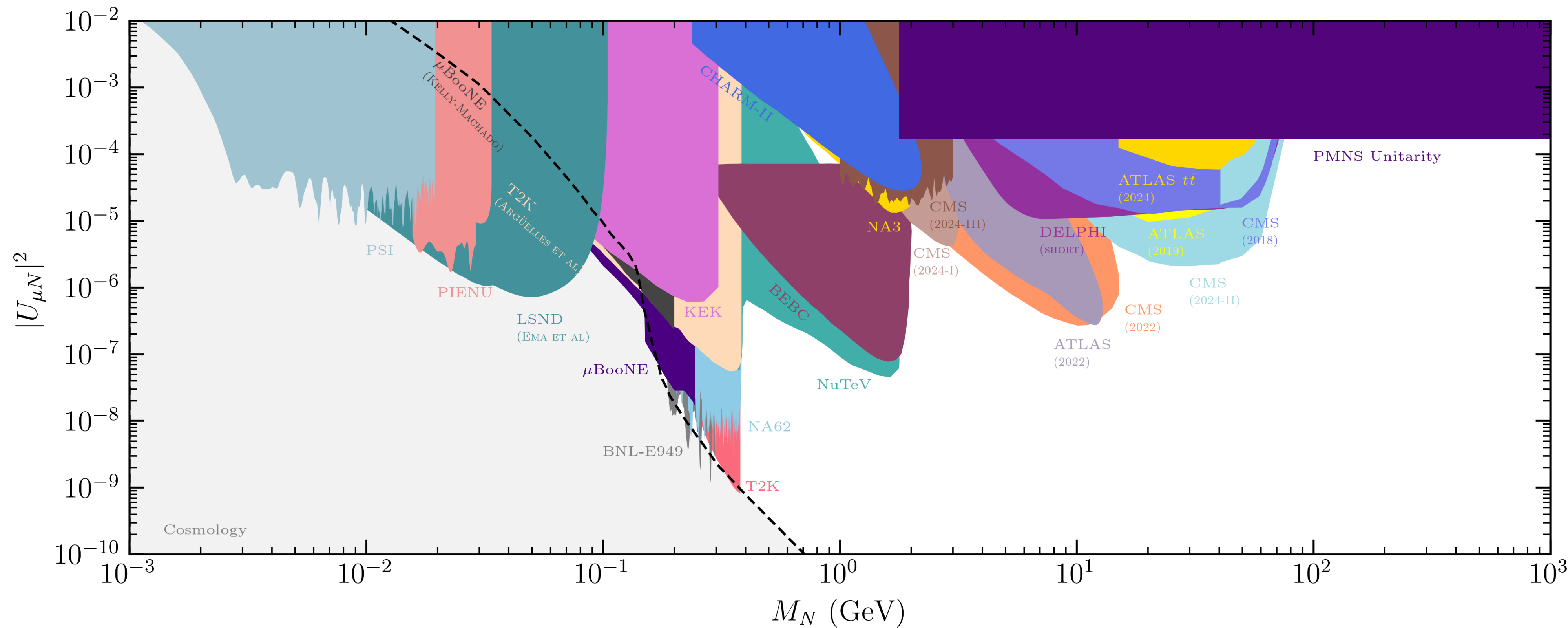
# Heavy Neutral Leptons: present bounds



Present bounds in the HNL mixing from  $\nu$  oscillations, kinks in beta-decay, peaks in meson decays, beam-dumps, colliders, and non-unitarity of the PMNS.



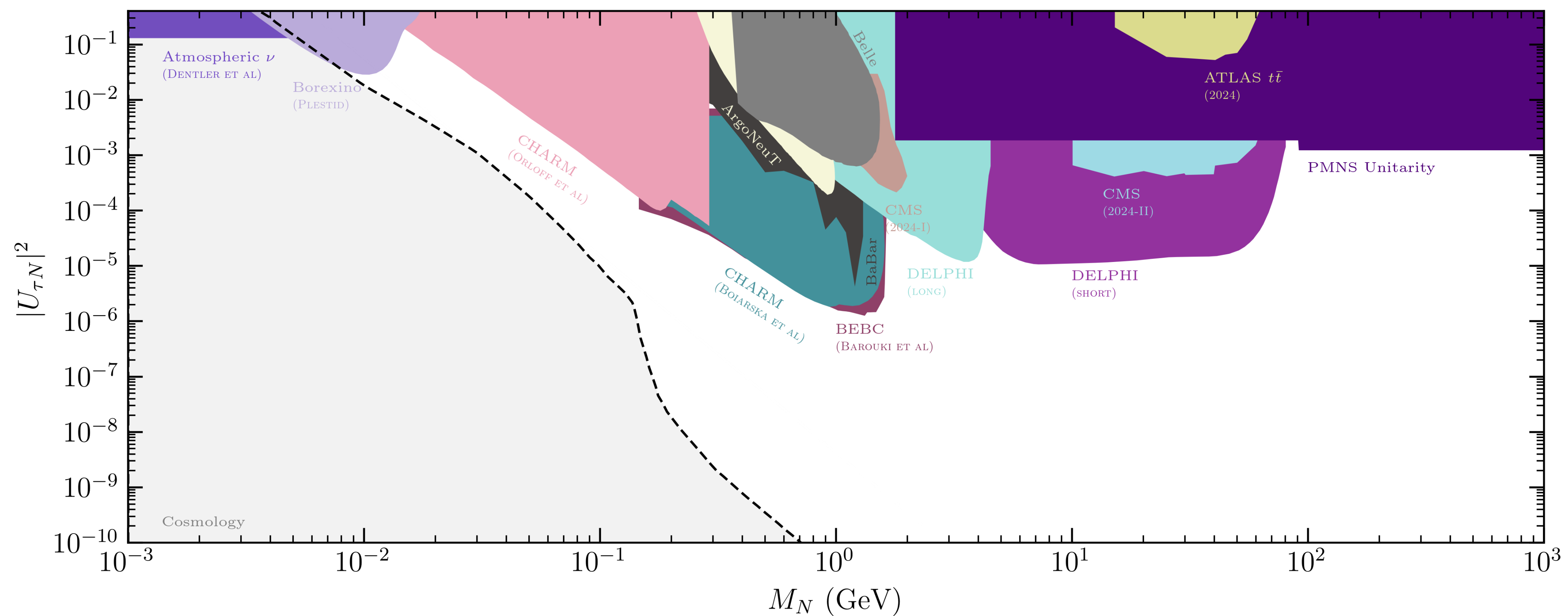
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# Heavy Neutral Leptons: present bounds



Present bounds in the HNL mixing from  $\nu$  oscillations, kinks in beta-decay, peaks in meson decays, beam-dumps, colliders, and non-unitarity of the PMNS.



# Heavy Neutral Leptons: production and decay

- HNL production

Leptonic and semileptonic meson decays

	$\pi \rightarrow$	$K^0 \rightarrow$	$K \rightarrow$	$D \rightarrow$	$D_s \rightarrow$	$\tau \rightarrow$
$ U_{eN} ^2$	$eN$	—	$eN$	$eN$	$eN$	—
	—	$\pi^0 eN$	$\pi eN$	$K^0 eN$	—	—
$ U_{\mu N} ^2$	$\mu N$	—	$\mu N$	$\mu N$	$\mu N$	—
	—	$\pi^0 \mu N$	$\pi \mu N$	$K^0 \mu N$	—	—
$ U_{\tau N} ^2$	—	—	—	$\tau N$	$\tau N$	$\pi N$
	—	—	—	—	—	$\pi \pi^0 N$
	—	—	—	—	—	$e \nu N$
	—	—	—	—	—	$\mu \nu N$

	$\pi$	$K$	$D$	$D_s$	$\tau$
$P^+/\text{PoT}$	6.0	1.1	$1.2 \cdot 10^{-5}$	$3.3 \cdot 10^{-6}$	$2.1 \cdot 10^{-7}$
$P^-/\text{PoT}$	4.1	0.4	$1.9 \cdot 10^{-5}$	$4.6 \cdot 10^{-6}$	$3.0 \cdot 10^{-7}$

Table I. Averaged parent production yield at the NuMI target per PoT during its neutrino mode. A pre-selection of light mesons that would produce neutrinos towards the region where ICARUS is placed has been done.

- HNL decay

Visible HNL decay channels with higher branching ratio

	$N \rightarrow$				
$ U_{eN} ^2$	$ee\nu$	$\mu\mu\nu$	$e\mu\nu$	$e \text{ hadr.}$	$\nu \text{ hadr.}$
	$\nu\pi^0$	$e\pi$	$e\rho$	—	—
$ U_{\mu N} ^2$	$ee\nu$	$\mu\mu\nu$	$e\mu\nu$	$\mu \text{ hadr.}$	$\nu \text{ hadr.}$
	$\nu\pi^0$	$\mu\pi$	$\mu\rho$	—	—
$ U_{\tau N} ^2$	$ee\nu$	$\mu\mu\nu$	—	—	$\nu \text{ hadr.}$
	$\nu\pi^0$	—	—	$\nu\rho^0$	—



# Statistical analysis

- Feldman and Cousins

Prescription for a Poisson distribution with no background and under the hypothesis of no events being observed

$$N^{ex}(\phi) < 2.44, \quad \text{for 90 \% CL}$$

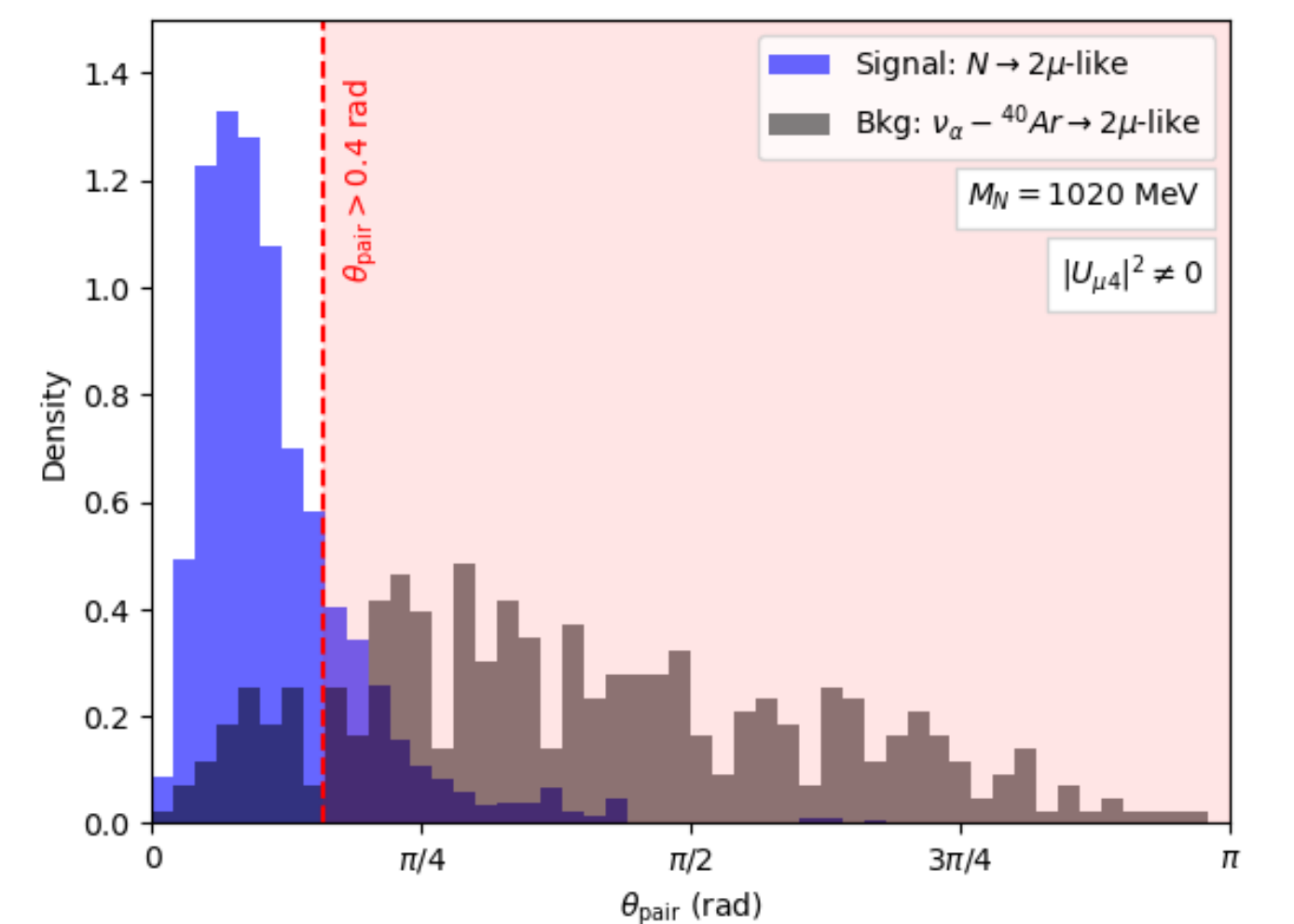
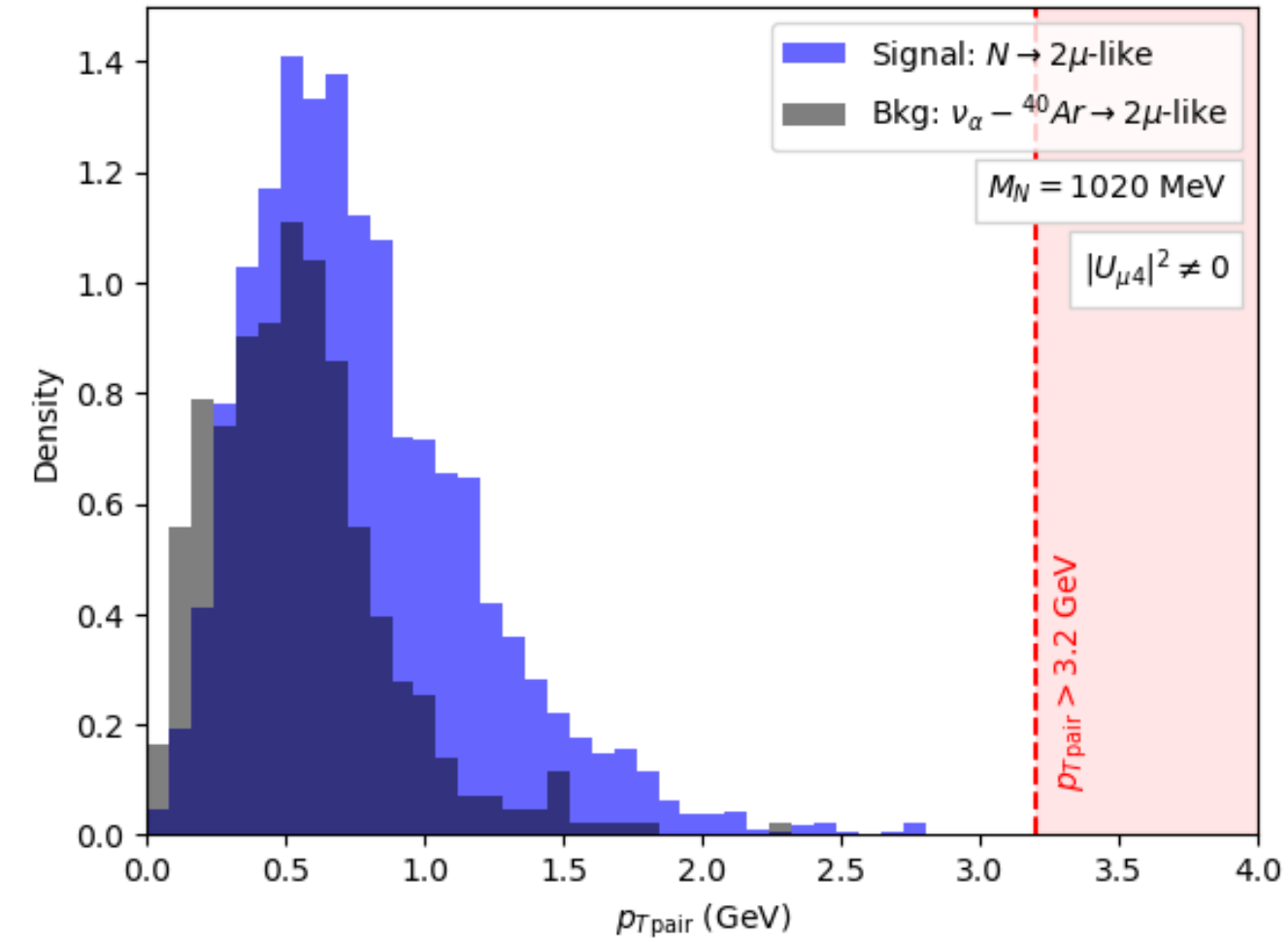
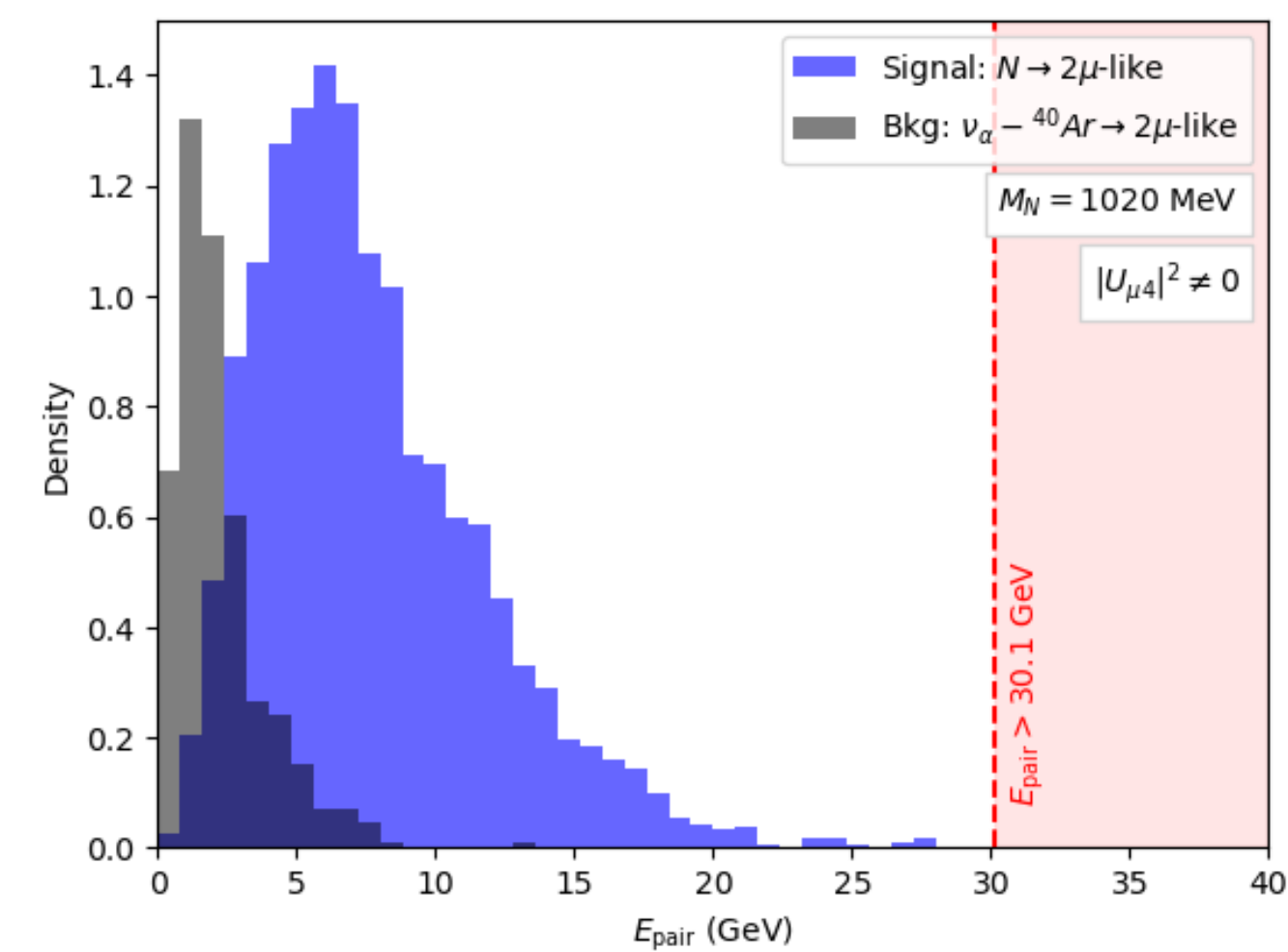
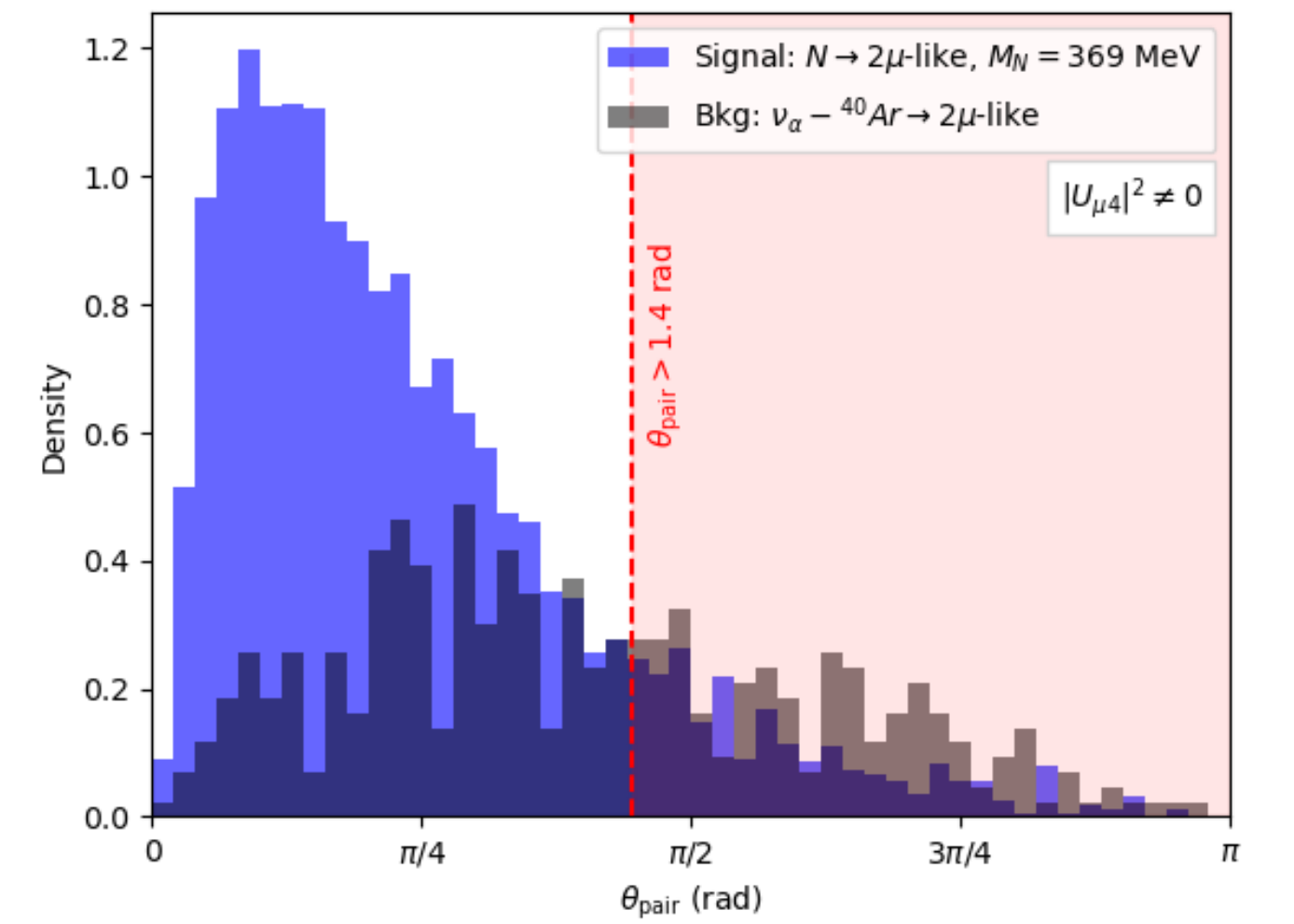
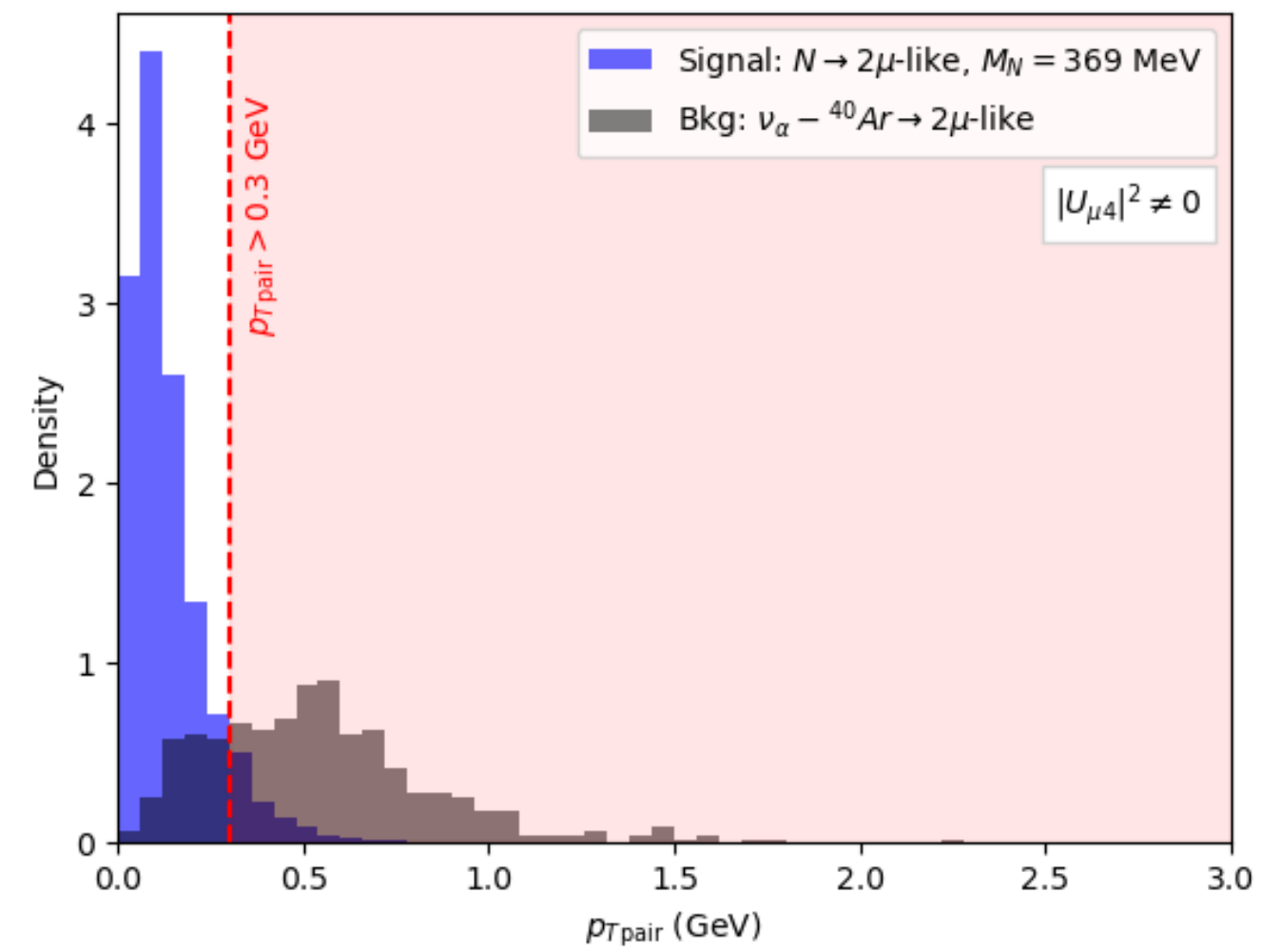
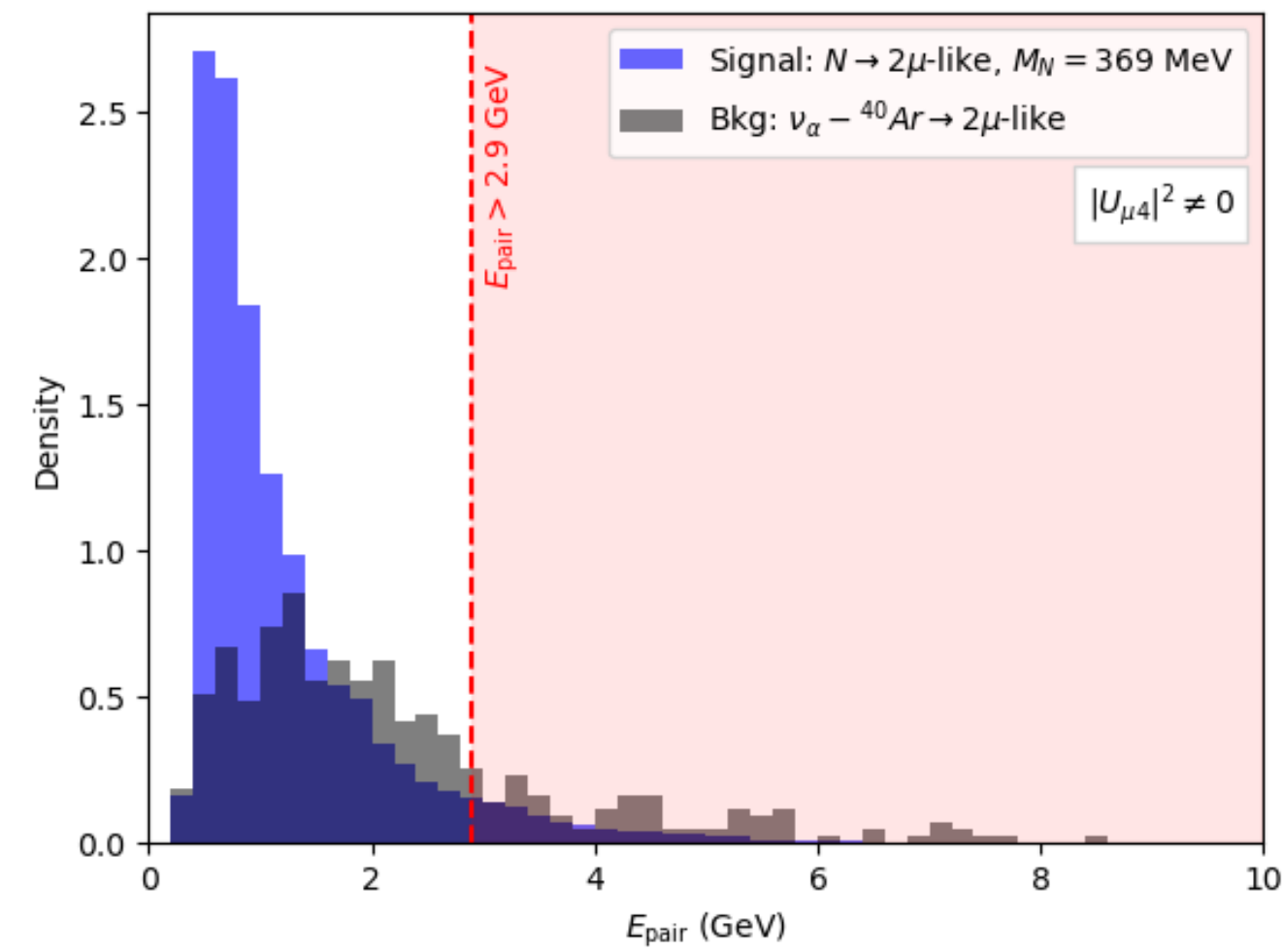
- Unbinned Gaussian  $\chi^2$

$$\chi^2 = \min_{\xi} \left\{ \left( \frac{N^{ex}(\phi, \xi) - N^{ob}(\xi)}{\sigma_N(\xi)} \right)^2 + \left( \frac{\xi}{\sigma_f} \right)^2 \right\}$$

where  $N^{ex}(\phi, \xi)$  is the total expected event rate including both signal and background events, and  $N^{ob}(\xi)$  is the observed events; with the uncertainty on the flux normalization,  $\sigma_f$ , taken as 20 % , and  $\sigma_N(\xi) = \sqrt{N^{ob}(\xi)}$  is the statistical error.

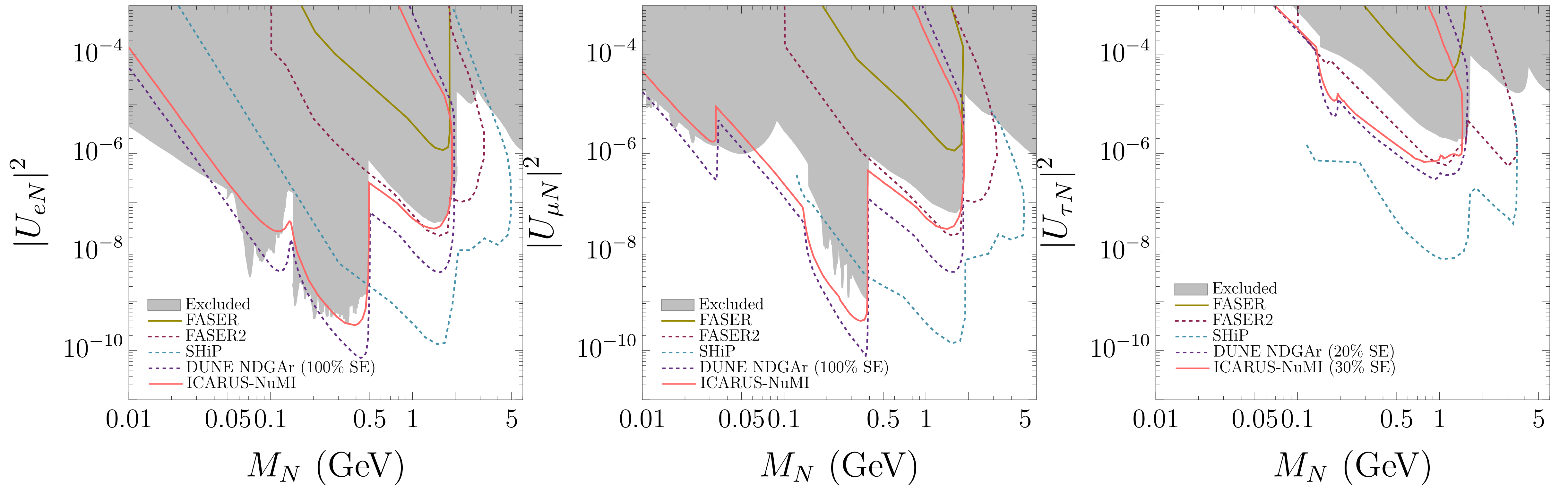
$$\chi^2 < 4.61, \quad \text{for 90 \% CL}$$

# Kinematical cuts



# Present and nHNL sensitivity: present and near-future

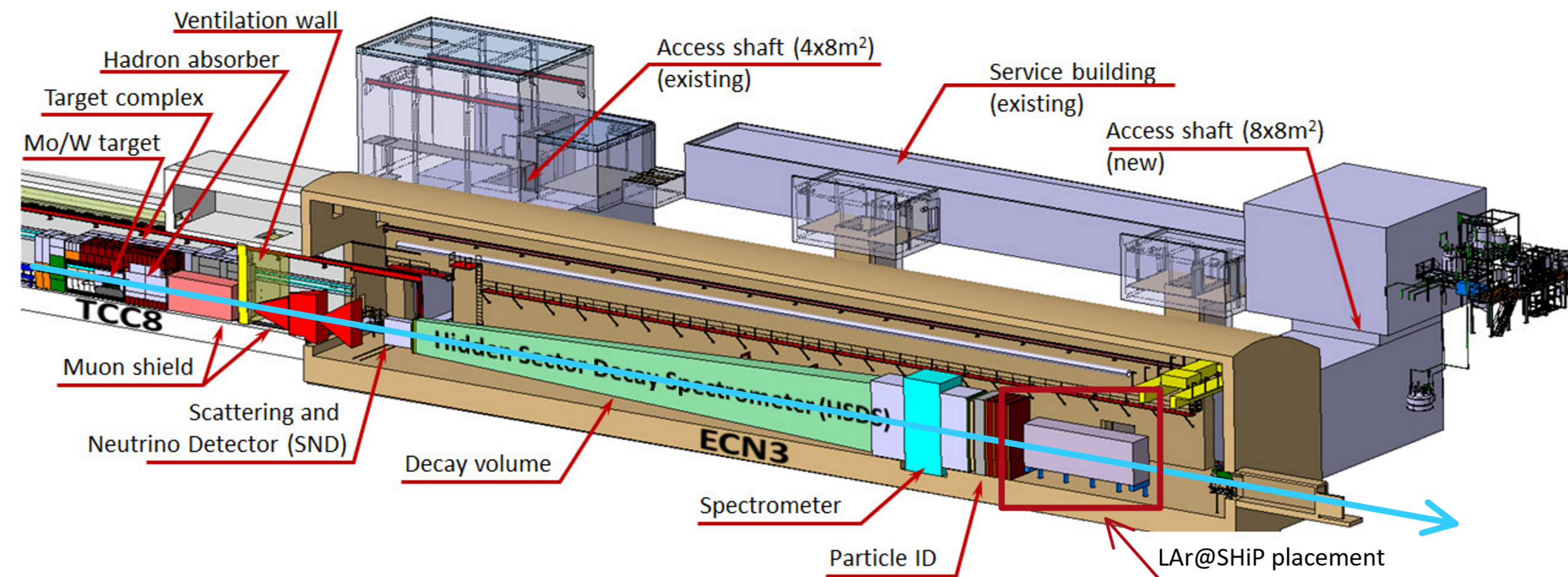
Expected sensitivities assuming signal only (no background)



2 years of ICARUS NuMI (2026). 4.5 years of DUNE Phase II (2040). 15 years of SHiP (2045).



# The SHiP experiment

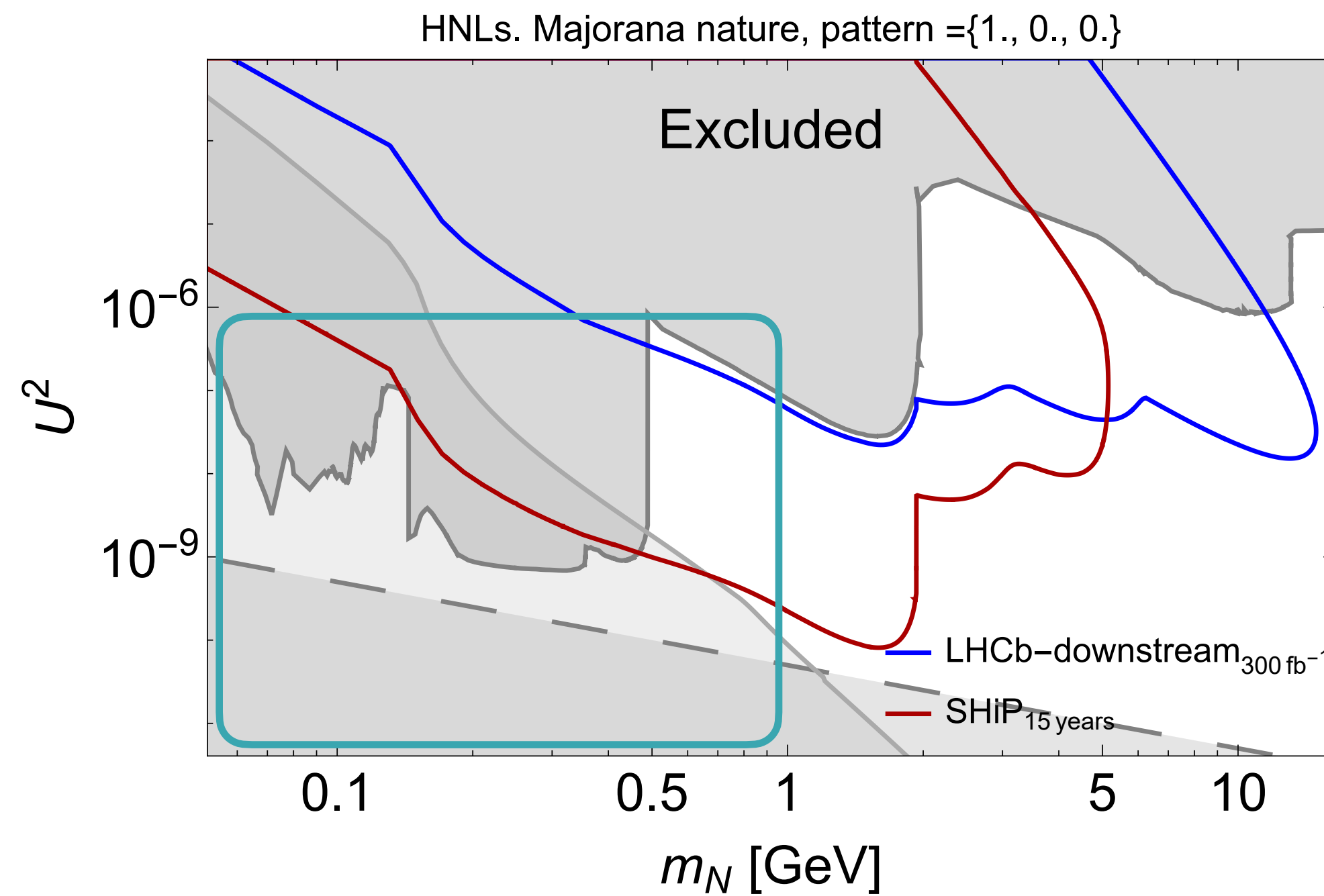
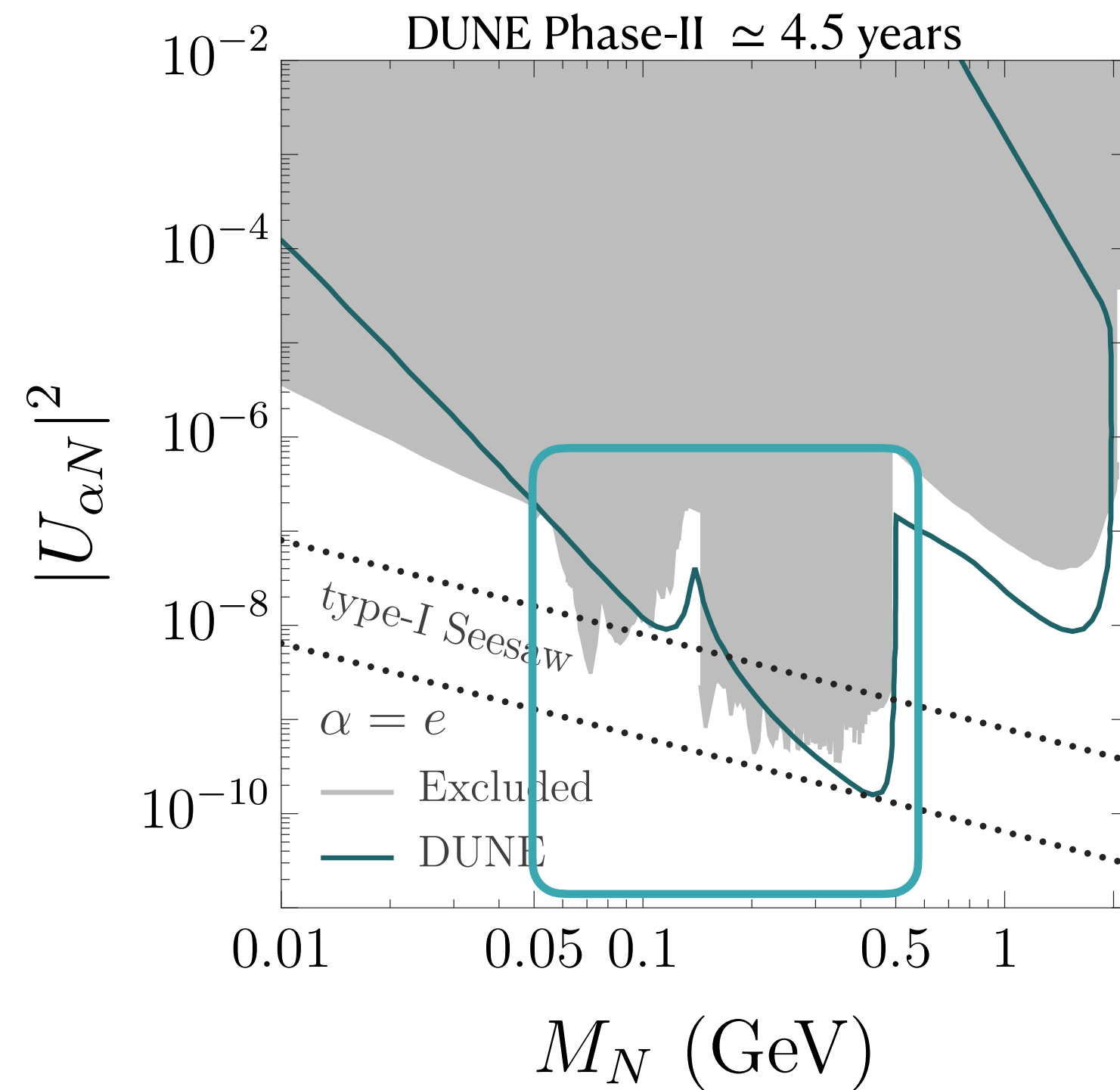


## SHiP experiment highlights:

- Proton energy of 400 GeV
- Expected running time  $\simeq 2030 - 2045$
- Large beam intensity  $4 \cdot 10^{19}$  PoT/year  $\Rightarrow$  an expected total of  $\simeq 6 \cdot 10^{20}$  PoT
- “Background-free” experiment for many scenarios
- Fully instrumented

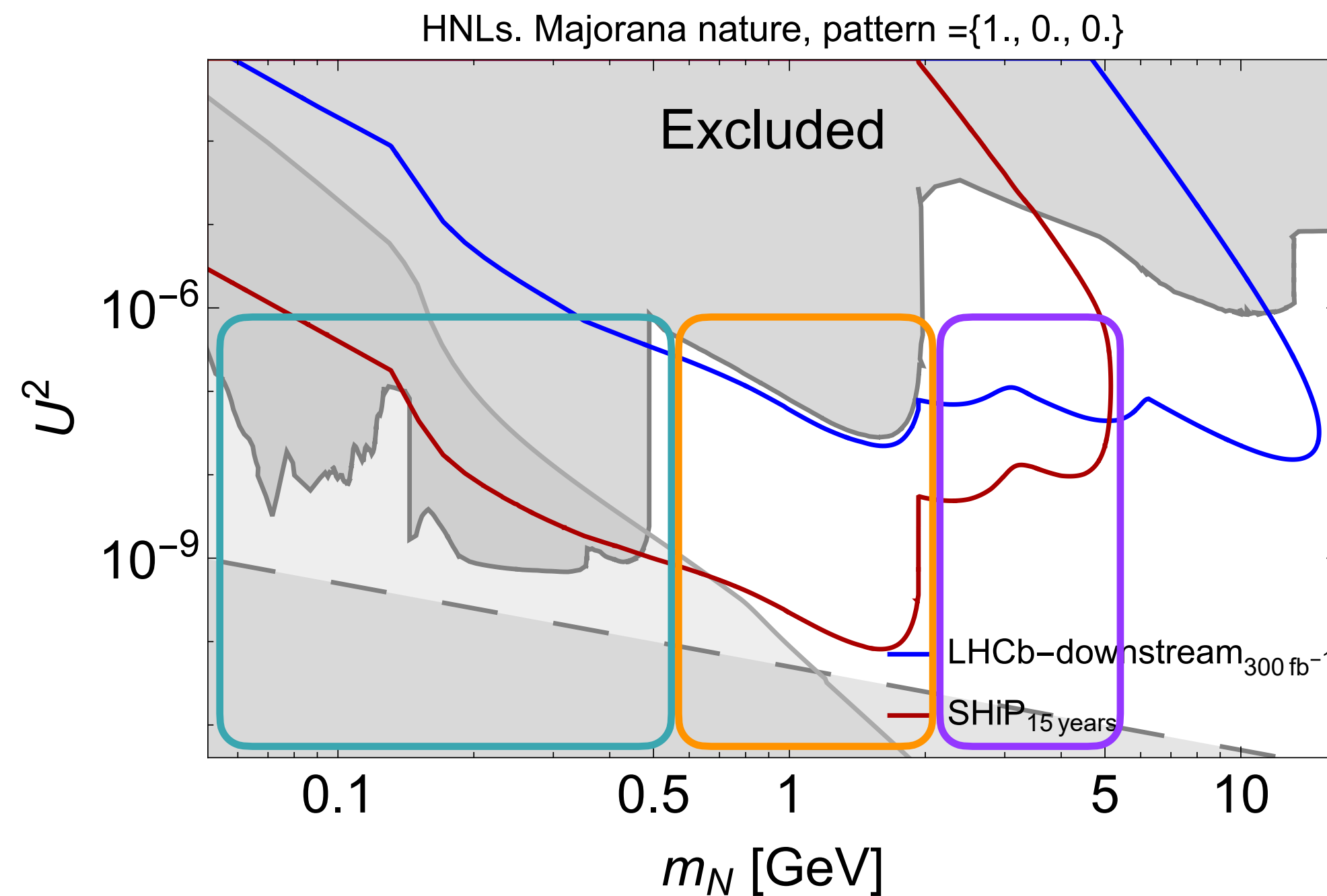
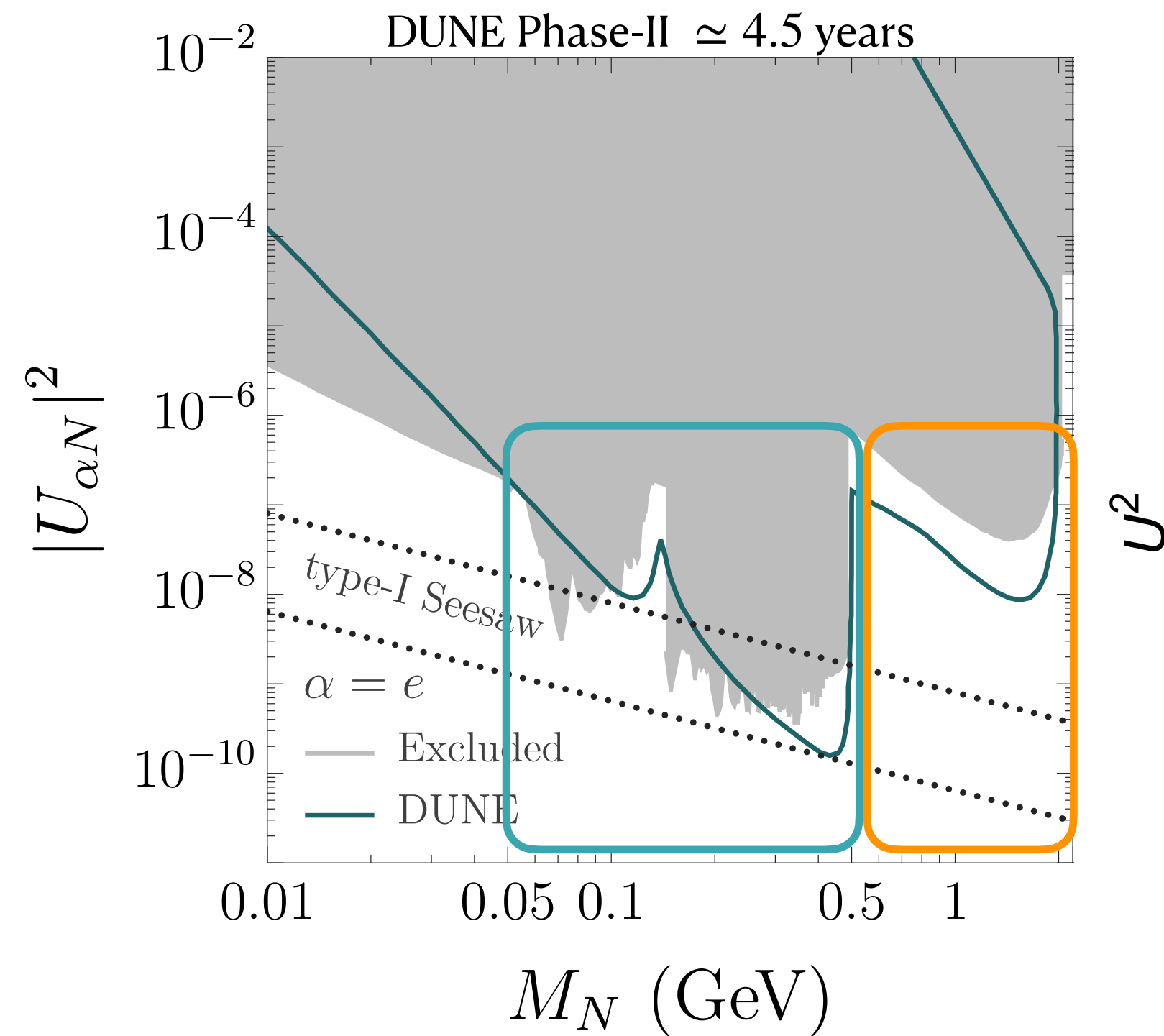
designed for  
LLP searches

# HNLs searches: DUNE ND-GAr vs SHiP



Better sensitivity of DUNE ND-GAr in the  $\pi$  and  $K$  regions

# HNLs searches: DUNE ND-GAr vs SHiP

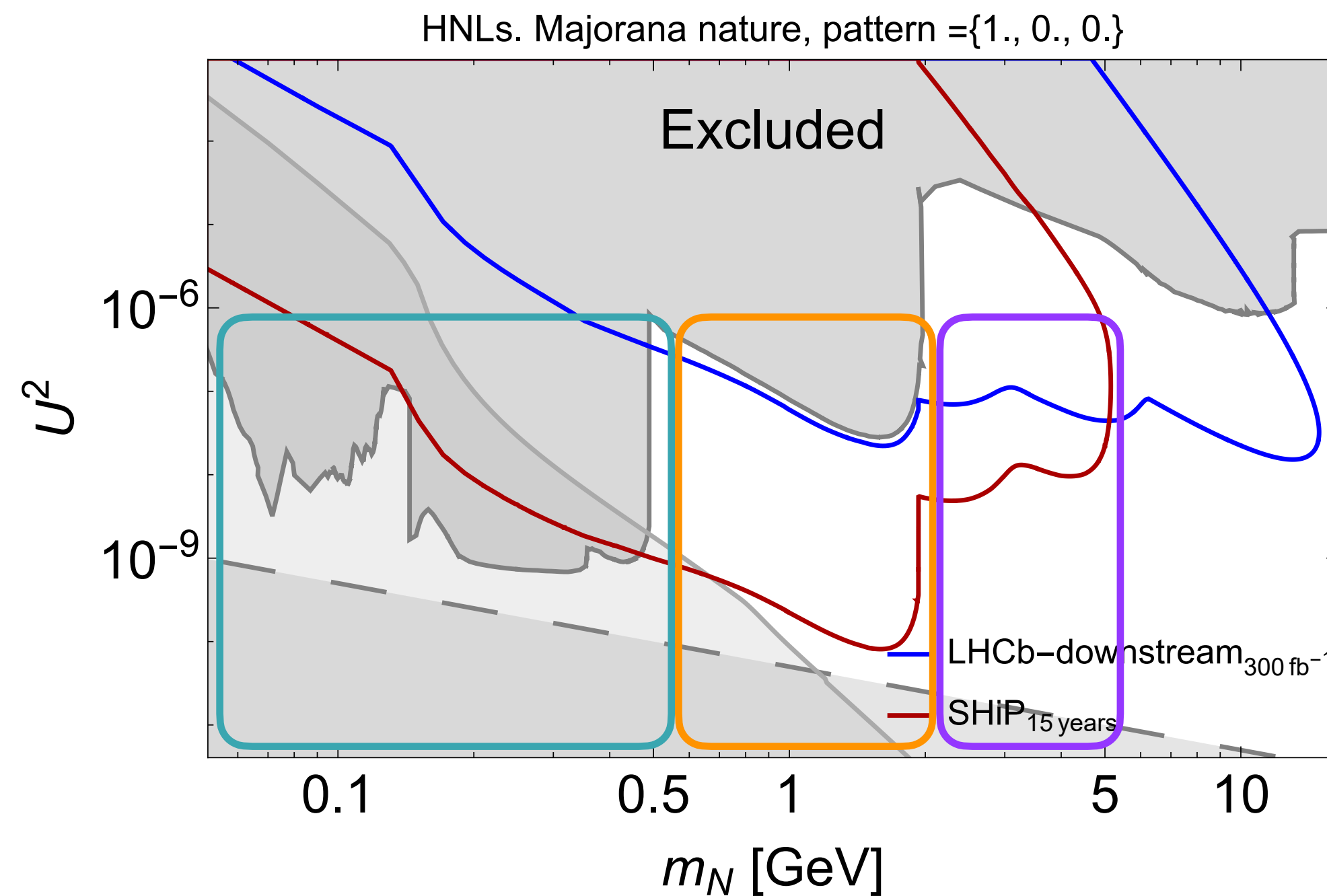
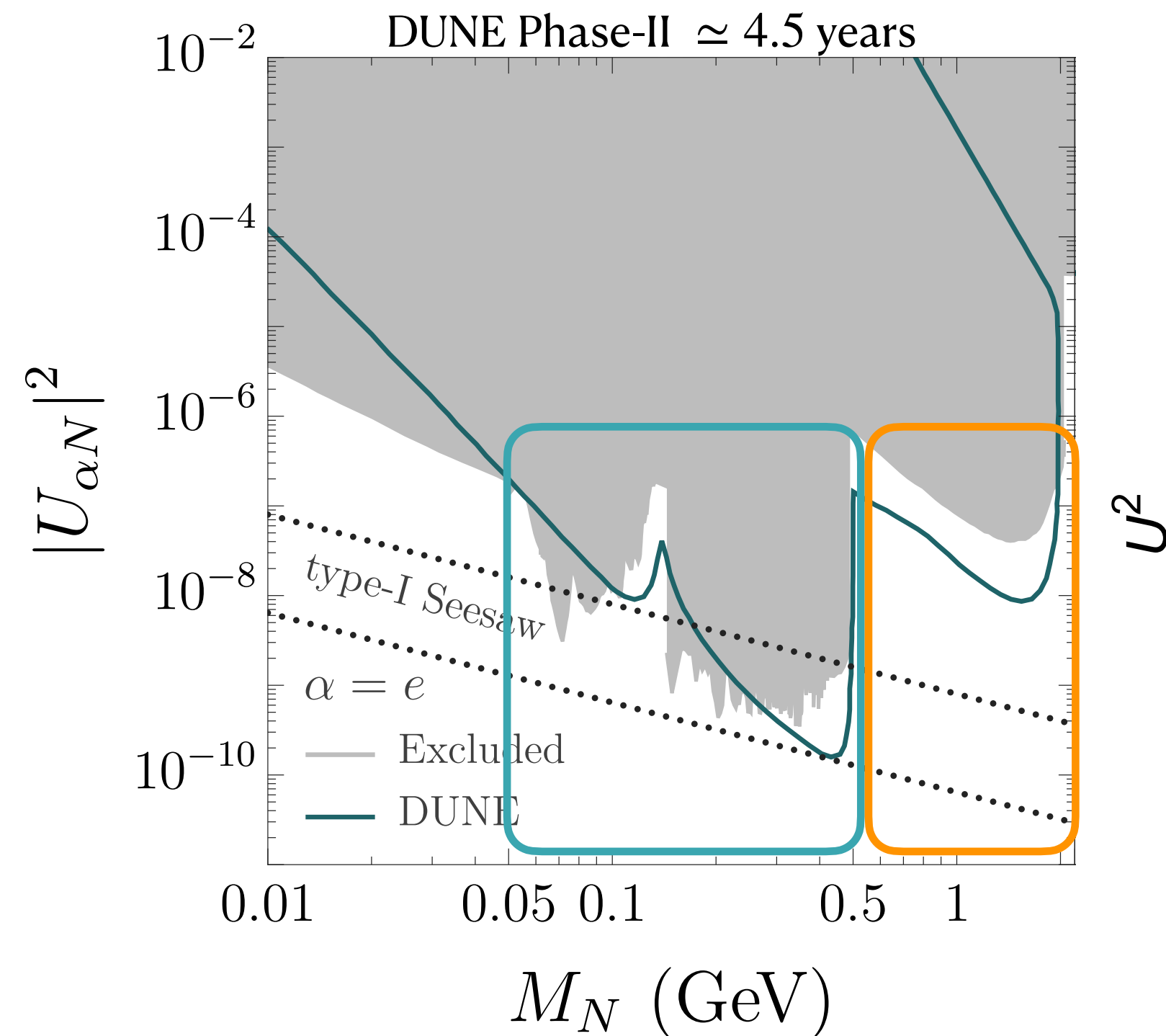


Better sensitivity of DUNE ND-GAr in the  $\pi$  and  $K$  regions

Better sensitivity of SHiP in the  $D$  and  $B$  regions



# HNLs searches: DUNE ND-GAr vs SHiP



Better sensitivity of DUNE ND-GAr in the  $\pi$  and  $K$  regions

Better sensitivity of SHiP in the  $D$  and  $B$  regions

Similar results for the  $|U_{\mu N}|^2$ . Better sensitivity of SHiP in the  $|U_{\tau N}|^2$  since it comes from  $D$  decays.