## Mesogenesis

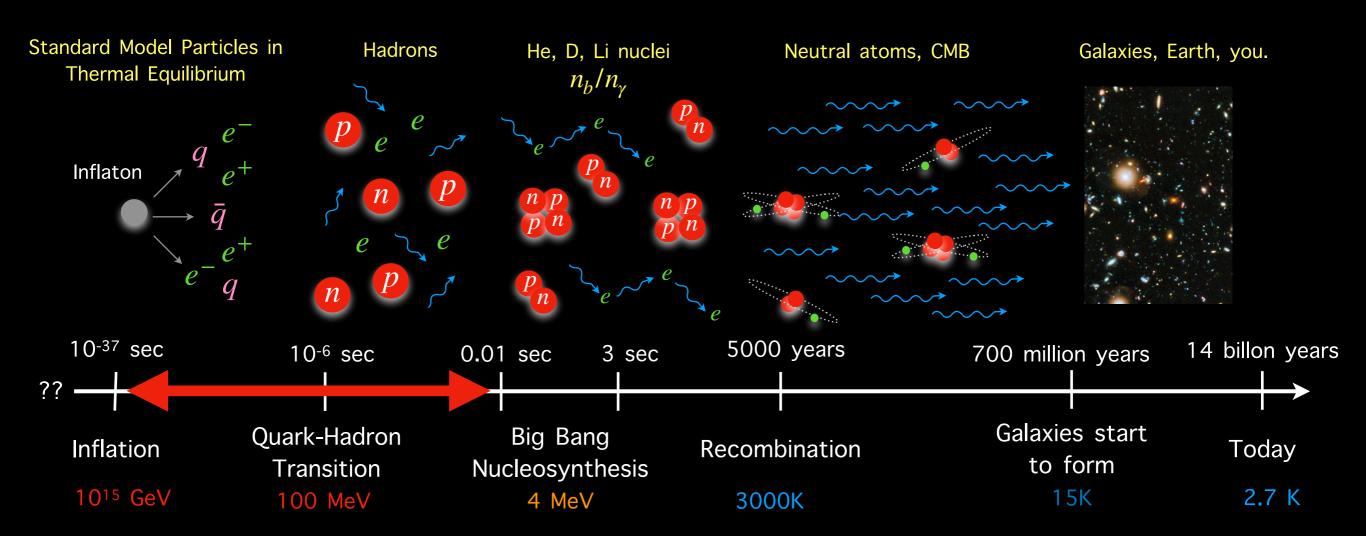
Gilly Elor

Mainz Institute for Theoretical Physics, JGU

IFT Workshop, Madrid

October 4 2022

## Baryogenesis

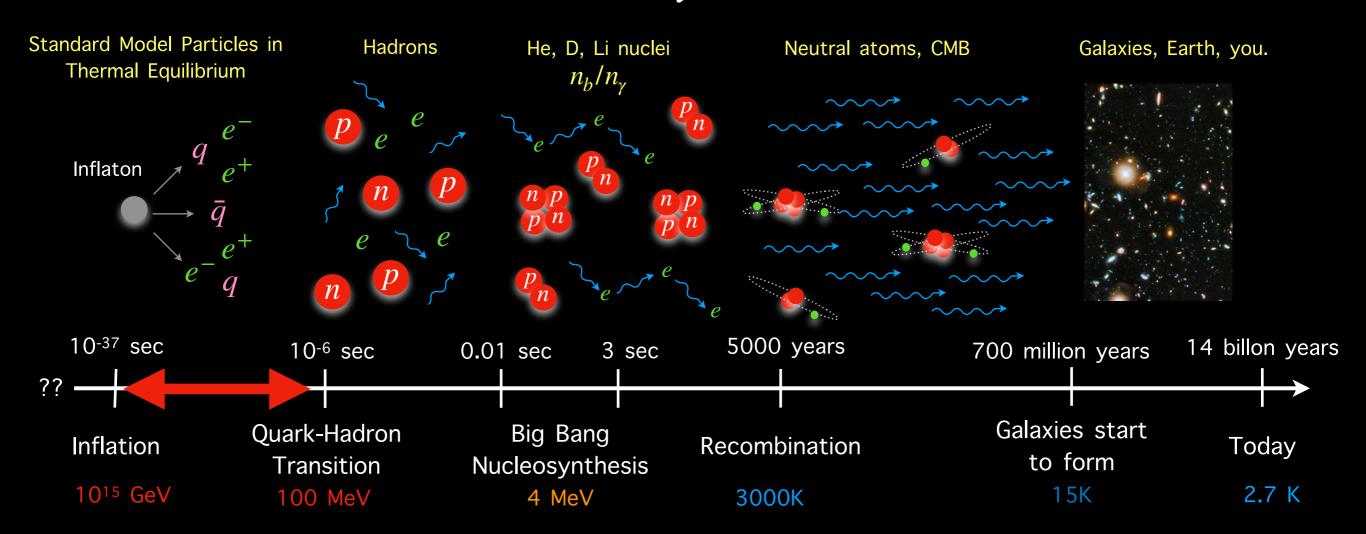


What mechanism generated the initial asymmetry? Observed to be (BBN, CMB):

$$Y_B^{obs} \equiv \frac{n_{\mathcal{B}} - n_{\bar{\mathcal{B}}}}{s} \sim 8 \times 10^{-11}$$

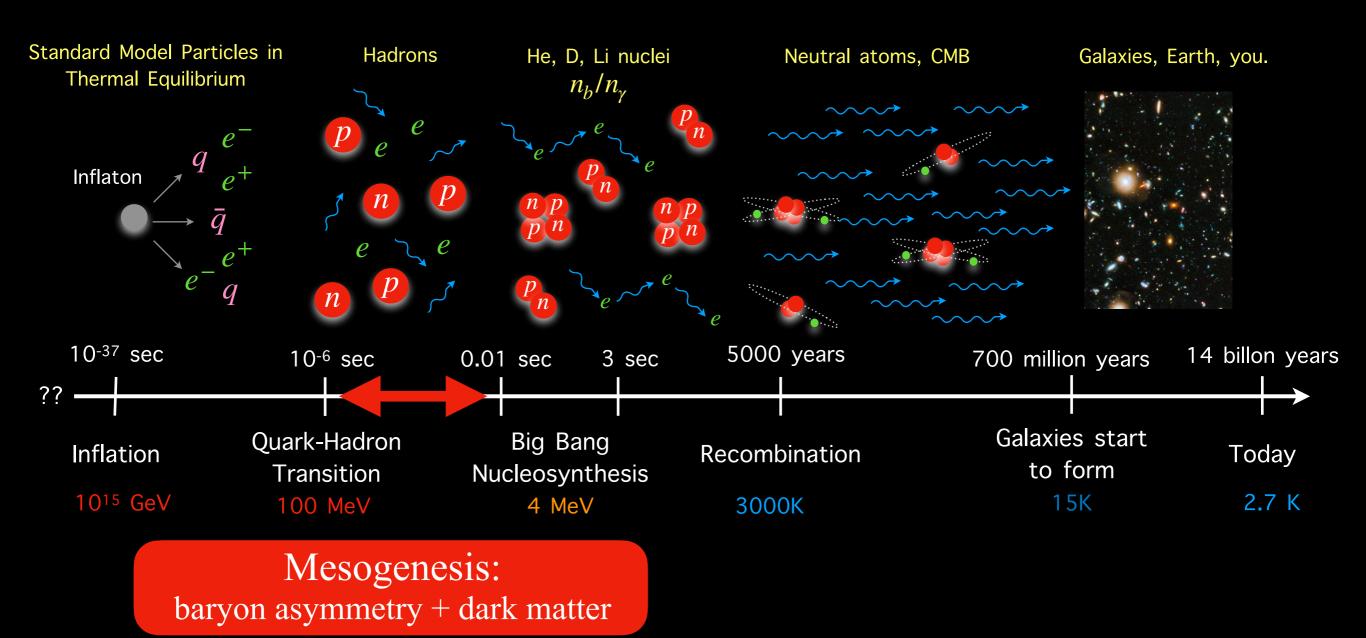
#### Traditional Baryogenesis Mechanisms

#### are traditionally hard to test



- GUT baryogenesis
- Electroweak baryogenesis
- Leptogenesis

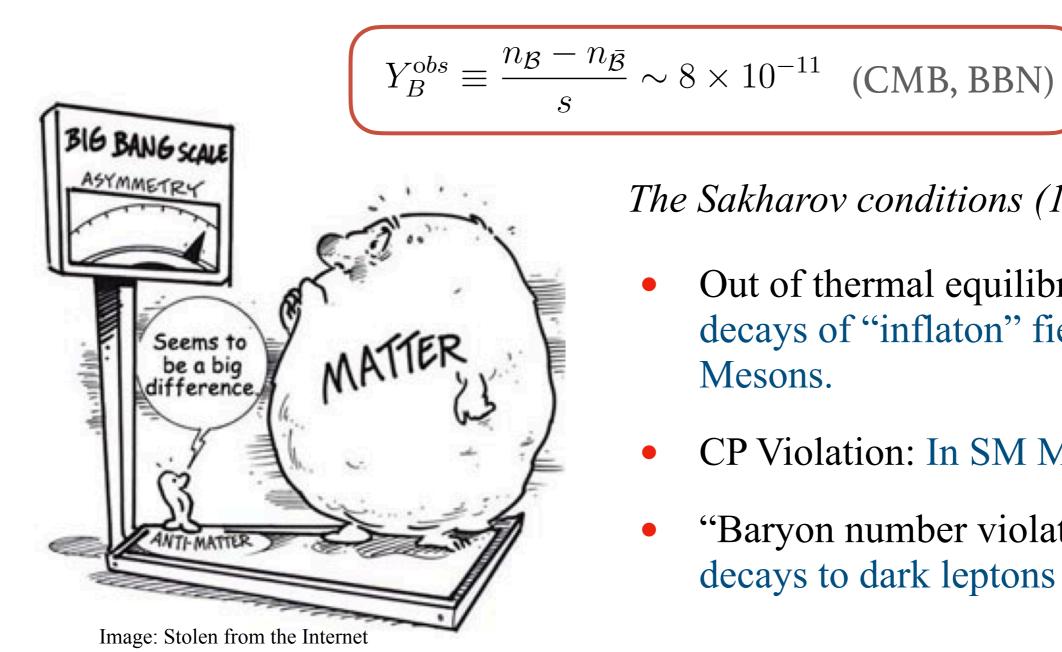
## Making the Universe at 20MeV



- Controlled by experimental observables. Signals!
- Theoretically appealing e.g. Relaxion and Nnaturalness require low scale baryogenesis.

## Mesogenesis

#### How to generate a matter/antimatter asymmetry



The Sakharov conditions (1967):

- Out of thermal equilibrium: Late decays of "inflaton" field to SM Mesons.
- CP Violation: In SM Meson systems.
- "Baryon number violation": SM Meson decays to dark leptons or baryons.

# The Many Flavors of Mesogenesis

Mechanism	CPV	Dark Sector	Observables	Relevant Experiments
$B^0$ Mesogenesis	$B_s^0 \& B_d^0$	Dark baryons	$A_{sl}^{s,d}$	LHCb
	oscillations		$Br(B \to \mathcal{B} + X)$	B Factories, LHCb
			$A^D_{CP}$	B Factories, LHCb
$D^+$ Mesogenesis	$D^{\pm}$ decays	Dark leptons	${\rm Br}_{D^+}$	B Factories, LHCb
		and/or baryons	$\operatorname{Br}(\mathcal{M}^+ \to \ell^+ + X)$	peak searches e.g. PSI, PIENU
			$A_{CP}^B$	B Factories, LHCb
$B^+$ Mesogenesis	$B^{\pm}$ decays	Dark leptons	${\rm Br}_{B^+}$	B Factories, LHCb
		and/or baryons	$Br(\mathcal{M}^+ \to \ell^+ + X)$	peak searches e.g. PSI, PIENU
			$A_{CP}^{B_c}$	LHCb, FCC
$B_c^+$ Mesogenesis	$B_c^{\pm}$ decays	Dark baryons	$\operatorname{Br}_{B_c^+}$	LHCb, FCC
			$\operatorname{Br}_{B^+ \to \mathcal{B}^+ + X}$	B Factories, LHCb

GE, M. Escudero, A. E. Nelson, PRD, [1810.00880]

G. Alonso-Alvarez, GE, A. E. Nelson, H. Xiao, JHEP, [1907.10612]

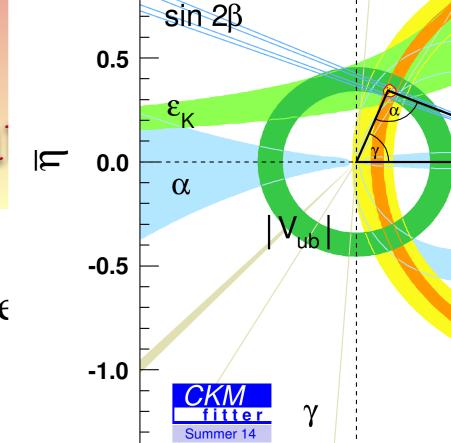
GE, R. McGehee, PRD [2011.06115]

G. Alonso-Alvarez, GE. M. Escudero, PRD, [2101.02706]

F. Elahi, GE, R. McGehee, [2109.09751]

G. Alonso-Alvarez, GE, M. Escudero, B. Fornal, B. Grinstein, J.M. Camalich [arXiv:2111.12712]

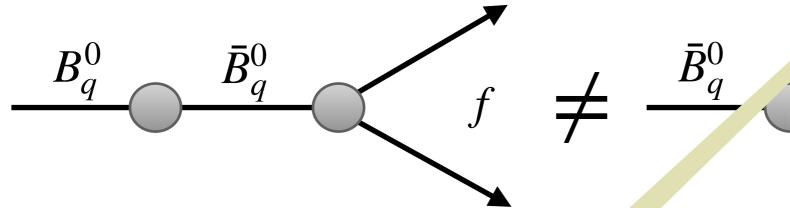
## CP Violation in B Meson Oscilla



-0.5

0.0

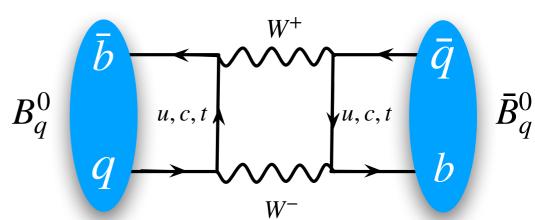
B meson/anti-meson mixing has sizable



#### Experimental Observable:

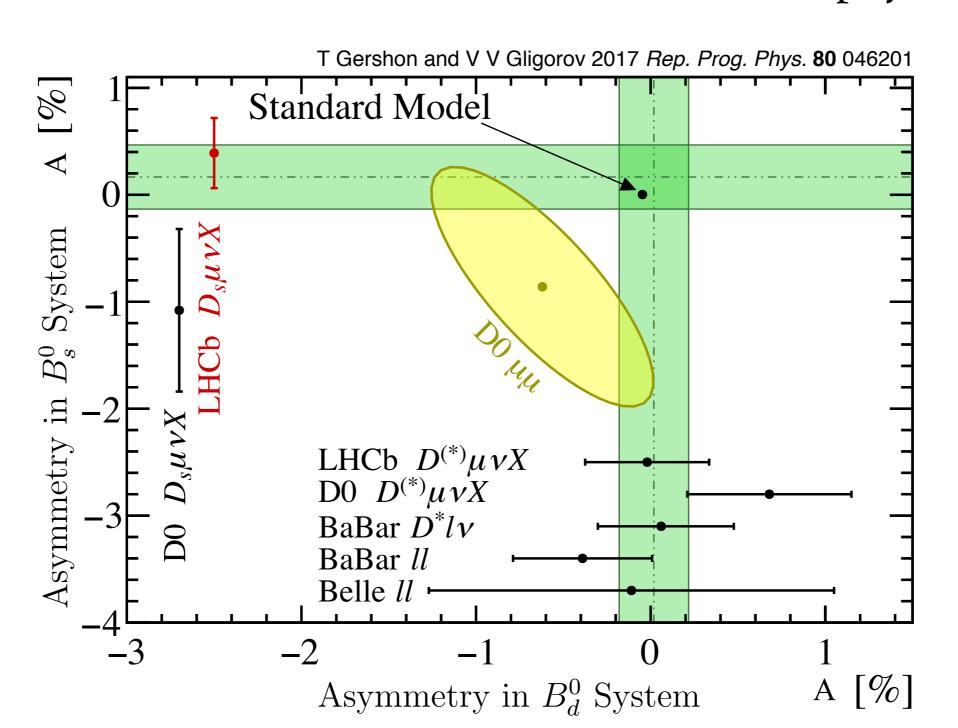
$$A_{\mathrm{SL}}^{q} = \frac{\Gamma\left(\bar{B}_{q}^{0} \to B_{q}^{0} \to f\right) - \Gamma\left(B_{q}^{0} \to \bar{f}\right)}{\Gamma\left(\bar{B}_{q}^{0} \to B_{q}^{0} \to f\right) + \Gamma\left(B^{r} \to \bar{B}_{q}^{0} \to \bar{f}\right)}$$

#### e.g. in Standard Model



## CP Violation in B Meson Oscillations

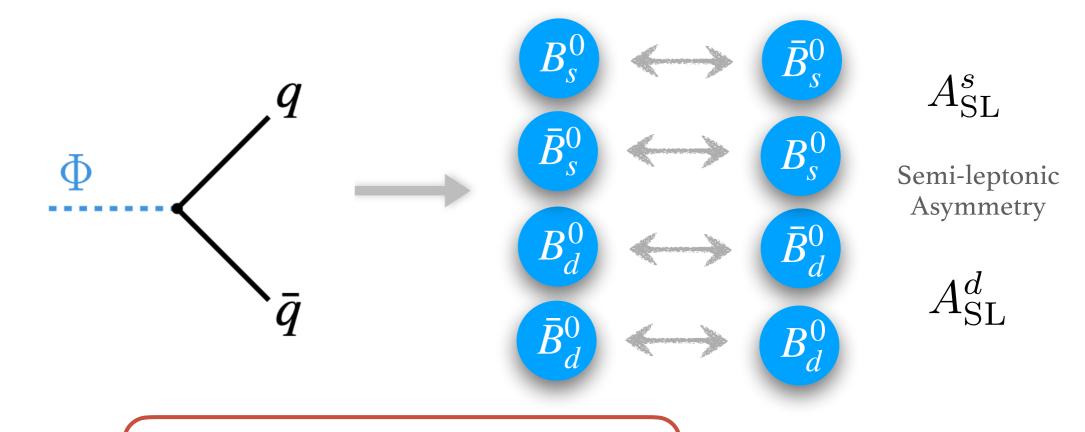
#### Can accommodate contributions from new physics



## Sakharov I & II: Out of thermal equilibrium and CPV

Late decay of an "inflaton-like" field

Decays at:  $\Gamma_{\Phi} = H(T_R)$  to quarks  $m_{\Phi} \in [5 \, \mathrm{GeV}, 100 \, \mathrm{GeV}]$ 



 $3.5\,\mathrm{MeV} \lesssim T_\mathrm{R} \lesssim 100\,\mathrm{MeV}$ 

Before BBN

After QCD phase transition

#### Sakharov III. B Violation?

Need a way to change baryon number



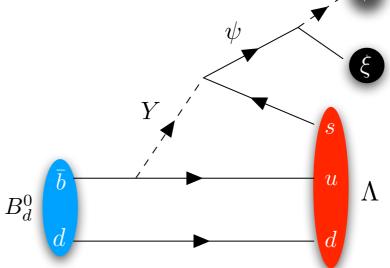
Hide baryon number in a dark sector rather than violate it



## An Explic Dark Sector Baryon

Field	Spin	$Q_{EM}$	Baryon no.	$\mathbb{Z}_2$	Hige
Φ	0	0	0	+1	$11-100\mathrm{GeV}$
Y	0	-1/3	-2/3	+1	$\mathcal{O}(\mathrm{TeV})$
$\psi$	1/2	0	-1	+1	$\mathcal{O}(\mathrm{GeV})$
ξ	1/2	0	0	-1	$\mathcal{O}(\mathrm{GeV})$
$\phi$	0	0	$\begin{bmatrix} & B \end{bmatrix}$	<b>VIES</b> -1	$Son_{\mathcal{O}(GeV)}$
			Ç	$J_h =$	<del>-</del> U

baryon number in a dark sector



Kinematics:  $m_{\psi} < m_B - m_{\rm Baryon} < 4.3 \, {\rm GeV}$ 

Proton stability:  $m_{\psi} > m_{p} - m_{e} \simeq 937.8 \,\mathrm{MeV}$ 

Equal and opposite dark and visible baryon as yn mætries generated.

$$Y_{\mathcal{B}} - Y_{\bar{\mathcal{B}}} = -\left(Y_{\psi} - Y_{\bar{\psi}}\right)$$

## An Explicit Mos

Field	Spin	$Q_{EM}$	Baryon no.	$\mathbb{Z}_2$	Mass
Y	0	-1/3	-2/3	+1	$\mathcal{O}(\mathrm{TeV})$
$\psi_{\mathcal{B}}$	1/2	0	-1	+1	$\mathcal{O}(\mathrm{GeV})$

SUSY Squark

Kinematics, forbid

proton decay

Allowed by all the symmetries:

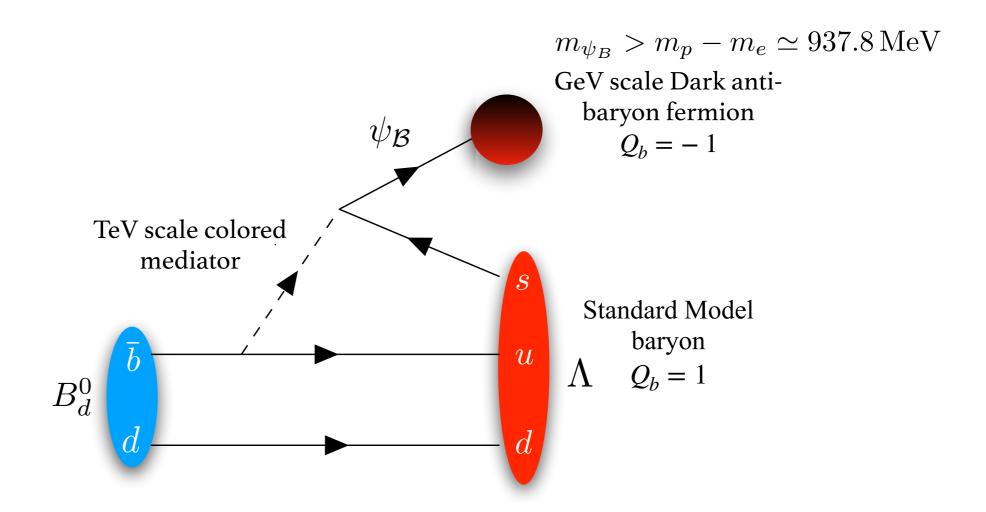
$$\mathcal{L}_{-1/3} = -\sum_{i,j} y_{u_i d_j} Y^* \bar{u}_{iR} d_{jR}^c - \sum_k y_{\psi d_k} Y d_{kR}^c \bar{\psi}_{\mathcal{B}} + \text{h.c.}$$

Effective four fermion operator at MeV scales:

$$\mathcal{H}_{eff} = \frac{\kappa}{m_Y^2} b \, u \, s \, \psi_{\mathcal{B}}$$

This interaction *does not* change baryon number

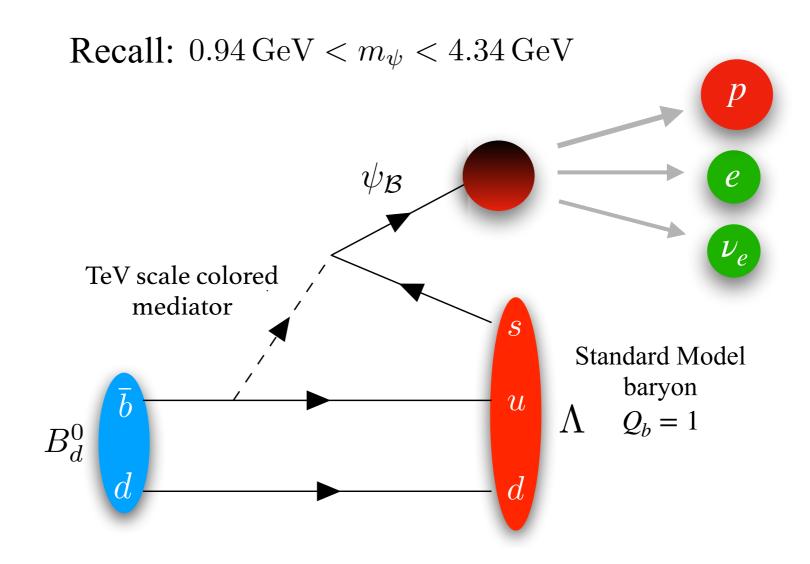
### New decay of the B Meson



Equal and opposite dark and visible baryon asymmetries generated.

$$Y_{\mathcal{B}} - Y_{\bar{\mathcal{B}}} = -\left(Y_{\psi} - Y_{\bar{\psi}}\right)$$

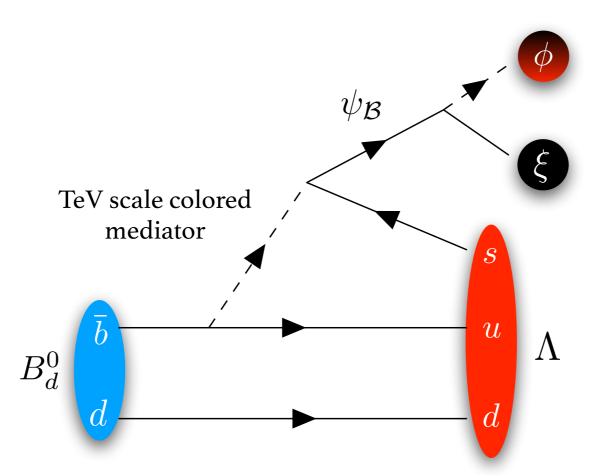
#### Dark Matter?



New dark baryon is unstable and will decay to baryonic matter, washing out the asymmetry in the process. It cannot be the dark matter.

#### Dark Matter

Dark fermion must quickly decay within the dark sector.



Dark scalar anti-baryon

$$Q_b = -1$$

Dark Fermion

$$Q_b = 0$$

DM stability/asymmetry preserved if:

$$m_{\phi} < m_p + m_e + m_{\xi}$$

Generated asymmetry:

$$(Y_{\mathcal{B}} - Y_{\bar{\mathcal{B}}} = -(Y_{\phi} - Y_{\phi^*}))$$

## Boltzmann Equations

$$\frac{dn_{\Phi}}{dt} + 3Hn_{\Phi} = -\Gamma_{\Phi}n_{\Phi}$$

"Inflaton": 
$$\frac{dn_{\Phi}}{dt} + 3Hn_{\Phi} = -\Gamma_{\Phi}n_{\Phi}$$
 Hubble: 
$$\frac{d\rho_{\rm rad}}{dt} + 4H\rho_{\rm rad} = +\Gamma_{\Phi}m_{\Phi}n_{\Phi}$$
 
$$H^2 = \frac{8\pi}{3M_{\rm Pl}^2}\left(\rho_{\rm rad} + m_{\Phi}n_{\Phi}\right)$$

$$H^2 = \frac{8\pi}{3M_{\rm Pl}^2} \left( \rho_{\rm rad} + m_{\Phi} n_{\Phi} \right)$$

Symmetric component of the dark scalar baryon

$$\frac{dn_{\phi+\phi^*}}{dt} + 3H n_{\phi+\phi^*} = 2\Gamma_{\Phi}^B n_{\Phi} - 2\langle \sigma v \rangle_{\phi} \left( n_{\phi+\phi^*}^2 - n_{\text{eq},\phi+\phi^*}^2 \right)$$

Anti-symmetric dark sector baryon makes up the baryon asymmetry

$$\frac{dn_{\phi-\phi^*}}{dt} + 3Hn_{\phi-\phi^*} = 2\Gamma_{\Phi}^B \sum_{q} \operatorname{Br}\left(\overline{b} \to B_q^0\right) A_{\operatorname{SL}}^q f_{\operatorname{deco}}^q n_{\Phi}$$

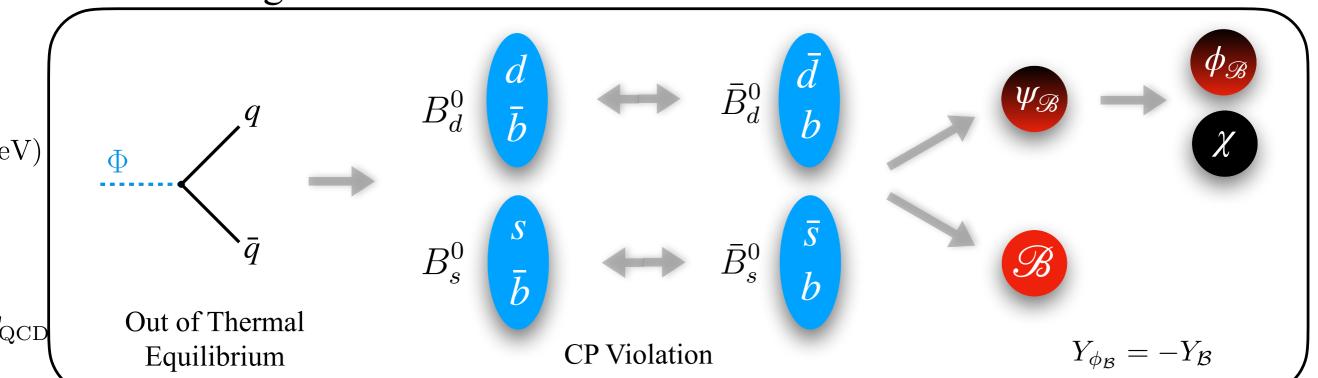
(20)Example Benchmark Point 1S encl neck work ve constan in table II: GeV, 0.3}. The left  $panel_{\infty}^{(2)}$  corresponds the  $m_{\xi} = 1.8 \,\mathrm{GeV}$ = 1-51GGV:11WeVtake hour 114 BOTai hánge in bel e Bascilla Br  $(B \to \psi \mathcal{B} \mathcal{M}) = 5.6 \times 10^{-3}$  is the standard with the standard particles, as we take  $m_{\epsilon} \neq 1$  GeV,  $5.6 \times 10^{-3}$ , 3.3 GeV, 0.3 and the properties of the particles, as we take  $m_{\epsilon} \neq 1$  GeV reported to 1.5 GeV  $-\frac{10^{-4}}{m_{\xi}} = Y_{10}$  $m_{\phi} = 1$  $m_{\phi} = 1 3 \, \mathrm{Ge}$ contributions to the leptonic asymmetry to be positive  $A_{\ell\ell} = 10^{-4}$  at  $T \sim 15$  MeV corresponds to decommende effects speciments MATERIAL PAR be underste remark to a mainly of stark party on  $\phi + \phi^*$ , with  $m_{\phi} = 1.3$ at∲reproduc bility 10 Faultions n(5) Ann (5) Ann (7). -We × 10<sup>-4</sup> - the dip in the asymmetry 24 in be under this case to correspond to the SM prediction. 1Both benchmark 2011 to reprod at  $T \sim 15 \,\mathrm{Me}$ **fy**]\* The latest of the simplifying the phenomena of the second The parameter space of our mixed highway with the leptonic way wi  $m_{\epsilon} < p_{54} \kappa_{\rm III}/2$ one thing the formula  $m_{\phi} < n_{35}$  of  $m_$  $t_{\rm III}/2$  $\Omega_{\rm DM} h^2 (29) .12$ news physisis, convibutions, marizes the pa-.12 Coherent Bs  $\Omega_{\rm DM}h^2 = 0.12$ Barvon states w can constrained by kinematics, proteined usual = \$2 EAL 1

In visible values. For mestance, and the constrained by kinematics, proteined as a constrained by kinematics, proteined by kinematics, proteined as a constrained by kinematics, proteined by kinematics, proteined as a constrained by kinematics, proteined by kinematics,  $Y_B = 8.7 \times 10^{-11}$ 45 y Buryon when 10  $T_{L_{as}} \lesssim 20 \, \mathrm{MeV} \ T_{B_d} \lesssim 10 \, \mathrm{MeV} \ 30$ 10 EAL III Coherent Bd oscillations start to  $T_{\gamma}$  / MeV

deplete the as Figure 120 Wheel  $T_{B_d} \lesssim 10 \,\mathrm{MeV}$ 

## Neutral B Mesogenesis

#### B<sup>0</sup> Mesogenesis



$$T \lesssim 20\,\mathrm{MeV}$$

Asymmetry is related to observables:

$$\begin{array}{l}
B^{0} = \sum_{P} \Delta m_{R}^{0} = 8.7 \times 10^{-11} \frac{\text{Br} (B \to \psi \mathcal{B} \mathcal{M})}{10^{-2}} \sum_{q=s,d} \alpha_{q} \frac{A_{\text{SL}}^{q}}{10^{-4}} \\
\frac{T}{20 \,\text{MeV}} = \sum_{p} \left( \frac{\langle r_{B_{0}}^{2} \rangle}{0.187} \right)^{2} & \text{Br} (B \to \psi \mathcal{B} \mathcal{M}) \\
\frac{T}{20 \,\text{MeV}} = \sum_{p} \alpha_{q} \frac{A_{\text{SL}}^{q}}{10^{-4}} & \text{Br} (B \to \psi \mathcal{B} \mathcal{M}) \\
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\frac{T}{20 \,\text{MeV}} = \sum_{p} \alpha_{q} \frac{A_{\text{S$$

succesful B Mesogenesis requires:

$$A_{\rm SL}^{s,d} \times \operatorname{Br}\left(B^0 \to \psi \,\mathcal{B} \,\mathcal{M}\right) > 10^{-6}$$

### Signals of Neutral B-Mesogenesis

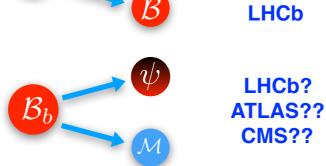
For successful baryogenesis:  $A_{\rm SL}^{s,d} \times \text{Br}\left(B^0 \to \psi \,\mathcal{B} \,\mathcal{M}\right) > 10^{-6}$ 

Collider Signals of Baryogenesis and Dark Matter from B Mesons (*B-Mesogenesis*)

#### **Direct Signals**

## Semileptonic asymmetry: $A_{\rm SL}^q>10^{-5}$ Belle II LHCb ATLAS CMS Rew B meson decay: Belle II Belle II

New b-Baryon decay:



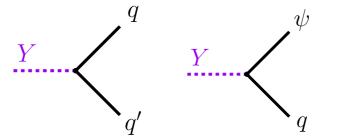
#### **Indirect Signals**

B<sup>0</sup> meson CPV and oscillation observables:

$$\phi_{12}^{d,s}$$
  $\Delta M_{d,s}$   $\Delta \Gamma_{d,s}$ 

Belle II ATLAS CMS

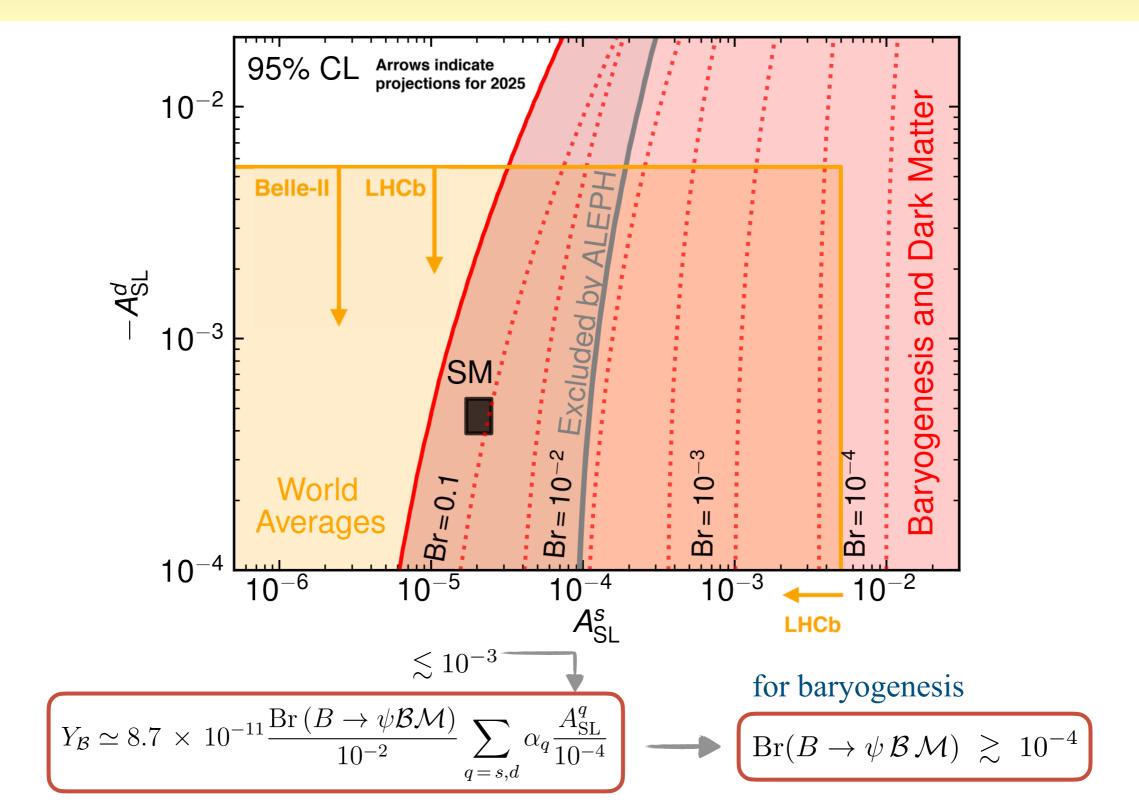
New TeV-scale color-triplet scalar, Y



ATLAS CMS

**LHCb** 

## The Semi-Leptonic Asymmetry



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#### New b-Hadron Decays

$$\mathcal{L}_{-1/3} = -\sum_{i,j} y_{u_i d_j} Y^* \bar{u}_{iR} d^c_{jR} - \sum_k y_{\psi d_k} Y d^c_{kR} \bar{\psi} + \text{h.c.}$$

#### Flavorful variations:

Operator/Decay	Initial State	Final state
$egin{aligned} \mathcal{O} &= \psi  b  u  d \ ar{b} & ightarrow \psi  u  d \end{aligned}$	$B_d$ $B_s$ $B^+$	$\psi + n (udd)$ $\psi + \Lambda (uds)$ $\psi + p (duu)$
	$\Lambda_b$	$\bar{\psi} + \pi^0$
$\mathcal{O} = A \cdot b \cdot a \cdot a$	$B_d$	$\psi + \Lambda (usd)$ $\psi + \Xi^0 (uss)$
$ \mathcal{O} = \psi  b  u  s \\ \bar{b} \to \psi  u  s $	$B_s$ $B^+$	$\psi + \Sigma^+ (uus)$
	$\Lambda_b$	$\bar{\psi} + K^0$
$\mathcal{O} = \psi  b  c  d$	$egin{array}{c} B_d \ B_s \end{array}$	$\begin{vmatrix} \psi + \Lambda_c + \pi^- (cdd) \\ \psi + \Xi_c^0 (cds) \end{vmatrix}$
$\bar{b}  o \psi  c  d$	$B^+$	$\psi + \Lambda_c (dcu)$
	$\Lambda_b$	$\overline{\psi} + \overline{D}^0$
$\mathcal{O} = \psibcs$	$egin{array}{c} B_d \ B_s \end{array}$	$\psi + \Xi_c^0 (csd)$ $\psi + \Omega_c (css)$
$\bar{b} \rightarrow \psi  c  s$	$B^+$	$\psi + \Xi_c^+ (csu)$
	$\Lambda_b$	$  \bar{\psi} + D^- + K^+  $

Directly related to baryon asymmetry

Indirectly constrains *B*-Mesogenesis.
Charged track is an advantage for searches

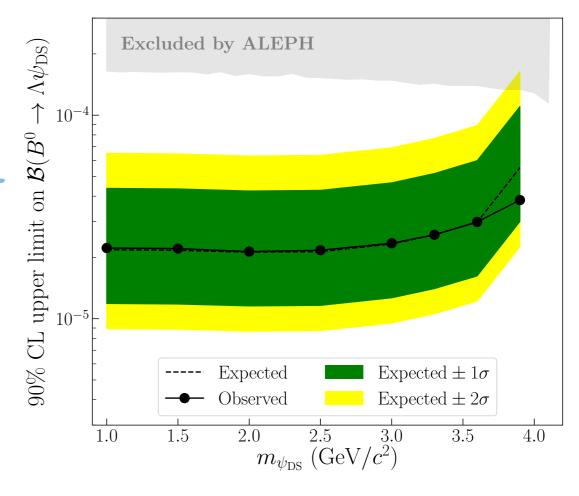
b-flavored baryon decays can yield indirect constraints.

#### New b-Hadron Decays

#### Targeted Searches at Colliders!

Operator/Decay	Initial State	Final state
	$B_d$	$\psi + n  (udd)$
$\mathcal{O} = \psi  b  u  d$	$B_s$	$\psi + \Lambda \left( uds \right)$
$\bar{b} \rightarrow \psi  u  d$	$B^+$	$\psi + p \left( duu \right)$
	$\Lambda_b$	$\bar{\psi} + \pi^0$
	$B_d$	$\psi + \Lambda \left( usd \right)$
$\mathcal{O} = \psi  b  u  s$	$B_s$	$\psi + \Xi^0 (uss)$
$\bar{b} \rightarrow \psi  u  s$	$B^+$	$\psi + \Sigma^{+} (uus)$
	$\Lambda_b$	$\bar{\psi} + K^0$
	$B_d$	$\psi + \Lambda_c + \pi^- \left( cdd \right)$
$\mathcal{O} = \psi  b  c  d$	$B_s$	$\psi + \Xi_c^0 \left( cds \right)$
$\overline{b} \rightarrow \psi  c  d$	$B^+$	$\psi + \Lambda_c (dcu)$
	$\Lambda_b$	$\bar{\psi} + \overline{D}^0$
	$B_d$	$\psi + \Xi_c^0 \left( csd \right)$
$\mathcal{O} = \psi  b  c  s$	$B_s$	$\psi + \Omega_c \left( css \right)$
$\overline{b} \rightarrow \psi  c  s$	$B^+$	$\psi + \Xi_c^+ (csu)$
	$\Lambda_b$	$\bar{\psi} + D^- + K^+$

#### Belle collaboration [arXiv:2110.14086]



Proposed searches for all modes at LHCb [2106.12870]

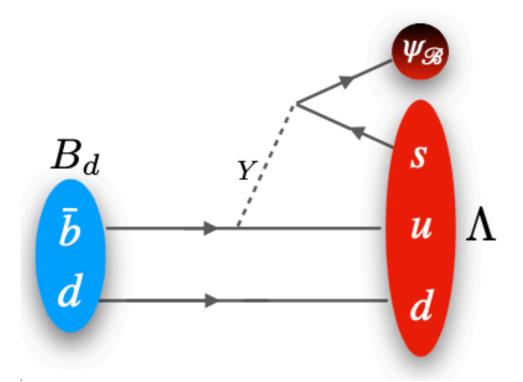
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#### New b-Hadron Decays

#### Form Factor Calculation from Light Cone QCD Sum Rules

[arXiv:22xx.xxxx] GE with Alfredo Guerrera

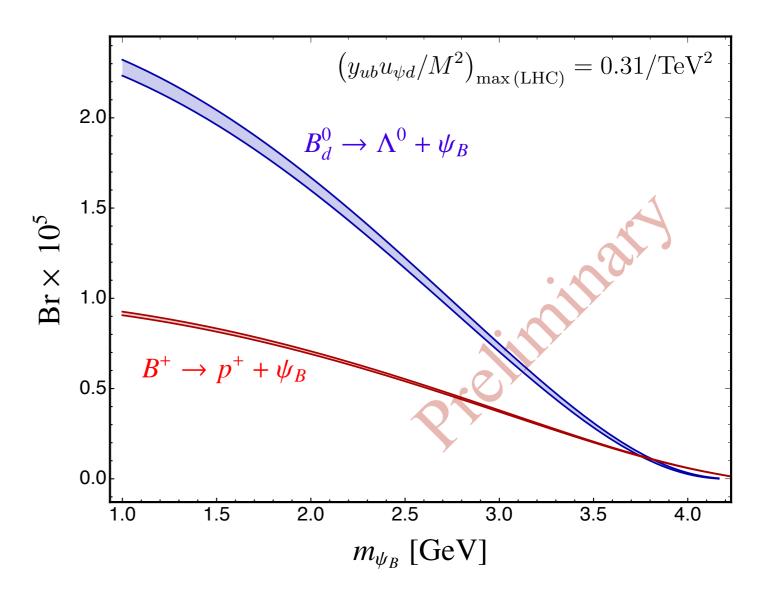
example: 
$$\mathcal{A}\left(B_s^0 \to \Lambda^0 \psi_B\right) = \frac{y_{ub} u_{\psi d}}{M^2} \left(b_R u_R\right) \langle \Lambda^0(P) | \mathcal{O}_{ub,d} | B_s^0(P+q) \rangle u_{\psi}^c(q)$$



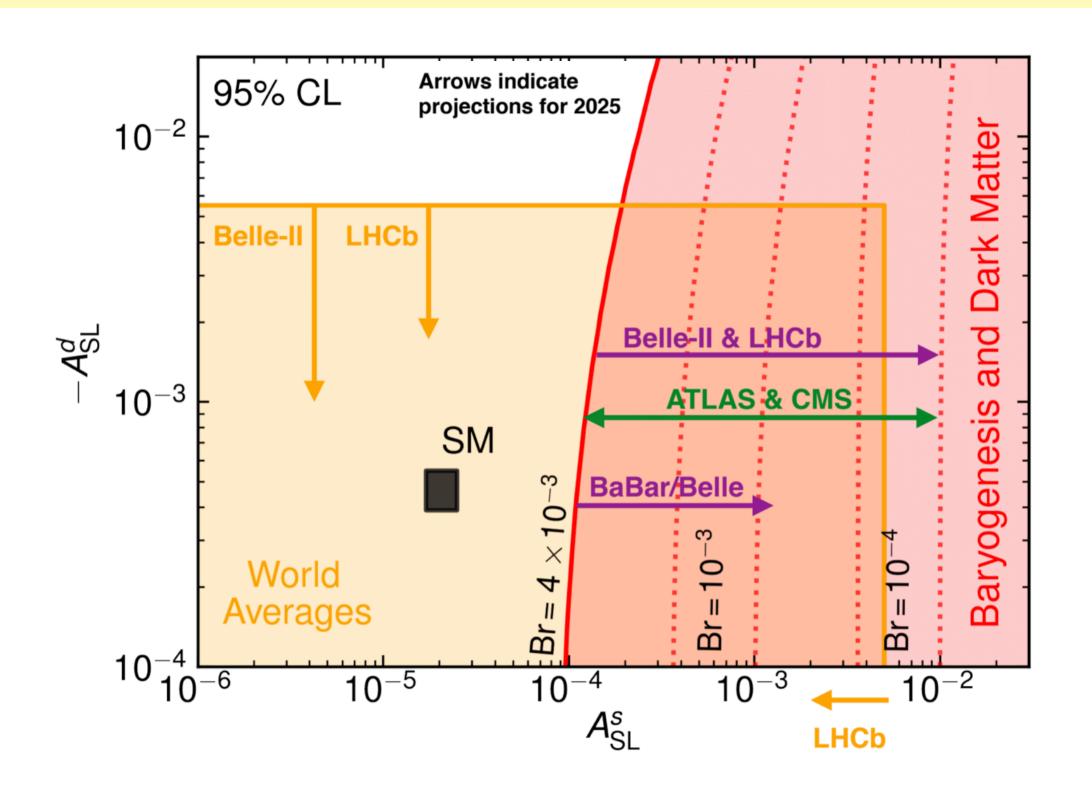
$$\mathcal{O}_{ub,d} = rac{y_{ub}u_{\psi d}}{M^2} \left(b_R u_R\right) \left(\psi_{B,R} d_R\right)$$

Recall for *B*-Mesogenesis:

$$Br(B \to \psi_B + \mathcal{B}_{SM}) \gtrsim 10^{-4}$$

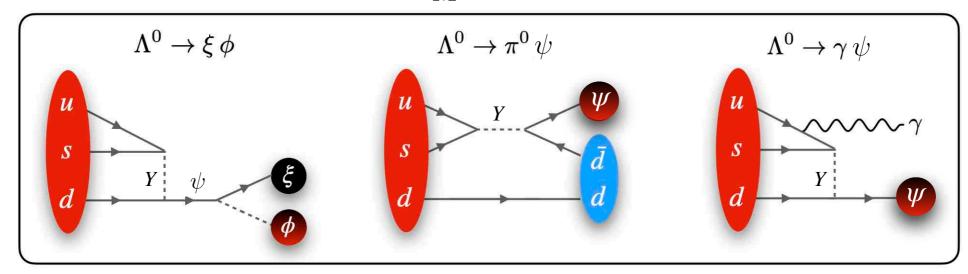


#### Discovering Neutral B-Mesogenesis



#### Indirect Signals: New Hyperon Decays

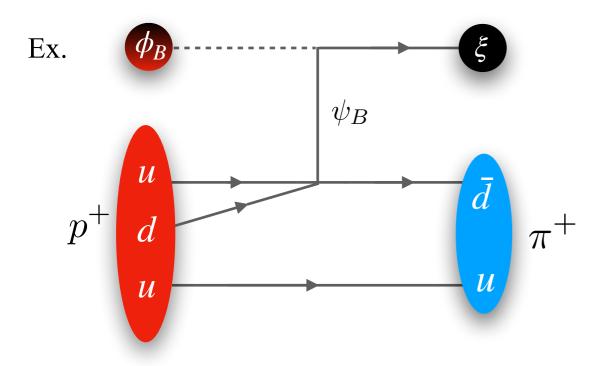
$$\mathcal{O}_{u_a d_b, d_c}^{L,R} = \frac{y_{u_a d_b} y_{\psi d_c}}{M^2} \left( u_a d_b \right)_{L,R} \left( \psi_B d_c \right)_R$$



Initial State	Final State	${ m Max}[m_\chi] \; ({ m MeV})$	Operator	$\operatorname{Br} \times \left[\frac{C_{ud_a,d_b}^{L/R}}{0.5  \text{TeV}^{-2}}\right]^2  (\text{LHC})$	SN 1987A Limit	BESIII
	$\chi + \gamma$	1116	$\mathcal{O}^{L/R}_{us,d}$	$3.5 \times 10^{-6}$	$\sim 10^{-7}$	_
$\Lambda\left(uds ight)$	$\xi + \phi$	1116	$\mathcal{O}^{L/R}_{ud,s}$	$5.0 \times 10^{-4}  y_{\xi\phi} ^2$	$\sim 10^{-8}$	$7 \times 10^{-5} [50]$
	$\chi + \pi^0$	981	$\mathcal{O}^L_{us,d}$	$1.1 \times 10^{-3}$	$\sim 10^{-7}$	_
$\Sigma^{+} (uus)$	$\chi + \pi^+$	1050	$\mathcal{O}^{L/R}_{us,d}$	$1.6 \times 10^{-3}$	$\sim 10^{-4}$	-
$\Sigma^{-} (dds)$	$\chi + \pi^-$	1058	$\mathcal{O}^{L/R}_{ud,s}$	$1.2 \times 10^{-3}$	$\sim 10^{-4}$	-
	$\chi + \gamma$	1315	$\mathcal{O}^{L/R}_{us,s}$	$1.5 \times 10^{-5}$	_	_
$\Xi^{0}\left(uss ight)$	$\xi + \phi$	1315	$\mathcal{O}^{L/R}_{us,s}$	$3.7 \times 10^{-4}  y_{\xi\phi} ^2$	_	_
	$\chi + \pi^0$	1180	$\mathcal{O}^R_{us,s}$	$6.2 \times 10^{-3}$	_	-
$\Xi^{-}\left( dss ight)$	$\chi + \pi^-$	1182	$\mathcal{O}^R_{us,s}$	$7.1 \times 10^{-3}$	_	_

## Indirect Signals: Induced Nucleon Decay at DUNE

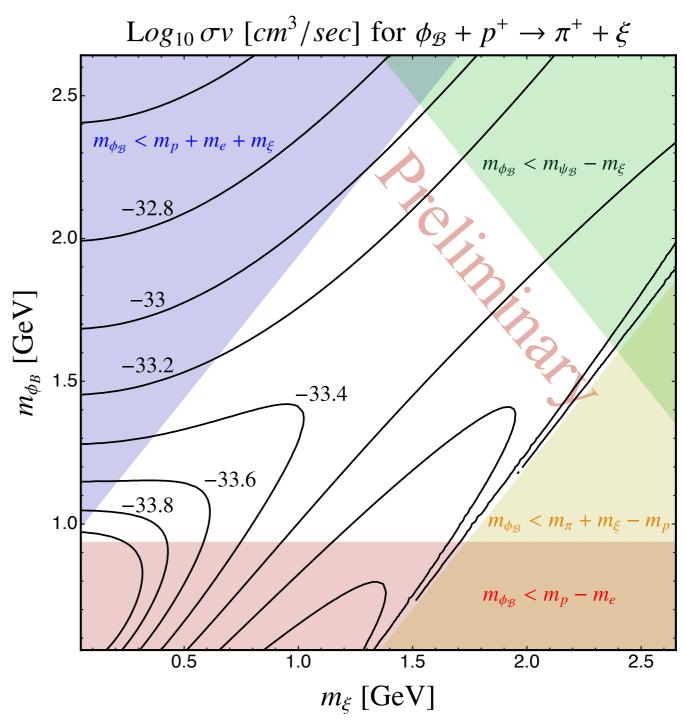
[arXiv:22xx.xxxx] GE with Josh Berger



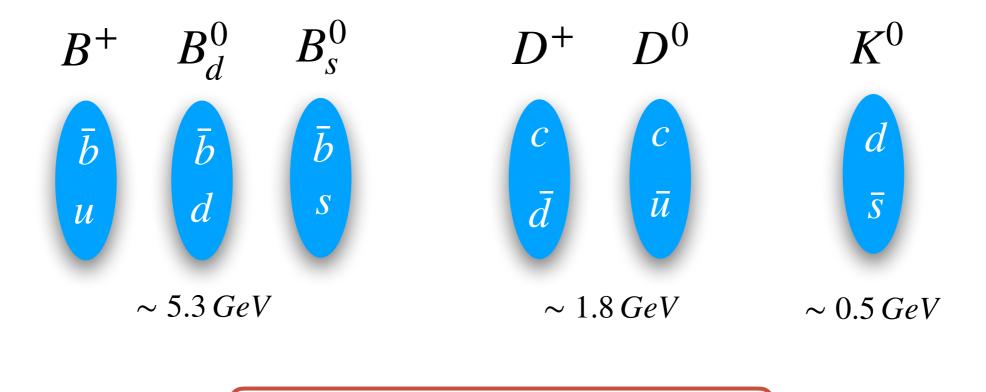
Mono-energetic meson:

$$E_{\mathcal{M}} = rac{m_{\mathcal{M}}^2 - m_{\xi}^2}{2(m_N + m_{\phi_B})} + rac{1}{2} \left( m_N + m_{\phi_B} 
ight) + \mathcal{O}(v_{\phi_B})$$

DUNE: 40 kiloton liquid argon detector. Expect sensitivity down to  $\sigma v \sim 10^{-40} \text{cm}^3/\text{sec}$  after one year of running.



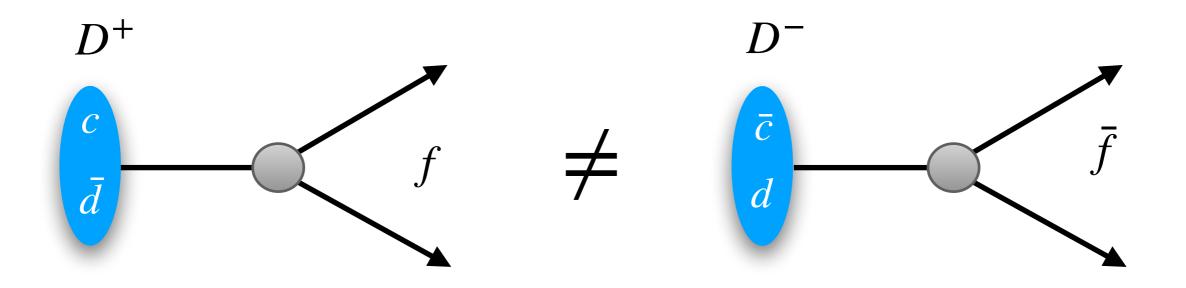
#### Why Neutral B Mesons?



• Kinematics: Dark baryons must be GeV scale. Only *B* mesons are heavy enough to decay into GeV scale. Charge dark particle under lepton number instead, then it can be light.

 $m_{\psi_B} > m_p - m_e \simeq 937.8 \,\mathrm{MeV}$ 

## Charged D Mesogenesis



Observable: 
$$A_{CP}^f = \frac{\Gamma(D^+ \to f) - \Gamma(D^- \to \bar{f})}{\Gamma(D^+ \to f) + \Gamma(D^- \to \bar{f})}$$

### CPV in Charged D Decays

Example: Standard Model decays to an odd number of charged pions

$D^+$ decay mode	$A_{CP}^f/10^{-2}$
$K_S^0\pi^+$	$-0.41 \pm 0.09$
$K^-\pi^+\pi^+$	$-0.18 \pm 0.16$
$K^-\pi^+\pi^+\pi^0$	$-0.3 \pm 0.6 \pm 0.4$
$K_S^0\pi^+\pi^0$	$-0.1 \pm 0.7 \pm 0.2$
$K_S^0 \pi^+ \pi^+ \pi^-$	$0.0 \pm 1.2 \pm 0.3$
$\pi^+\pi^0$	$2.4 \pm 1.2$
$\pi^+\eta$	$1.0 \pm 1.5$

$D^+$ decay mode	$A_{CP}^f/10^{-2}$
$\pi^+\eta$	$1.0 \pm 1.5$
$\pi^+\eta'(958)$	$-0.6 \pm 0.7$
$K^+K^-\pi^+$	$0.37 \pm 0.29$
$\phi\pi^+$	$0.01 \pm 0.09$
$a_0(1450)^0\pi^+$	$-19 \pm 12^{+8}_{-11}$
$\phi(1680)\pi^{+}$	$-9 \pm 22 \pm 14$
$\pi^+\pi^+\pi^-$	$-1.7 \pm 4.2$

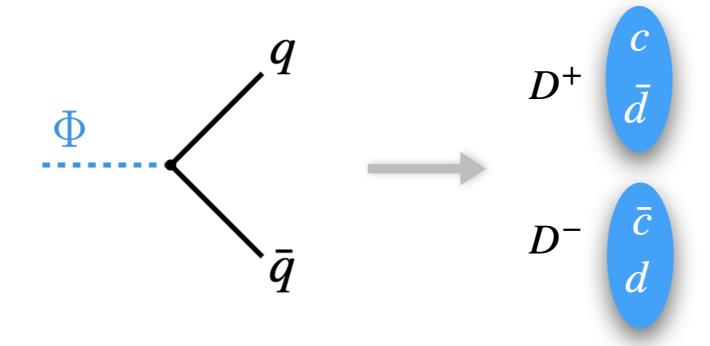
Not a small number if we want to explain

$$Y_B^{\text{obs}} = (8.718 \pm 0.004) \times 10^{-11}$$

#### Sakharov I. Out of Equilibrium

#### Late decay of an "inflaton-like" field

Decays at:  $\Gamma_{\Phi} = H(T_R)$  to quarks  $m_{\Phi} \in [5 \, \mathrm{GeV}, 100 \, \mathrm{GeV}]$ 



 $3.5\,\mathrm{MeV}\,\lesssim\,T_\mathrm{R}\,\lesssim\,20\,\mathrm{MeV}$ 



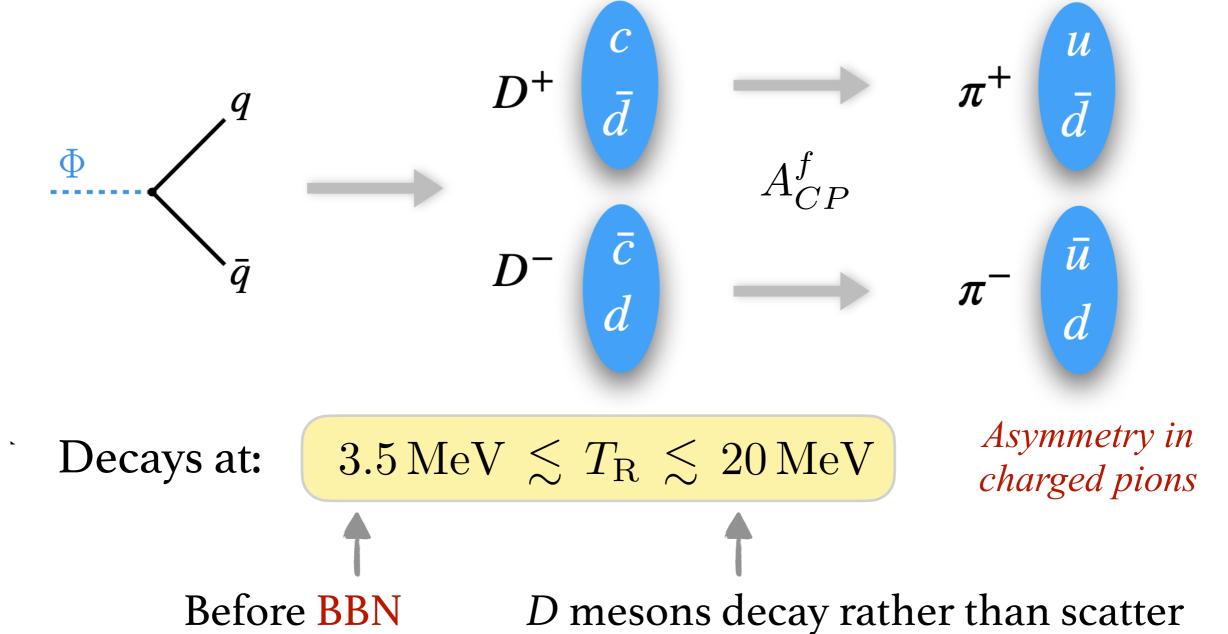
Before BBN



After QCD phase transition

#### Sakharov II. CP Violation

D mesons quickly undergo Standard Model decays to pions



#### Sakharov III. B Violation?

Need a way to change baryon number



Hide baryon *and lepton* number in a dark sector without violating either.

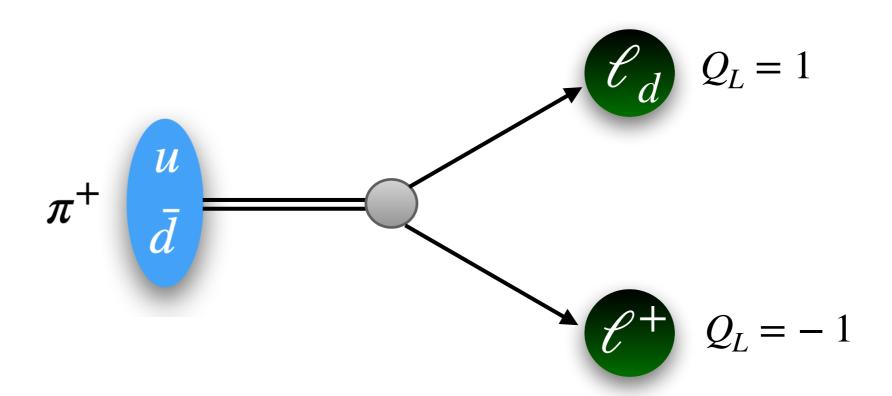


First generate a lepton asymmetry

### Dark Sector Lepton

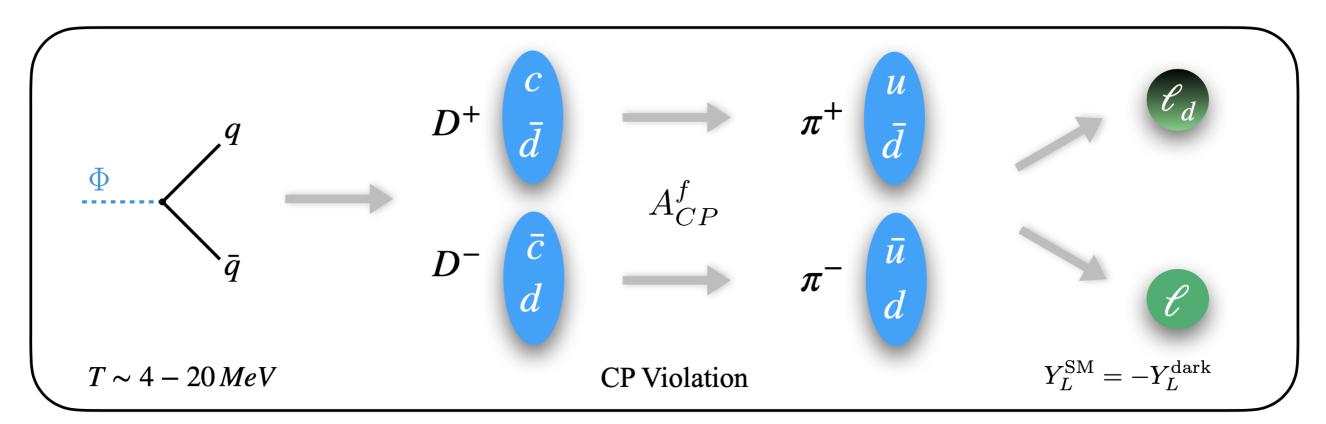
Portal Operator: 
$$\mathcal{O} = \frac{1}{\Lambda^2} \left[ \bar{d} \Gamma^{\mu} u \right] \left[ \bar{\ell}_d \Gamma_{\mu} \ell \right] + \text{h.c.}$$

Pion Decays:  $\pi^+ \to \ell_d + \ell^+$ ,  $m_{\ell_d} < m_{\pi^+} - m_{\ell}$  Can be light



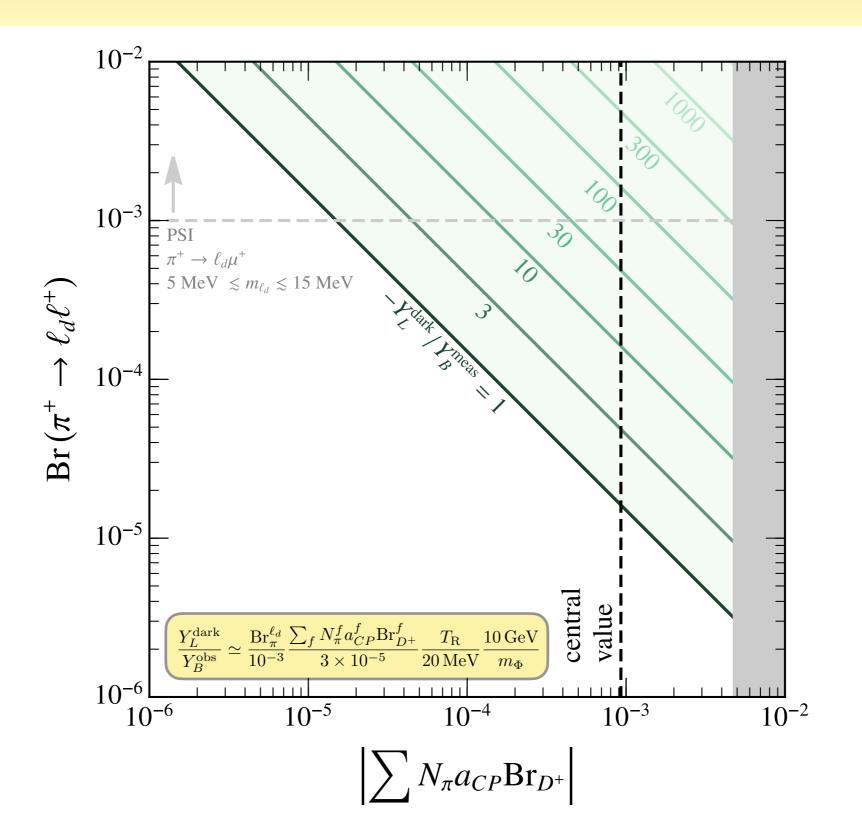
## Generating a Lepton Asymmetry

#### Equal and opposite dark/visible sector lepton asymmetry

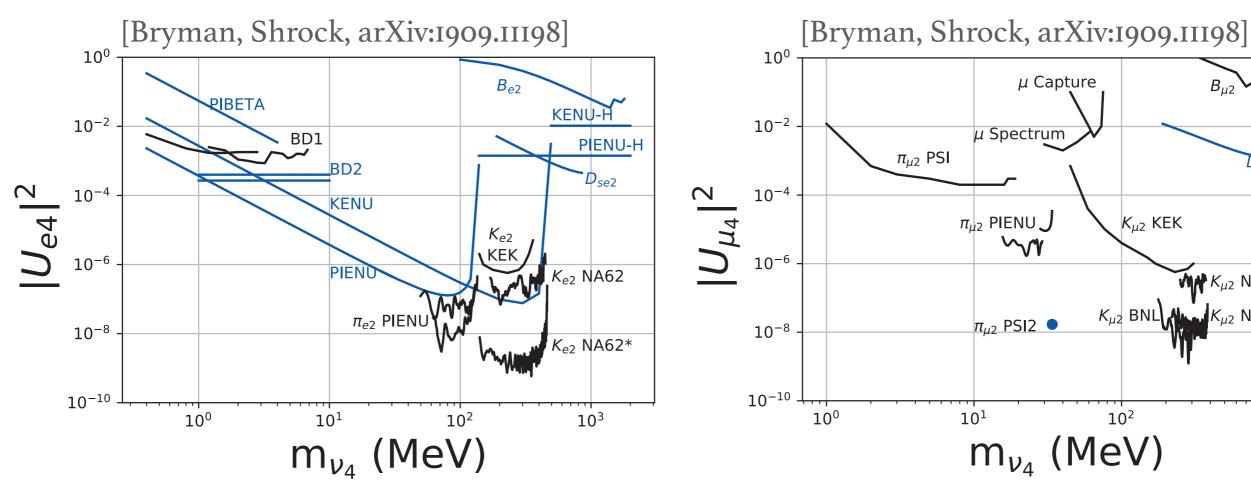


$$Y_L^{\text{dark}} \equiv \left(\frac{n_{\ell_d} - n_{\bar{\ell_d}}}{s}\right) \propto \text{Br}\left(\pi^+ \to \ell_d + \ell^+\right) \sum_f A_{\text{CP}}^f \times \text{Br}\left(D^+ \to f\right)$$

### Generating a Lepton Asymmetry



### Limits on Pion Decays



 $K_{\mu 2}$  KEK **γ**γγ κ<sub>μ2</sub> NA62 10<sup>3</sup>  $m_{\nu_4}$  (MeV)

Limit on  $|U_{\ell N}|^2 \Rightarrow \text{limit on } \frac{\Gamma(\pi^{\pm} \to \ell^{\pm} + \ell_d)}{\Gamma(\pi^{\pm} \to \ell^{\pm} + \nu_{\text{CM}})}$ 

[Shrock, Phys. Rev. D24, 1232 (1981)]

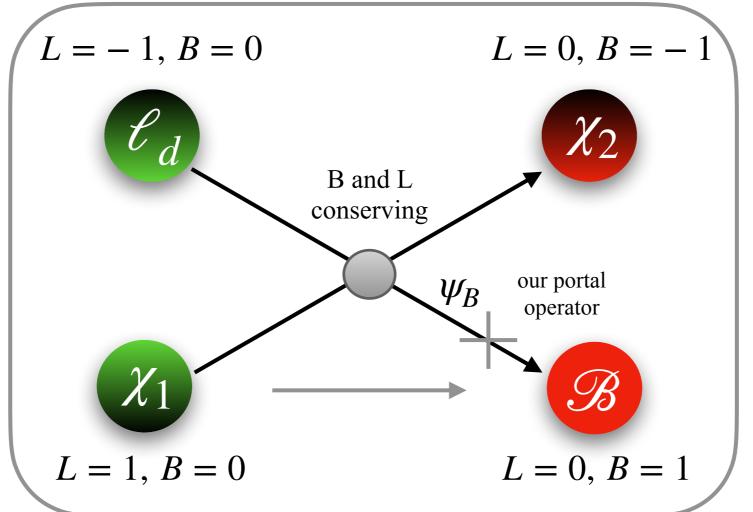
$$Br(\pi^{\pm} \to \mu^{\pm} + MET) \lesssim 10^{-3}$$
, for  $5 \, MeV < m_{\ell_d} < 15 \, MeV$ .

G. Elor

## Generating a Baryon Asymmetry?

At 20 MeV, you (of course) can not use Electroweak Sphalerons to transfer a lepton into a baryon asymmetry.

You also don't need them...



Dark Baryon

SM Baryon

Dark Leptons

### Freezing-In a Baryon Asymmetry

#### **Example Benchmark point:**

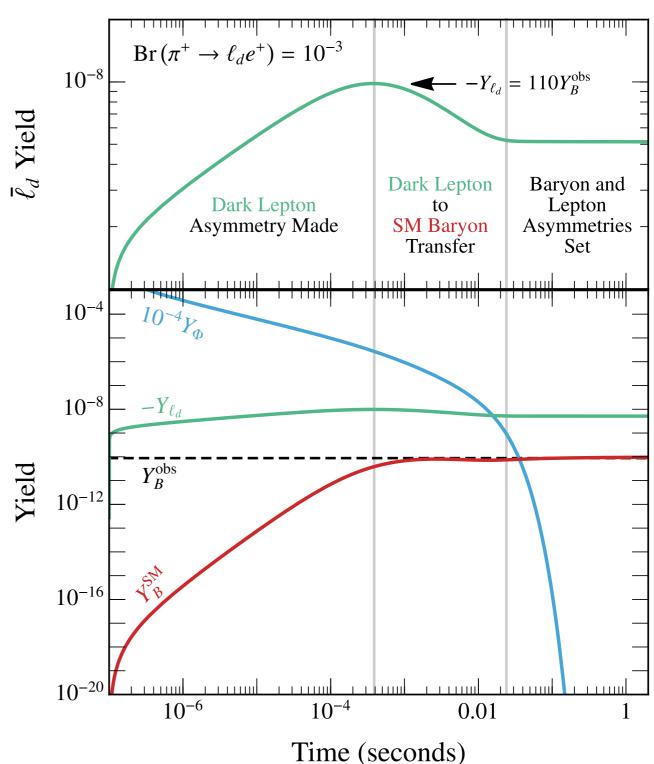
$$T_R = 10 \text{ MeV}, m_{\Phi} = 6 \text{ GeV}$$
  
 $\langle \sigma v \rangle = 1 \times 10^{-15} \text{ GeV}^{-2}$ 

$$Br\left(\Phi \to \chi_1 \bar{\chi}_1\right) = 0.1$$

$$\sum_{f} N_{\pi}^{f} a_{CP}^{f} Br_{D^{+}}^{f} = \left(-9.3 \times 10^{-4}\right)$$

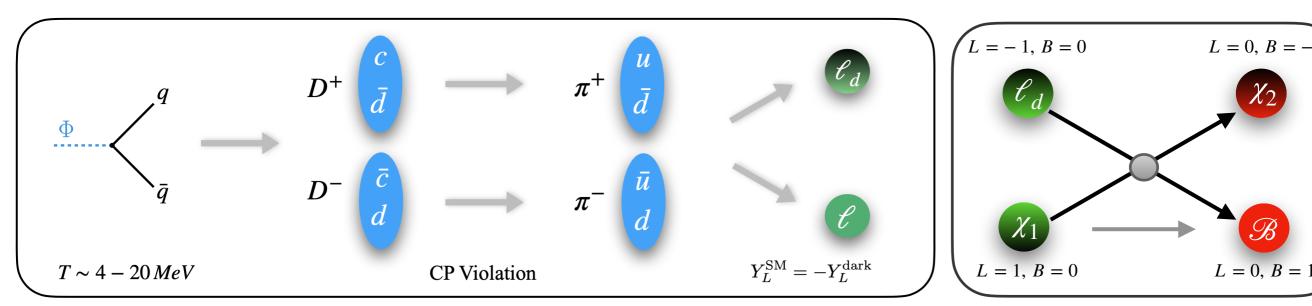
$$\frac{d}{dt} (n_{\mathcal{B}} - n_{\overline{\mathcal{B}}}) + 3H (n_{\mathcal{B}} - n_{\overline{\mathcal{B}}}) =$$
$$-\langle \sigma v \rangle n_{\chi_1} (n_{\ell_d} - n_{\bar{\ell}_d})$$

$$\left. \frac{n_{\chi_1} \langle \sigma v \rangle}{H(T)} \right|_{T=T_R} \gtrsim \frac{Y_B^{\text{obs}}}{Y_L^{\text{dark}}}$$



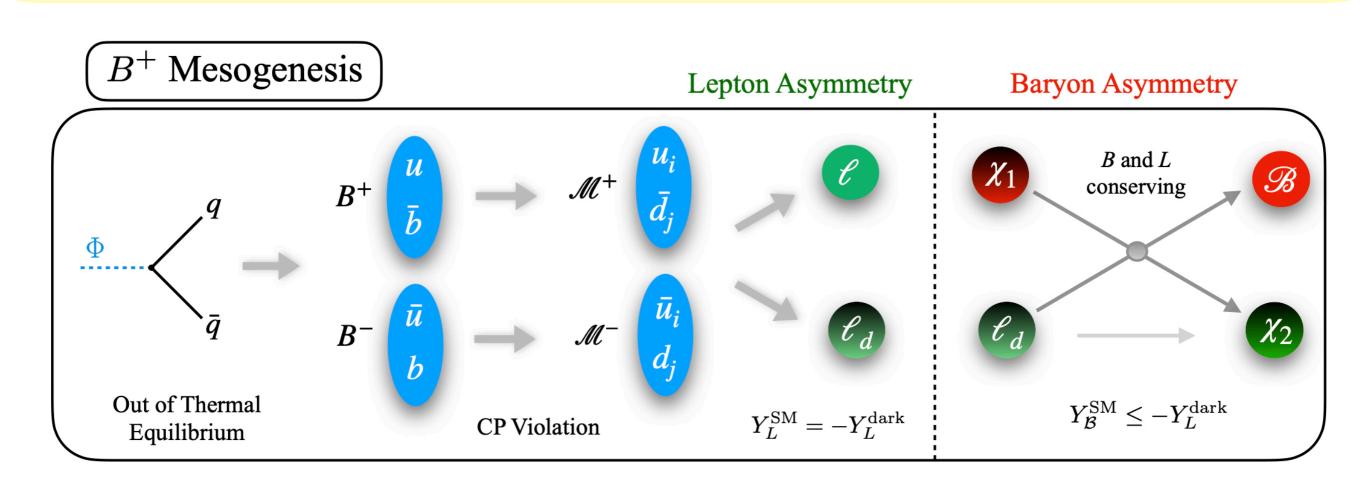
G. Elor

## D<sup>+</sup> Mesogenesis



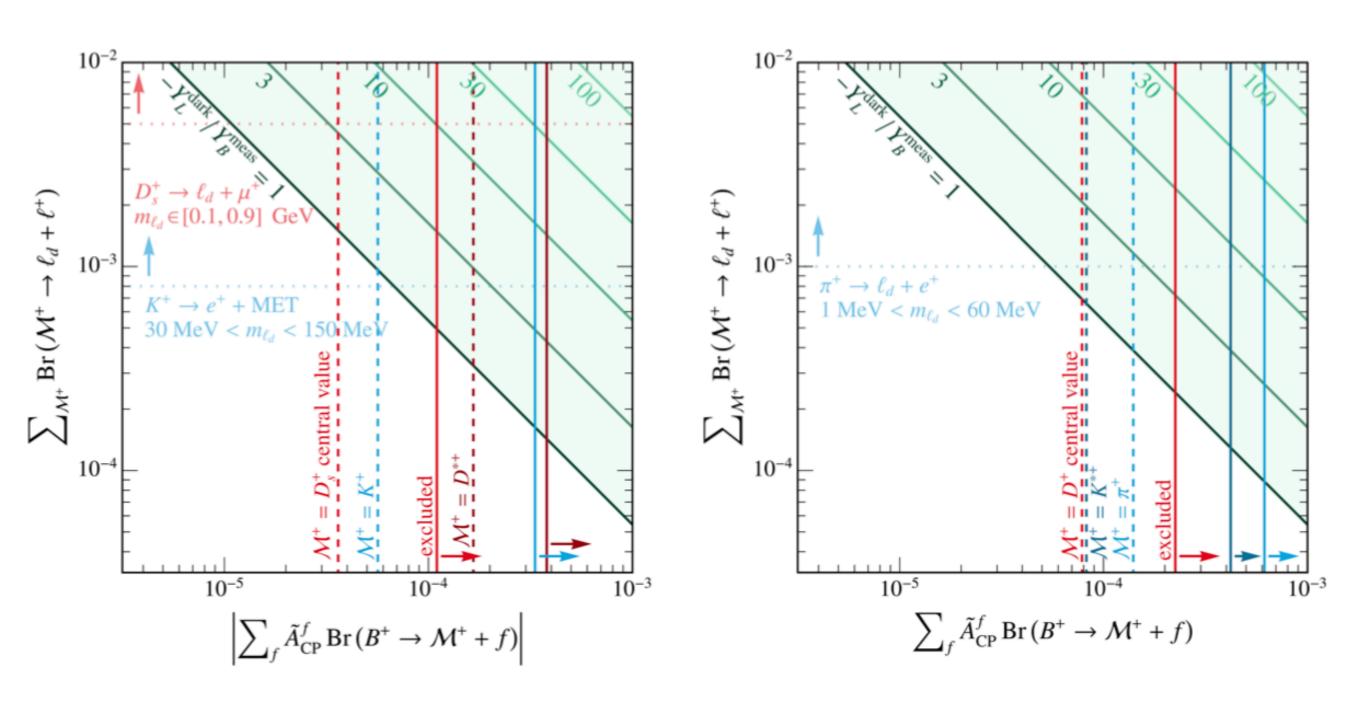
- First generates a lepton asymmetry and then freezes in a baryon asymmetry through dark sector scatterings.
- Baryogenesis and dark matter production are controlled by experimental observables of the charged *D* Mesons system.

# B<sup>+</sup> Mesogenesis

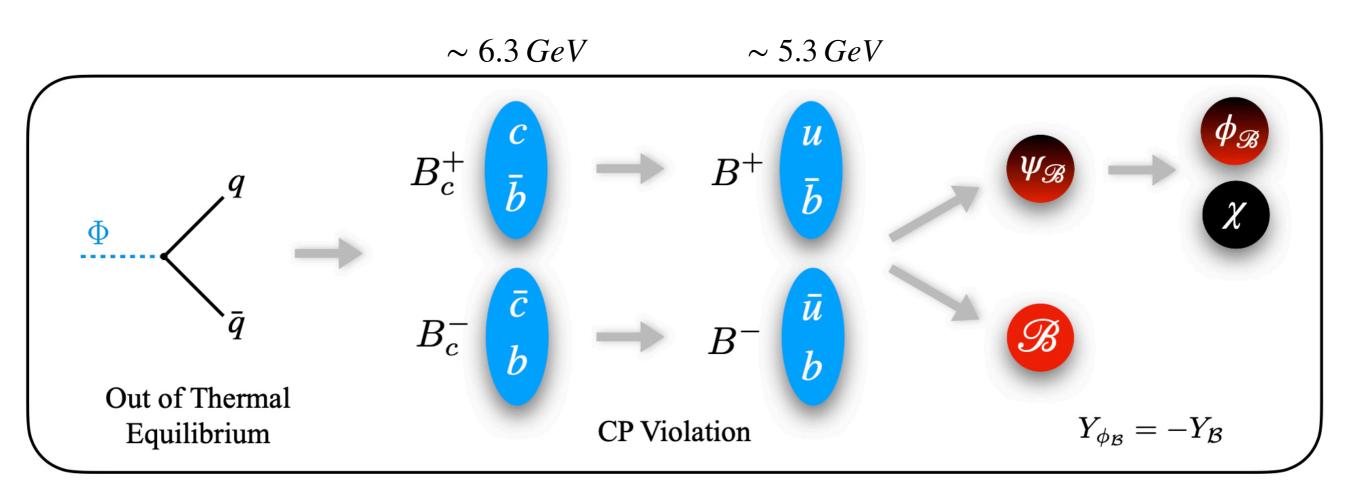


$$Y_{\ell_d} \propto \sum_{\mathcal{M}^+} \operatorname{Br}_{\mathcal{M}^+}^{\ell_d} \sum_f \tilde{A}_{\operatorname{CP}}^f \operatorname{Br}_{B^+}^f$$
 
$$\tilde{A}_{\operatorname{CP}}^f = \frac{\Gamma(B^+ \to f) - \Gamma(B^- \to f)}{\Gamma(B^+ \to f) + \Gamma(B^- \to f)}$$

# B<sup>+</sup> Mesogenesis



# B<sub>c</sub><sup>+</sup> Mesogenesis



$$A_{\text{CP}}^{f} = \frac{\Gamma(B_c^+ \to f) - \Gamma(B_c^- \to \bar{f})}{\Gamma(B_c^+ \to f) + \Gamma(B_c^- \to \bar{f})} \qquad \mathcal{O} = \frac{y^2}{M_{\phi}^2} \bar{\psi}_{\mathcal{B}} b \bar{u}_i^{\text{c}} d_j + \text{h.c.},$$

$$m_{\psi_B} > m_p - m_e \simeq 937.8 \,\text{MeV}$$

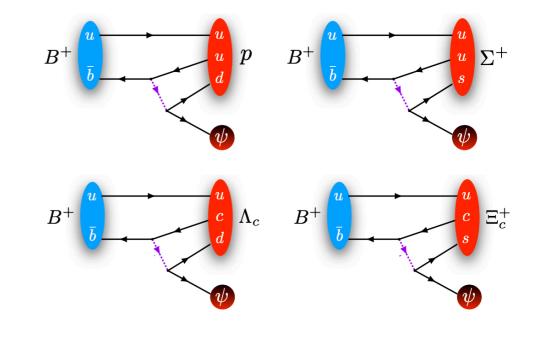
## $B^+$ Decay

UV Model: 
$$\mathcal{L}_{-1/3} = -\sum_{i,j} y_{u_i d_j} Y^* \bar{u}_{iR} d^c_{jR} - \sum_k y_{\psi d_k} Y d^c_{kR} \bar{\psi} + \text{h.c.}$$

Operator/Decay	Initial State	Final state		
	$B_d$	$\psi + n  (udd)$		
$\mathcal{O} = \psi  b  u  d$	$B_s$	$\psi + \Lambda \left( uds \right)$		
$\bar{b} \to \psi  u  d$	$B^+$	$\psi + p \left( duu \right)$		
	$\Lambda_b$	$\bar{\psi} + \pi^0$		
	$B_d$	$\psi + \Lambda \left( usd \right)$		
$\mathcal{O} = \psi  b  u  s$	$B_s$	$\psi + \Xi^0 \left( uss \right)$		
$\bar{b} \rightarrow \psi  u  s$	$B^+$	$\psi + \Sigma^+ (uus)$		
	$\Lambda_b$	$\bar{\psi} + K^0$		
	$B_d$	$\psi + \Lambda_c + \pi^- (cdd)$		
$\mathcal{O} = \psi  b  c  d$	$B_s$	$\psi + \Xi_c^0 \left( cds \right)$		
$\overline{b} \rightarrow \psi  c  d$	$B^+$	$\psi + \Lambda_c \left( dcu \right)$		
	$\Lambda_b$	$\overline{\psi} + \overline{D}^0$		
	$B_d$	$\psi + \Xi_c^0 \left( csd \right)$		
$\mathcal{O} = \psi  b  c  s$	$B_s$	$\psi + \Omega_c \left( css \right)$		
$\bar{b} \rightarrow \psi  c  s$	$B^+$	$\psi + \Xi_c^+ (csu)$		
	$\Lambda_b$	$\bar{\psi} + D^- + K^+$		

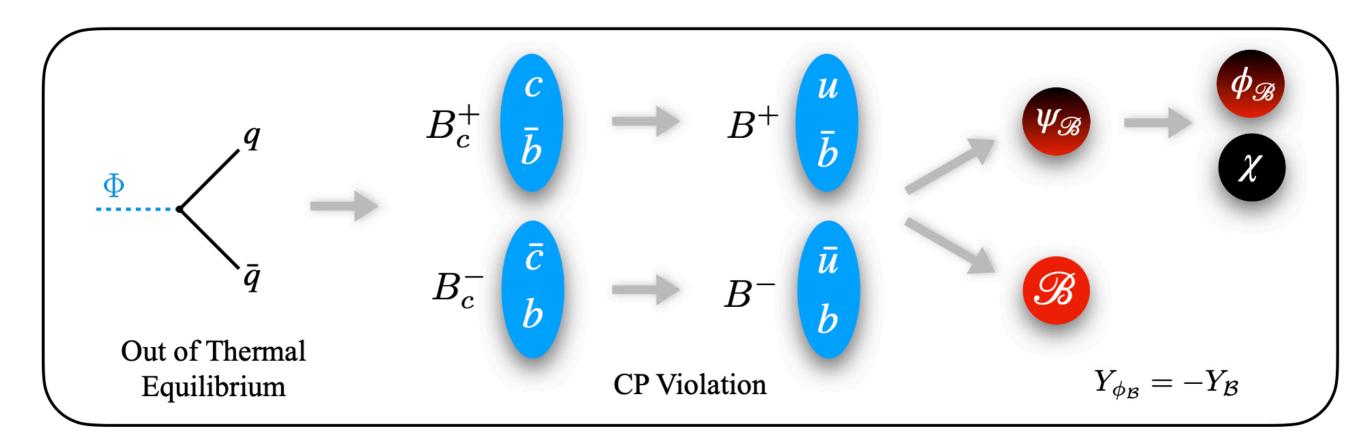
Directly related to neutral B Mesogenesis, and indirectly related  $B^+$  Mesogenesis.

Directly related to  $B^+$  Mesogenesis.



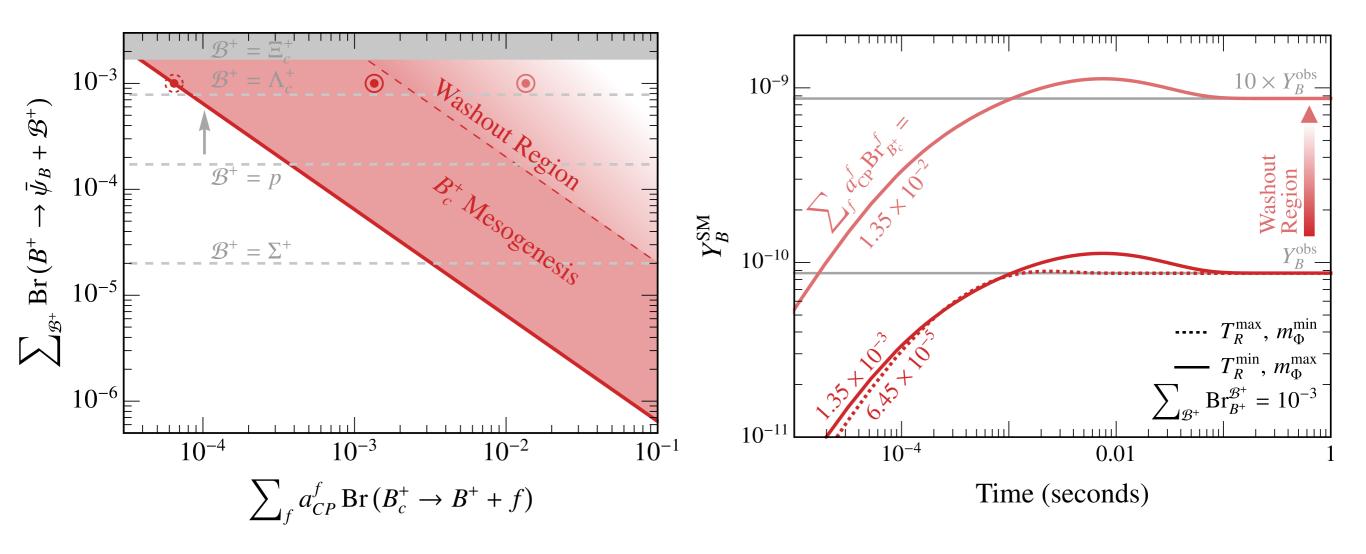
Indirect signal of charged and neutral *B*Mesogenesis
G. Elor

# $B_c^+$ Mesogenesis



$$Y_{\mathcal{B}} \equiv \frac{n_{\mathcal{B}} - n_{\bar{\mathcal{B}}}}{s} \propto \sum_{f} A_{\mathrm{CP}}^{f} \operatorname{Br} \left( B_{c}^{+} \to B^{+} + f \right) \times \sum_{\mathcal{B}^{+}} \operatorname{Br} \left( B^{+} \to \bar{\psi}_{\mathcal{B}} + \mathcal{B}^{+} \right)$$

# B<sub>c</sub><sup>+</sup> Mesogenesis



$$\frac{Y_{\mathcal{B}}}{Y_{\mathcal{B}}^{\text{obs}}} \simeq \frac{\sum_{\mathcal{B}^{+}} \operatorname{Br}_{B^{+}}^{\mathcal{B}^{+}}}{10^{-3}} \frac{\sum_{f} a_{\text{CP}}^{f} \operatorname{Br}_{B_{c}^{+}}^{f}}{6.45 \times 10^{-5}} \frac{T_{R}}{20 \text{ MeV}} \frac{2m_{B_{c}^{+}}}{m_{\Phi}}$$

# The Many Flavors of Mesogenesis

Mechanism	CPV	Dark Sector	Observables	Relevant Experiments		
$B^0$ Mesogenesis	$B_s^0 \& B_d^0$	Dark baryons	$A_{sl}^{s,d}$	LHCb		
	oscillations		$Br(B \to \mathcal{B} + X)$	B Factories, LHCb		
			$A^D_{CP}$	B Factories, LHCb		
$D^+$ Mesogenesis	$D^{\pm}$ decays	Dark leptons	$\mathrm{Br}_{D^+}$	B Factories, LHCb		
		and/or baryons	$Br(\mathcal{M}^+ \to \ell^+ + X)$	peak searches e.g. PSI, PIENU		
			$A_{CP}^B$	B Factories, LHCb		
$B^+$ Mesogenesis	$B^{\pm}$ decays	Dark leptons	${\rm Br}_{B^+}$	B Factories, LHCb		
		and/or baryons	$Br(\mathcal{M}^+ \to \ell^+ + X)$	peak searches e.g. PSI, PIENU		
			$A_{CP}^{B_c}$	LHCb, FCC		
$B_c^+$ Mesogenesis	$B_c^{\pm}$ decays	Dark baryons	$\operatorname{Br}_{B_c^+}$	LHCb, FCC		
			$\operatorname{Br}_{B^+ \to \mathcal{B}^+ + X}$	B Factories, LHCb		

GE, M. Escudero, A. E. Nelson, PRD, [1810.00880]

G. Alonso-Alvarez, GE, A. E. Nelson, H. Xiao, JHEP, [1907.10612]

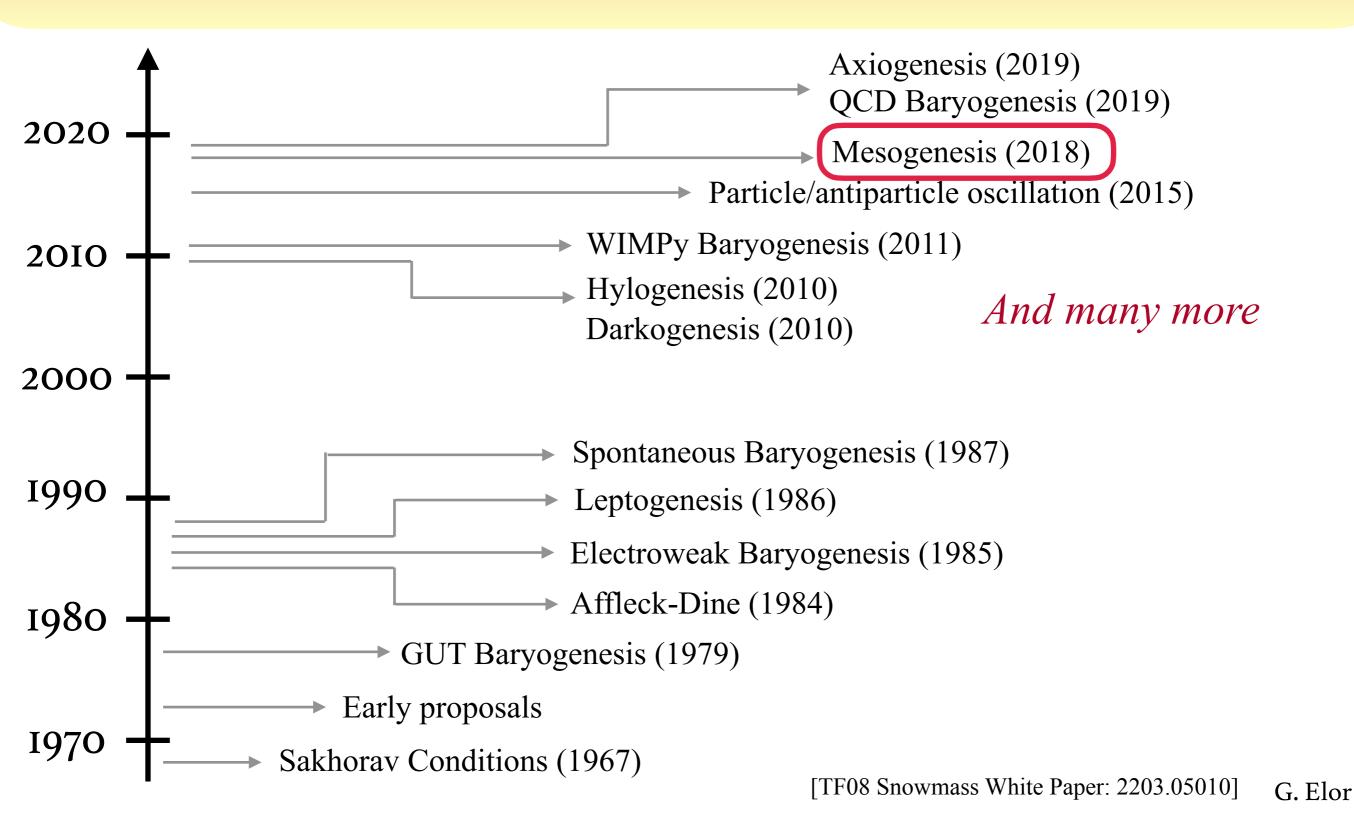
GE, R. McGehee, PRD [2011.06115]

G. Alonso-Alvarez, GE. M. Escudero, PRD, [2101.02706]

F. Elahi, GE, R. McGehee, [2109.09751]

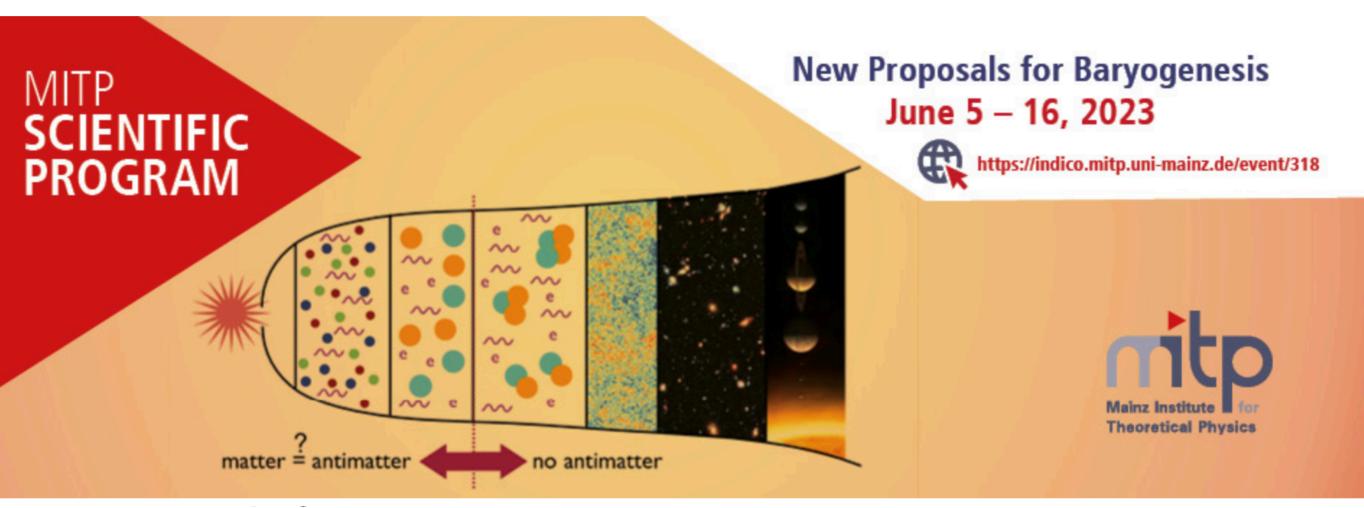
G. Alonso-Alvarez, GE, M. Escudero, B. Fornal, B. Grinstein, J.M. Camalich [arXiv:2111.12712]

# New Ideas Baryogenesis



## MITP Program 2023

Join us in Mainz in 2023!



Organizing with Seyda Ipek and Bibhushan Shakya

# The Many Flavors of Mesogenesis

Mechanism	CPV	Dark Sector	Observables	Relevant Experiments		
$B^0$ Mesogenesis	$B_s^0 \& B_d^0$	Dark baryons	$A_{sl}^{s,d}$	LHCb		
	oscillations		$\operatorname{Br}(B \to \mathcal{B} + X)$	B Factories, LHCb		
			$A^D_{CP}$	B Factories, LHCb		
$D^+$ Mesogenesis	$D^{\pm}$ decays	Dark leptons	$\mathrm{Br}_{D^+}$	B Factories, LHCb		
		and/or baryons	$\operatorname{Br}(\mathcal{M}^+ \to \ell^+ + X)$	peak searches e.g. PSI, PIENU		
			$A_{CP}^{B}$	B Factories, LHCb		
$B^+$ Mesogenesis	$B^{\pm}$ decays	Dark leptons	$\mathrm{Br}_{B^+}$	B Factories, LHCb		
		and/or baryons	$\operatorname{Br}(\mathcal{M}^+ \to \ell^+ + X)$	peak searches e.g. PSI, PIENU		
			$A_{CP}^{B_c}$	LHCb, FCC		
$B_c^+$ Mesogenesis	$B_c^{\pm}$ decays	Dark baryons	$\operatorname{Br}_{B_c^+}$	LHCb, FCC		
			$\operatorname{Br}_{B^+ \to \mathcal{B}^+ + X}$	B Factories, LHCb		

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G. Alonso-Alvarez, GE, M. Escudero, B. Fornal, B. Grinstein, J.M. Camalich [arXiv:2111.12712]

Thanks!

# Backups

# Emphasis on Testability

#### New Ideas in Baryogenesis: A Snowmass White Paper

Editors: Gilly Elor, 1 Julia Harz, 2 Seyda Ipek, 3 Bibhushan Shakya. 4

Authors: Nikita Blinov,<sup>5</sup> Raymond T. Co,<sup>6</sup> Yanou Cui,<sup>7</sup> Arnab Dasgupta,<sup>8</sup> Hooman Davoudiasl,<sup>9</sup> Fatemeh Elahi,<sup>1</sup> Gilly Elor,<sup>1</sup> Kåre Fridell,<sup>2</sup> Akshay Ghalsasi,<sup>8</sup> Keisuke Harigaya,<sup>10</sup> Julia Harz,<sup>2</sup> Chandan Hati,<sup>2</sup> Peisi Huang,<sup>11</sup> Seyda Ipek,<sup>3</sup> Azadeh Maleknejad,<sup>10</sup> Robert McGehee,<sup>12</sup> David E. Morrissey,<sup>13</sup> Kai Schmitz,<sup>10</sup> Bibhushan Shakya,<sup>4</sup> Michael Shamma,<sup>13</sup> Brian Shuve,<sup>14</sup> David Tucker-Smith,<sup>15</sup> Jorinde van de Vis,<sup>4</sup> Graham White.<sup>16</sup>

Model	Key Ingredients	Observable scale	Observables	
Axiogenesis	Axion misalignment, sphalerons	axion scale ~ O( $10^{8-11}$ GeV) axion mass ~ $\mu$ eV	Gravitational waves	
$W_R$ baryogenesis	axion inflation, $W_R$ interactions with the inflaton	LR symmetry breaking ~ O( 1010 GeV)	Gravitational waves	
QCD Baryogenesis	Singlet scalar coupled to the gluon field strength, axion, sphalerons	masses ~ O(10 GeV) temperature ~ O(TeV)	Scalar field mixing with the Higgs Graviational waves	
Wash-in Leptogenesis	Charge asymmetry instead of <i>B-L</i> , out of equilibrium decays	Right-handed neutrino masses ~ O(100 TeV)	Charged lepton flavor violation	
Hylogenesis	Long-lived dark baryons	GeV-TeV	Induced nucleon decay collider signatures	
WIMPy Baryogenesis	Metastable WIMPs	O(100 GeV)	Long-lived particles	
Gaugino Portal Baryogenesis	hidden sector gaugino-bino mixing, R-parity violation	masses ~ O(10 - 108 GeV)	Neutron-antineutron oscillation, LLP (RPV decays) at colliders	
Freeze-in Baryogenesis	DM oscillations	masses ~ O(TeV), O(10 keV)	missing momentum searches structure formation, X-ray signals	
Pseudogenesis	Pseudogenesis Pseudo-Dirac fermions, particle-antiparticle oscillations		LLPs, dilepton asymmetry	
Mesino-genesis Mesino-antimesino oscillations, SU(3)-charged scalars		$masses \sim O(TeV)$ $temperature \sim O(100 MeV)$	LLPs, same-sign top quark decays, mullti-jet signals	
Mesogenesis CPV from SM Meson systems, dark states charged under SM B and L number		masses ~ O(1-100 GeV) temperature ~ O(5-100 MeV)	CPV observables at B factories, LHC, decays of hadrons to dark baryons, peak searches at colliders	
Baryogenesis from Quantum Statistics dark matter chemical potential		_	_	

#### Flavorful Variations

No a priori reason to expect a particular flavor structure.

Most general interactions:

$$\left(\mathcal{L}_{-1/3} = -\sum_{i,j} y_{u_i d_j} Y^* \bar{u}_{iR} d_{jR}^c - \sum_k y_{\psi d_k} Y d_{kR}^c \bar{\psi} + \text{h.c.}\right)$$

Possible operators:

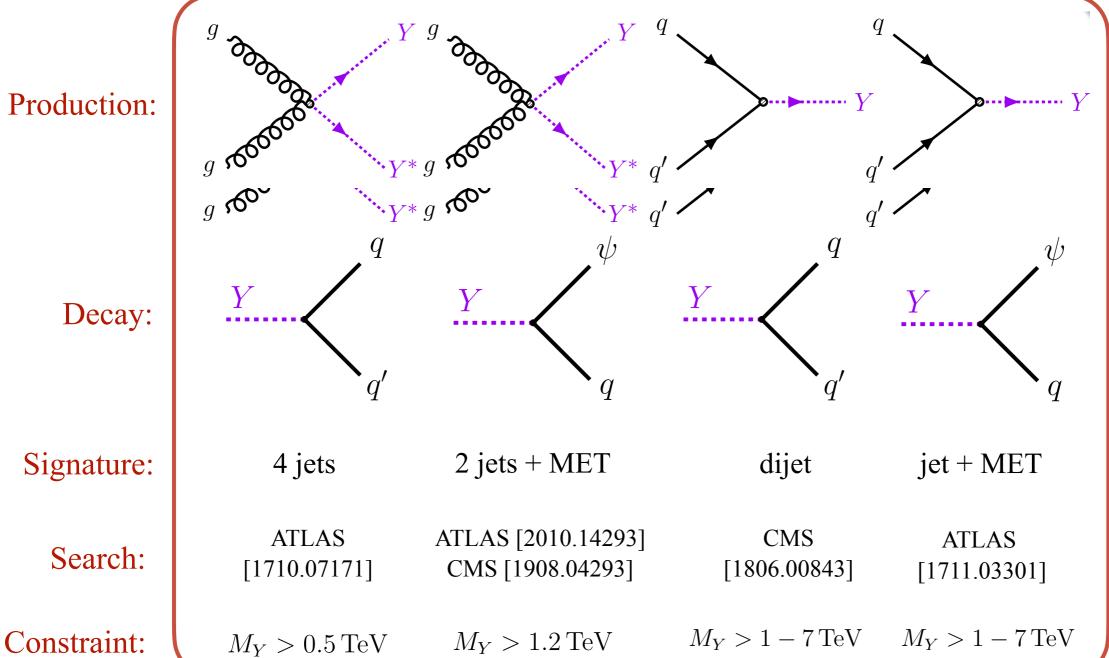
$$\mathcal{O}_{ud} = \psi \, b \, u \, d$$
 $\mathcal{O}_{us} = \psi \, b \, u \, s$ 
 $\mathcal{O}_{cd} = \psi \, b \, c \, d$ 
 $\mathcal{O}_{cs} = \psi \, b \, c \, s$ 

*B*-Mesogenesis requires:

 $Br(B \to \psi \mathcal{B} \mathcal{M}) \gtrsim 10^{-4}$ 

## Colored Triplet Scalar

#### Constraints from LHC squark searches



 $d_j$ 

 $d_i$ 

### Searching for new b-Hadron Decays

#### Possibilities at LHCb

[See our white paper on "Stealth Physics at LHCb" 2105.12668]

- No handle on initial energy of decaying *B* meson so measuring missing energy is non-trivial.
- But, LHCb has advantages: larger number of *B* mesons produced than at Belle, excellent vertex resolution, and good particle reconstruction efficiencies.
- Some possibilities for searches do exist. e.g. new paper just last week!

#### Prospects on searches for baryonic Dark Matter produced in b-hadron decays at LHCb [2106.12870]

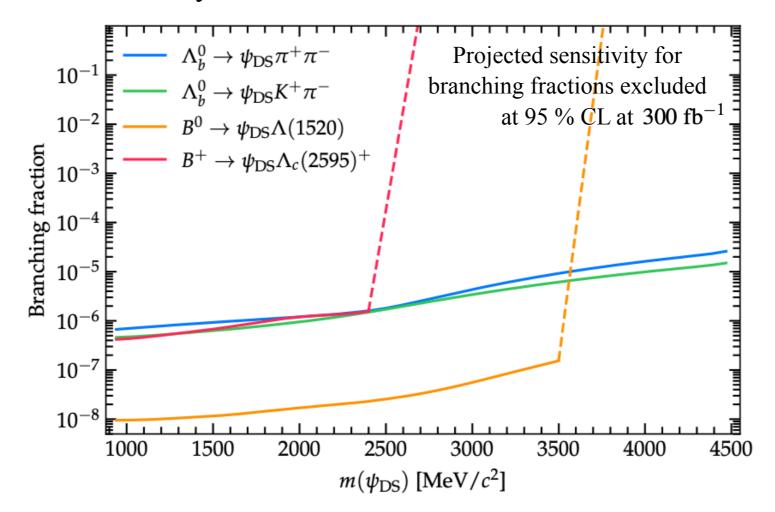
Alexandre Brea Rodríguez <sup>a,1</sup>, Veronika Chobanova <sup>b,1</sup>, Xabier Cid Vidal <sup>c,1</sup>, Saúl López Soliño <sup>d,1</sup>, Diego Martínez Santos <sup>e,1</sup>, Titus Mombächer <sup>f,1</sup>, Claire Prouvé <sup>g,1</sup>, Emilio Xosé Rodríguez Fernández <sup>h,1</sup>, Carlos Vázquez Sierra <sup>i,2</sup>

<sup>&</sup>lt;sup>1</sup>Instituto Galego de Física de Altas Enerxías (IGFAE), Universidade de Santiago de Compostela, 15782 Santiago de Compostela, Spain <sup>2</sup>European Organization for Nuclear Research (CERN), Geneva, Switzerland

#### Searching for new b-Hadron Decays

#### Proposed Search at LHCb [2106.12870]

- Search for decays of B mesons and b-Flavored baryons into an excited baryon in the final state  $B \to \psi \mathcal{B}^*$
- The excited baryon promptly decay at the same decay point as original decay, allowing one to trigger on this decay.



## Freezing-In a Baryon Asymmetry

#### Boltzmann Equations with scattering: $\bar{\ell}_d + \chi_1 \rightarrow \chi_2 + \mathcal{B}$

• New dark lepton/lepto-baryon:  $m_{\Phi} \gtrsim m_{\chi_1} \quad m_{\Phi} \gtrsim m_{\chi_2} + m_{\mathcal{B}}$ 

$$\frac{dn_{\chi_1}}{dt} + 3Hn_{\chi_1} = \Gamma_{\Phi} n_{\Phi} \operatorname{Br} \left(\Phi \to \chi_1 \bar{\chi}_1\right) - \langle \sigma v \rangle n_{\bar{\ell}_d} n_{\chi_1}$$

Dark lepton:

$$\frac{d}{dt} \left( n_{\ell_d} - n_{\bar{\ell}_d} \right) + 3H \left( n_{\ell_d} - n_{\bar{\ell}_d} \right) = 2\Gamma_{\Phi}^D n_{\Phi} \operatorname{Br}_{\pi}^{\ell_d} \sum_{f} N_{\pi}^f a_{CP}^f \operatorname{Br}_{D^+}^f - \langle \sigma v \rangle n_{\chi_1} \left( n_{\ell_d} - n_{\bar{\ell}_d} \right)$$

Baryon asymmetry:

$$\left(\frac{d}{dt}\left(n_{\mathcal{B}} - n_{\overline{\mathcal{B}}}\right) + 3H\left(n_{\mathcal{B}} - n_{\overline{\mathcal{B}}}\right) = -\langle \sigma v \rangle n_{\chi_{1}} \left(n_{\ell_{d}} - n_{\overline{\ell}_{d}}\right)\right)$$

To efficiently transfer the asymmetry  $\left. \frac{n_{\chi_1} \left< \sigma v \right>}{H(T)} \right|_{T=T_R} \ \gtrsim \ \frac{Y_B^{
m obs}}{Y_L^{
m dark}}$ 

#### Boltzmann Equations: Lepton Asymmetry

• Inflaton: 
$$\frac{dn_{\Phi}}{dt} + 3Hn_{\Phi} = -\Gamma_{\Phi}n_{\Phi}$$

• Radiation: 
$$\frac{d\rho_{\rm rad}}{dt} + 4H\rho_{\rm rad} = +\Gamma_{\Phi}m_{\Phi}n_{\Phi}$$

• Hubble: 
$$H^2 = \frac{8\pi}{3M_{\rm Pl}^2} \left( \rho_{\rm rad} + m_{\Phi} n_{\Phi} \right) \qquad \Gamma_{\Phi} = 4H \left( T_R \right)$$

The dark lepton asymmetry:

$$\Gamma_{\Phi}^{D} \equiv \Gamma_{\Phi} \operatorname{Br}(\Phi \to c) \operatorname{Br}(c \to D)$$

$$\frac{d}{dt} \left( n_{\ell_d} - n_{\bar{\ell}_d} \right) + 3H \left( n_{\ell_d} - n_{\bar{\ell}_d} \right) = 2 \Gamma_{\Phi}^D n_{\Phi} \operatorname{Br}_{\pi}^{\ell_d} \sum_{f} N_{\pi}^f a_{CP}^f \operatorname{Br}_{D^+}^f$$

#### **Experimental Observables:**

$$a_{CP}^f \equiv A_{CP}^f/(1+A_{CP}^f) \approx A_{CP}^f$$
 LHCb, B  
 $\operatorname{Br}_{D^+}^f \equiv \operatorname{Br}(D^+ \to f)$ 

 $\operatorname{Br}_{\pi}^{\ell_d} \equiv \operatorname{Br}(\pi^+ \to \ell_d + \ell^+)$  PIENU, PSI, etc. Charged pion decays:

#### Dark Possibilities

$$\bar{\ell}_d + \chi_1 \to \chi_2 + \bar{\psi}_B$$

Field	L	В	Field	L	В
$\chi_1$	1	0	$\chi_1$	1	1
$\chi_2$	0	-1	$\chi_2$	0	0
$\chi_1$	0	1	$\chi_1$	0	0
$\chi_2$	1	0	$\chi_2$	-1	-1

## Models

## Proof of concept that what I have told you thus far is not (too) crazy.

Some example models/dark sector charge assignments.

$$\bar{\ell}_d + \chi_1 \to \chi_2 + \mathcal{B}$$

• Estimation of the scattering cross section to confirm it can be large enough to transfer the asymmetry given current constraints.

$$\langle \sigma v \rangle \gtrsim 10^{-16} \,\mathrm{GeV}^{-2} \, \frac{Y_B^{\mathrm{obs}}}{Y_L^{\mathrm{dark}}} \times \frac{10 \,\mathrm{GeV}}{m_\Phi} \frac{20 \,\mathrm{MeV}}{T_R} \frac{10^{-1}}{\mathrm{Br}(\Phi \to \chi_1 \bar{\chi_1})}$$

#### Portal to the Dark Sector

Model Build for: 
$$\bar{\ell}_d + \chi_1 \rightarrow \chi_2 + \mathcal{B}$$

New fields: (Same model as for *B*-Mesogenesis[arXiv:1810.00880])

Color triplet scalar mediator

Dark Baryon

Field	Spin	L	В	$\mathbb{Z}_2$	Mass
Y	0	0	-2/3	+1	$\gtrsim 1\mathrm{TeV}$
$\ell_d$	1/2	1	0	+1	$\mathcal{O}(10-140\mathrm{MeV})$
$\psi_B$	1/2	0	-1	+1	$\gtrsim 1.2\mathrm{GeV}$

Collider bounds (as just discussed)

Stability of matter, neutron star bounds

**Allowed Interactions:** 

$$\mathcal{L} \supset y_{u_i d_j} Y^* \bar{u}_i d_j^c + y_{\psi d_k} Y \bar{\psi}_B d_k^c + h.c.$$

$$\longrightarrow$$

$$\mathcal{L}_{\mathrm{eff}} = rac{y^2}{M_Y^2} ar{u}_i^c d_j d_k^c \psi_B \qquad egin{array}{l} \mathit{dark \ baryon \ 'mixing''} \ \mathit{baryon \ 'mixing''} \end{array}$$

## Example Charge Assignment

 $m_{\chi_2} + m_{\xi} > m_{\psi_B} > m_{\mathcal{B}}$ L = -1, B = 0L = 0, B = -1B and L conserving our portal  $\psi_B$ operator L = 1, B = 0L = 0, B = 1

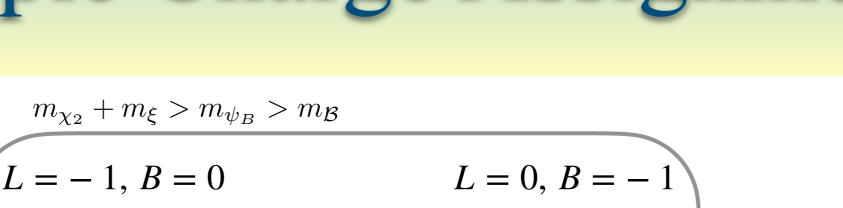
Dark Leptons

Dark Baryons

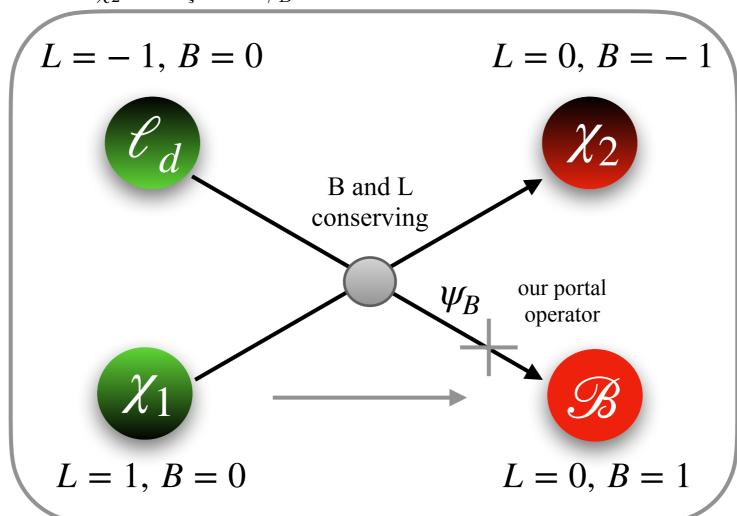
$$\mathcal{L} \supset y_b \, \bar{\psi}_B \xi \chi_2 + y_l \, \bar{\ell}_d \xi \chi_1 + \text{h.c.}.$$

MeV scale Dirac Fermion mediator

## Example Charge Assignment



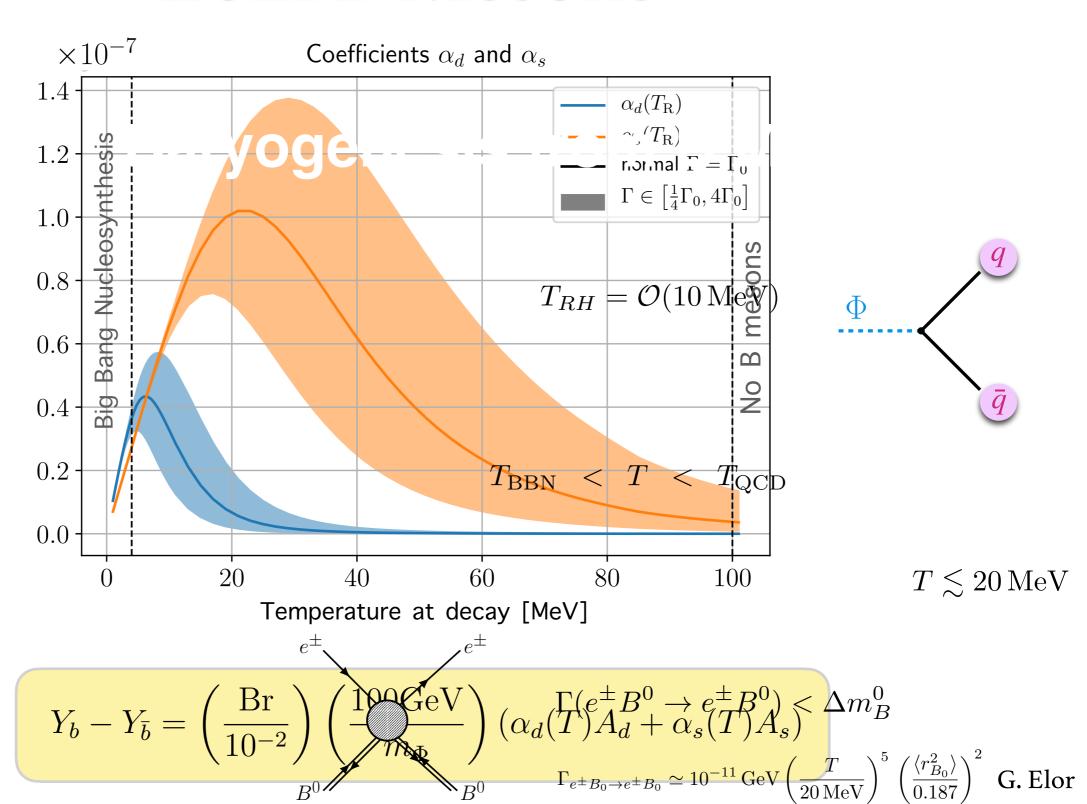
Dark Leptons



Dark Baryons

$$\langle \sigma v \rangle \simeq 10^{-15} \,\mathrm{GeV}^{-2} \,(y_l \, y_b)^2 \times \left(\frac{10 \,\mathrm{MeV}}{m_{\ell_d}}\right) \left(\frac{20 \,\mathrm{GeV}}{m_{\chi_1}}\right) \left(\frac{10 \,\mathrm{GeV}}{m_{\chi_2}}\right)$$

# Baryogenesis and Dark Matter from B Mesons



## A Supersymmetric Theory

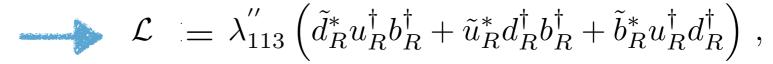
#### MSSM, R Symmetry, and Dirac Gauginos and Sterile Neutrios

Superfield	R-Charge	L no.
$\mathbf{U}^c,\mathbf{D}^c$	2/3	0
Q	4/3	0
$\mathbf{H}_u, \mathbf{H}_d$	0	0
$\mathbf{R}_u,\mathbf{R}_d$	2	0
S	0	0
${f L}$	1	1
$\mathbf{E}^c$	1	-1
$\mathbf{N}_R^c$	1	-1

"RPV" 
$$\mathbf{W} = y_u \mathbf{Q} \mathbf{H}_u \mathbf{U}^c - y_d \mathbf{Q} \mathbf{H}_d \mathbf{D}^c - y_e \mathbf{L} \mathbf{H}_d \mathbf{E}^c + \frac{1}{2} \lambda_{ijk}'' \mathbf{U}_i^c \mathbf{D}_j^c \mathbf{D}_k^c$$

$$+ \mu_u \mathbf{H}_u \mathbf{R}_d + \mu_d \mathbf{R}_u \mathbf{H}_d$$

$$+ \lambda_u^t \mathbf{H}_u \mathbf{T} \mathbf{R}_d + \lambda_d^t \mathbf{R}_u \mathbf{T} \mathbf{H}_d + \lambda_d^s \mathbf{S} \mathbf{R}_u \mathbf{H}_d.$$



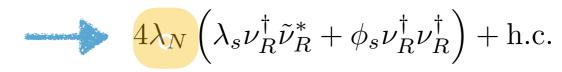
#### Gauge:

$$\mathcal{L}_{\text{gauge}} = -\sqrt{2}g(\phi T^a \psi^{\dagger}) \lambda^{a\dagger} + \text{h.c.}$$

$$\Rightarrow -\sqrt{2}g(\tilde{d}_R^* d_R \tilde{B}^{\dagger}) - \sqrt{2}g(\tilde{d}_L d_L^{\dagger} \tilde{B}^{\dagger}) + \text{h.c.}$$

#### Neutrio:

$$\mathbf{W} = \frac{\lambda_N}{4} \mathbf{S} \mathbf{N}_R^c \mathbf{N}_R^c + \mathbf{H}_u \mathbf{L}^i y_N^{ij} \mathbf{N}_R^{c,j} + \frac{1}{2} \mathbf{N}_R^c M_M \mathbf{N}_R^c + \text{h.c.},$$



Parameter space: "RPV" couplings and squark mass mixing

## A Supersymmetric Theory

Superpartners and SM particles have different charge under an unbroken R-symmetry. We can identify this with Baryon number.

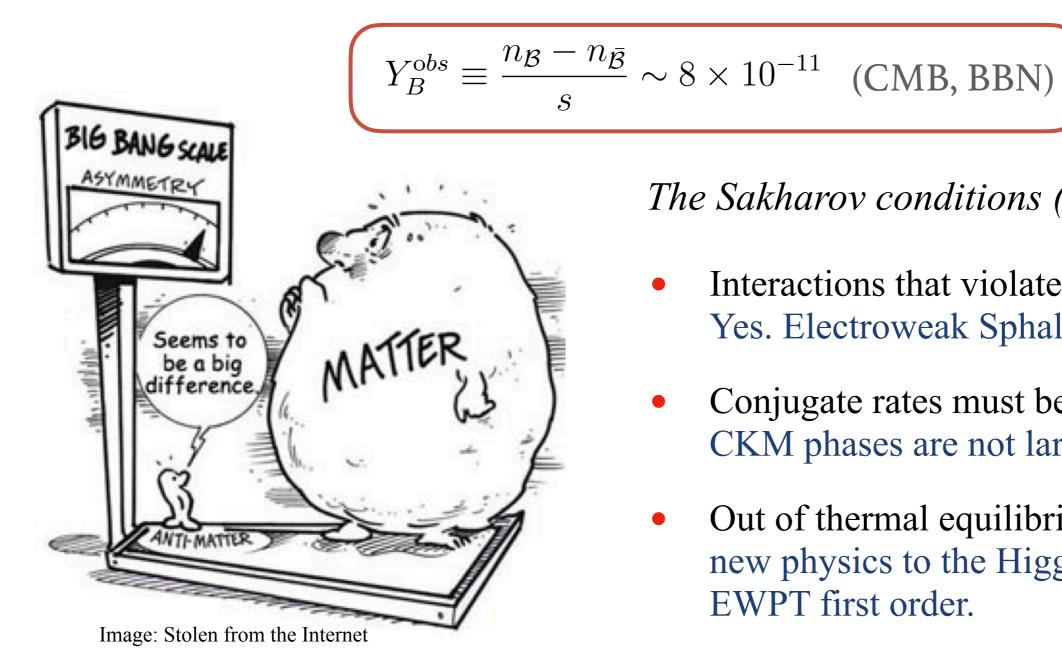


Superpartners as dark baryons.

	Field	Spin	$Q_{EM}$	Baryon no.	$\mathbb{Z}_2$	Mass
	Φ	0	0	0	+1	11 - 100  GeV
MSSM Squark	$ ilde{d}_R$	0	-1/3	-2/3	+1	$\mathcal{O}(\mathrm{TeV})$
Dirac Bino	$\left[egin{array}{c}  ilde{B} \ \lambda_s^\dagger \end{array} ight]$	1/2	0	-1	+1	$\mathcal{O}(\mathrm{GeV})$
Right handed	$ u_R$	1/2	0	0	-1	$\mathcal{O}(\mathrm{GeV})$
neutrino multiplet	$ ilde{ u}_R$	0	0	-1	-1	$\mathcal{O}(\mathrm{GeV})$

## Baryogenesis from the SM?

#### How to generate a matter/antimatter asymmetry



The Sakharov conditions (1967):

- Interactions that violate Baryon number. Yes. Electroweak Sphalerons.
- Conjugate rates must be different. CPV CKM phases are not large enough.
- Out of thermal equilibrium. Need to add new physics to the Higgs sector to make EWPT first order.

# A Need for BSM Physics

How to generate a matter/antimatter asymmetry

#### **B** violation

Sphalerons
Explicit B violation
Explicit L violation
Other particle number violation

Sakharov Conditions

Cosmological phase transitions Out-of-equilibrium decays

Out of equilibrium

#### **CP** violation

New CP violation in scalars New CP violation in quarks New CP violation in leptons CP violation in a dark sector